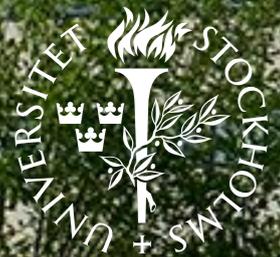


# DarkSUSY 6

## Tutorial – Part I

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# Question

- Have you managed to install DarkSUSY and run dstest?

Go to [menti.com](https://www.menti.com) and enter the code 98 63 15  
to answer

# Question

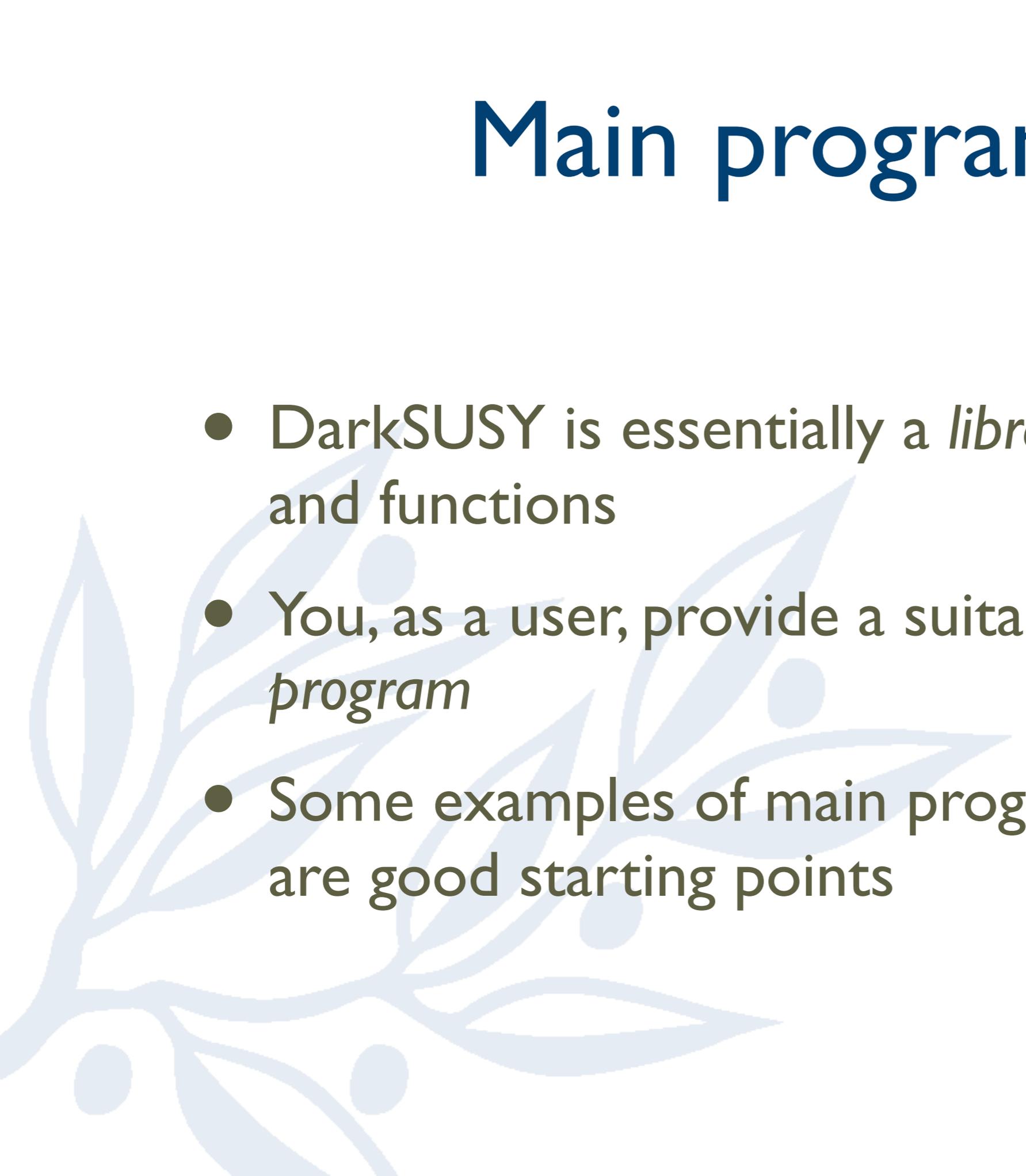
- How well do you know Fortran?

