Relativistic Hydrodynamics in High-Energy Heavy Ion Collisions



for the Origin of Particles and the Universe

Kobayashi Maskawa Institute Department of Physics, Nagoya University *Chiho NONAKA*

December 13, 2013@KMI 2013, Nagoya

Relativistic Heavy Ion Collisions

RHIC:2000





Observables: a	a lot of experimental data at RHIC and LHC	photons/leptons
		bulk property
_		Jets
		heavy quarkonia

Phenomenological model

sQGP Initial condition → Hydrodynamic model → Freezeout process Experimental data NONAKA

Higher Harmonics



 $\frac{dN}{dyd\phi} \propto 1 + 2v_1 \cos(\phi - \Theta_1) + 2v_2 \cos 2(\phi - \Theta_2) + 2v_3 \cos 3(\phi - \Theta_3) + 2v_4 \cos 4(\phi - \Theta_4) + \cdots$





Initial Conditions





Ollitrault

Initial Conditions





Ollitrault

Initial Conditions





Numerical Scheme



Numerical algorithm for hydrodynamic evolution

- ✓ shock-wave capturing scheme✓ stable
- less numerical viscosity



Hydrodynamic Expansion



Bulk property transport coefficients

importance of numerical algorithm !



Akamatsu, Inutsuka, CN, Takamoto: arXiv:1302.1665, J. Comp. Phys. (2014) 34

HYDRODYNAMIC MODEL



Viscous Hydrodynamic Model

- Relativistic viscous hydrodynamic equation $\partial_{\mu}T^{\mu\nu} = 0$
 - First order in gradient: acausality
 - Second order in gradient:
 - Israel-Stewart, Ottinger and Grmela, AdS/CFT,

Grad's 14-momentum expansion, Renomarization group

- Numerical scheme
 - Shock-wave capturing schemes: Riemann problem
 - Godunov scheme: analytical solution of Riemann problem
 - SHASTA: the first version of Flux Corrected Transport algorithm, Song, Heinz, Pang, Victor...
 - Kurganov-Tadmor (KT) scheme, McGill





Takamoto and Inutsuka, arXiv:1106.1732 Akamatsu, Inutsuka, CN, Takamoto, arXiv:1302.1665

• Israel-Stewart Theory

(ideal hydro) **1. dissipative fluid dynamics** = advection + dissipation



Riemann solver: Godunov method

Two shock approximation

Mignone, Plewa and Bodo, Astrophys. J. S160, 199 (2005)

Rarefaction wave \longrightarrow shock wave

2. relaxation equation = advection + stiff equation



Numerical Scheme

• Israel-Stewart Theory

Takamoto and Inutsuka, arXiv:1106.1732

1. Dissipative fluid equation

$$\partial_{\mu}T^{\mu\nu} = 0$$

$$T^{\mu\nu} = (\epsilon + p)u^{\mu}u^{\nu} - pg^{\mu\nu} + q^{\mu}u^{\nu} + q^{\nu}u^{\mu} + \tau^{\mu\nu}$$

$$= T_{\text{ideal}} + T_{\text{dissip}}$$

$$\partial_{t}U + \nabla \cdot F(U) = 0 \qquad U = U_{\text{ideal}} + U_{\text{dissip}}$$





• Shock Tube Test : Molnar, Niemi, Rischke, Eur. Phys. J.C65, 615 (2010)





Shocktube problem

• Ideal case





L1 Norm

• Numerical dissipation: deviation from analytical solution





Large ΔT difference

10







• SHASTA with small A_{ad} has large numerical dissipation



Artificial and Physical Viscosities



Molnar, Niemi, Rischke, Eur. Phys. J. C65, 615 (2010)

2

3

stability



Large ΔT difference



- Our algorithm is stable even with small numerical dissipation.







Our Dynamical Model



Our Dynamical Model



Initial Pressure Distribution

• MC-KLN (centrality 15-20%)





Time Evolution of v_n



Hydro + UrQMD

Transverse momentum spectrum



Pt distribution at LHC has flatter slope

Larger radial flow at LHC



Effect of Hadronic Interaction



- Effect of final state interactions is small
- Slope of proton Pt spectra become flatter



Higher harmonics from Hydro + UrQMD

Effect of hadronic interaction







Importance of numerical scheme in Hydrodynamic Models

- We develop a state-of-the-art numerical scheme
 - Shock wave capturing scheme: Godunov method

Our algorithm

- Less artificial diffusion: crucial for viscosity analyses
- Stable for strong shock wave
- Construction of a hybrid model
 - Fluctuating initial conditions + Hydrodynamic evolution +

UrQMD

- Higher Harmonics
 - Time evolution, hadron interaction