Go to [menti.com](https://www.menti.com) and enter the code 84 11 03  
to answer

# DarkSUSY 6.0.0

- DS 6.0.0 is a major re-organization and (partly) rewrite of the code.
  - It is tested on gfortran 5, 6 and 7 on Mac OS X and Ubuntu.
  - The manual is included in the distribution
- If you find problems/have questions, ask me during the school or e-mail me at [edsjo@fysik.su.se](mailto:edsjo@fysik.su.se)

# Main programs

- DarkSUSY is essentially a *library* of routines and functions
  - You, as a user, provide a suitable *main program*
  - Some examples of main programs exist and are good starting points
- 

# Fortran

- Fortran is rather basic, but gives fast code and should be fairly easy to follow/write
- DS6 consists of mainly Fortran 77 code, with some Fortran 95/03 additions

# Typical Fortran program

6 chars  
↔

```
program myprogram
implicit none
include 'dsver.h'
real*8 oh2,xf
integer unphys,war,ierr,iwar,nfc
real*8 dsrdomega

call dsinit
call dsgive_model(500.0d0,1000.d0,300.d0,10.d0,
& 3000.d0,0.d0,0.d0)
call dsmodelsetup(unphys,war)
oh2=dsrdomega(1,1,xf,ierr,iwar,nfc)
write(*,*) 'Relic density, omega h^2 = ',oh2

end
```

includes files, typically global variables (common blocks)

declarations

arguments (must match declaration)

subroutine call

function call

exponent for double precision (real\*8) number

program

continuation character (col 5)

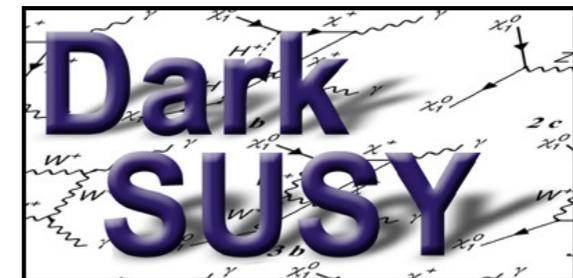
See example programs or tutorials on the web for more Fortran examples

# Manual

- In the DarkSUSY root folder, type `make pdf-manual` to produce the manual. It is available in `docs/Manual.pdf`

- There is also a long version (with routine headers) that you create with `make pdf-manual-long`

**Task:** create the manual and look at it!



DarkSUSY darksusy-6.0.0

Manual and description of routines

Documentation harvested from documentation and source directories with harvestdoc.pl

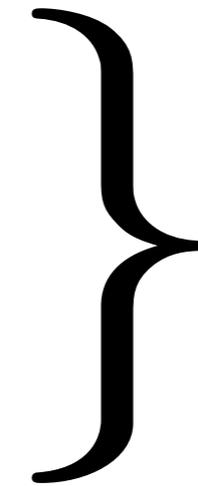
Wed Feb 28 02:57:36 2018

<http://www.darksusy.org>

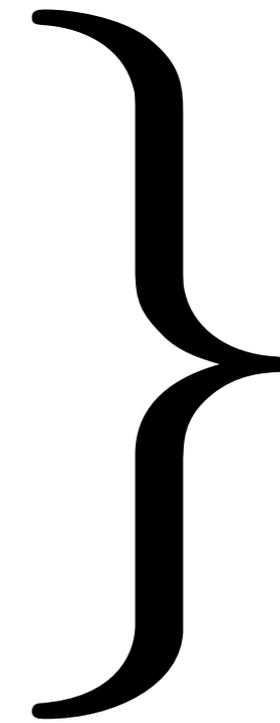
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# Outline of hands-on

1. dstest program
2. dsmain\_wimp program
  - MSSM
  - generic WIMP
3. Writing your own programs and using makefiles in DarkSUSY 6
4. Using dsmain\_wimp with SLHA files
5. Other example programs
6. Replaceable functions
7. Direct detection example
8. Creating a new particle physics module



Part I



Part II+III

# I. dstest program

- The dstest program is used to test your installation
- It calculates observables (masses, relic density, direct and indirect rates, ...) and compares with pre-computed values

```
cd examples/test  
./dstest
```

(already compiled with main make, takes about 60 seconds to run)

Output should end with

```
Total number of errors in dstest:      0
```

[Show code]

## 2. dsmain\_wimp.F

- In examples/ we have the file dsmain\_wimp.F which essentially does what dstest does, but in a more user-friendly way.
- **Task:** run it with  
`./dsmain_wimp`
- It will ask you which model you want to run:

```
What kind of SUSY model do you want to look at?  
1 = MSSM-7  
2 = cMSSM  
3 = as read from an SLHA2 file
```

# MSSM-7 example

- Pick 1: MSSM-7 and enter (e.g.)  
mu: 1000  
M2: 1000  
MA: 400  
tan( $\beta$ ): 10  
m0: 3000  
At/m0: 0  
Ab/m0: 0
- Then answer 0 to not write out an SLHA file (or something else if you want to)
- Observables are then calculated...

Task: run `dsmain_wimp` and inspect the output to see what is calculated

# Output

-----  
FeynHiggs 2.13.0

built on Feb 27, 2018

H. Bahl, T. Hahn, S. Heinemeyer, W. Hollik, S. Passehr, H. Rzehak, G. Weiglein

<http://feynhiggs.de>  
-----

FHHiggsCorr contains code by:

P. Slavich et al. (2L rMSSM Higgs self-energies)

H. Rzehak et al. (2L cMSSM asat Higgs self-energies)

S. Passehr et al. (2L cMSSM atat Higgs self-energies)

FHHiggsProd contains code by

- SM XS for VBF, WH, ZH, ttH are taken

from the LHC Higgs Cross Section Working Group,

<https://twiki.cern.ch/twiki/bin/view/LHCPhysics/CrossSections>

see there for an extensive list of references

- SM bbH XS: Harlander et al. hep-ph/0304035

- SM ggH XS: Grazzini et al.

<http://theory.fi.infn.it/grazzini/hcalculators.html>

see there for an extensive list of references

- 2HDM charged Higgs XS: Plehn et al.

- heavy charged Higgs XS:

Dittmaier et al. arXiv:0906.2648

Flechl et al. arXiv:1307.1347

g-2 amplitude [ $a_{\mu} = (g-2)/2 =$  ]: 3.8183990533120070E-011

Do you want to write out an SLHA2 file for your model?

0 = no

1 = yes, with full 6x6 sfermion mixing

2 = yes, with minimal flavour violation

```

***** MODEL: INTMOD000001          *****
istatus =          1  (=0 OK, <0 error, >0 warning, =1 experimentally excluded)
info =          4512
WIMP mass =          499.30971144910990

Calculating relic density without coannihilations, please be patient...
Oh2 =          0.96579250432421615          0          0
Calculating omega h^2 with coannihilations, please be patient...
with coannihilations Oh2 =          0.96579250432421615          0          0
Chemical decoupling (freeze-out) occurred at
T_f =          22.879465168299642          GeV.

Kinetic decoupling temperature, Tkd =          216.93575300730424          MeV
The resulting cutoff in the power spectrum corresponds to a mass of M_cut/M_sun =
2.2909066986614345E-009

Calculating DM-nucleon scattering cross sections...
sigsip (pb) =          8.8992537784068590E-010
sigsin (pb) =          9.2400044165495133E-010
sigsdp (pb) =          1.9718137439575352E-007
sigsdn (pb) =          1.4088379669121179E-007
proton: sigsi (pb) =          8.8992537784068590E-010  sigsd (pb) =
1.9718137439575352E-007
A=          23  Z=          11  sigsi (pb) =          2.3037796208861801E-004  sigsd (pb) =
5.1548841697066729E-006
sigsi*ff(pb)=          2.2292275772657175E-004  sigsd*ff(pb)=          4.9743110005459323E-006
A=          127  Z=          53  sigsi (pb) =          0.15253489069800388  sigsd (pb) =
1.7886424072791722E-005
sigsi*ff(pb)=          9.5441464487036914E-002  sigsd*ff(pb)=          1.1823115084786536E-005
A=          73  Z=          32  sigsi (pb) =          1.9706985553529958E-002  sigsd (pb) =
4.9292414333092543E-005
sigsi*ff(pb)=          1.6068638785545389E-002  sigsd*ff(pb)=          3.8390286999876545E-005
A=          1  Z=          1  sigsi (pb) =          8.8992537784068590E-010  sigsd (pb) =
1.9718137439575352E-007
sigsi*ff(pb)=          0.00000000000000000000  sigsd*ff(pb)=          0.00000000000000000000

```

Calculating gamma ray fluxes...

fluxgacdifff = 1.8515008372241465E-013 ph/(cm<sup>2</sup> s GeV)  
fluxgac = 3.8000236750318762E-011 ph/(cm<sup>2</sup> s)

Total number of photon lines in module MSSM:

2

photon flux from line No. 1 = 1.6539906105554819E-018 ph/(cm<sup>2</sup> s)  
[at E = 499.30971144910990 +/- 0.000000000000000000 GeV]  
photon flux from line No. 2 = 2.6166109285505223E-019 ph/(cm<sup>2</sup> s)  
[at E = 495.14637444449431 +/- 2.49000000000000002 GeV]

nsigvgacont = 8.3049860752652031E-026  
nsigvgacdifff = 3.6350954555843692E-026 GeV<sup>-1</sup>  
nsigvline 1 = 1.8074057116456577E-033  
nsigvline 2 = 2.8593134128061127E-034

Calculating antiproton fluxes...

solar modulated pbar flux at 0.35 GeV [GeV<sup>-1</sup> cm<sup>-2</sup> s<sup>-1</sup> sr<sup>-1</sup>):  
6.4724132322023058E-010  
solar modulated pbar flux at 1.76 GeV [GeV<sup>-1</sup> cm<sup>-2</sup> s<sup>-1</sup> sr<sup>-1</sup>):  
1.0017260278195632E-009  
solar modulated pbar flux at 3.00 GeV [GeV<sup>-1</sup> cm<sup>-2</sup> s<sup>-1</sup> sr<sup>-1</sup>):  
8.9088984206521731E-010

Calculating antideuteron fluxes...

solar modulated dbar flux at 1.00 GeV [GeV<sup>-1</sup> cm<sup>-2</sup> s<sup>-1</sup> sr<sup>-1</sup>):  
2.3931160388253188E-013

Calculating positron fluxes at 1 GeV...

phiep= 7.8088318707756649E-008 GeV<sup>-1</sup> cm<sup>-2</sup> s<sup>-1</sup> sr<sup>-1</sup>

Calculating rates in neutrino telescopes

Flux from the Earth = 1.9046606405318337E-008 km<sup>-2</sup> yr<sup>-1</sup>  
Flux from the Sun = 0.57508123135314870 km<sup>-2</sup> yr<sup>-1</sup>

Calculating neutrino-induced muon fluxes from the halo...

Muon flux from halo = 2.2035442898808658E-003 km<sup>-2</sup> yr<sup>-1</sup>

# Which module?

- At the end of the `dsmain_wimp` run we got

```
-----  
The DarkSUSY example program has finished successfully.  
Particle module that was used: MSSM  
-----
```

```
[simply call 'make -B dsmain_wimp DS_MODULE=<MY_MODULE>' if you want to try  
with a different module <MY_MODULE>]
```

- Try compiling again with

```
make -B dsmain_wimp DS_MODULE=generic_wimp  
./dsmain_wimp
```

- Enter e.g.  
mass: 100  
self-conjugate: 0  
ann cross section: 3e-26  
PDG: 5  
scattering cross section: 1e-42

# Output

```
Calculating omega h^2 without coannihilations, please be patient...  
without coannihilations Oh2 =      8.5782015186659649E-002      0      0  
Chemical decoupling (freeze-out) occurred at  
T_f =      4.4034841137539358      GeV.
```

etc

