Isotropization in Heavy Ion Collisions at High Energy

Kobayashi-Maskawa Institute, Nagoya, March 2014

T. Epelbaum, FG: arXiv:1307.1765 arXiv:1307.2214

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Outline

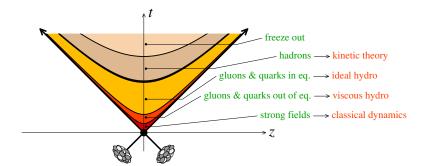


- 1 The surprising success of relativistic hydrodynamics
- 2 Modern ideas for the thermalization of quantum systems
- 3 CGC description of heavy ion collisions
- 4 Isotropization in Heavy Ion Collisions

hydrodynamics

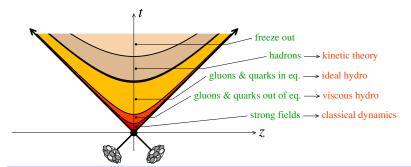
Relativistic

Stages of a nucleus-nucleus collision

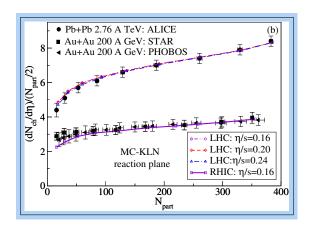


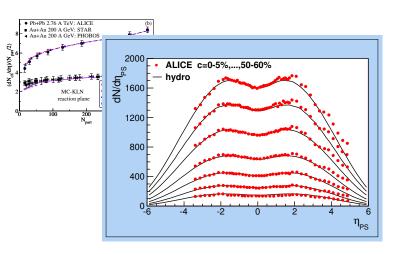
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Stages of a nucleus-nucleus collision

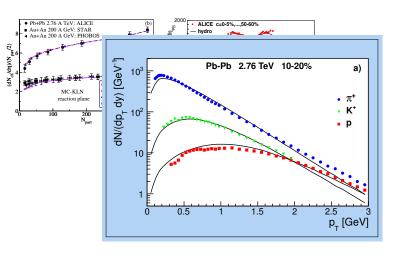


Well described as a fluid expanding into vacuum according to relativistic hydrodynamics

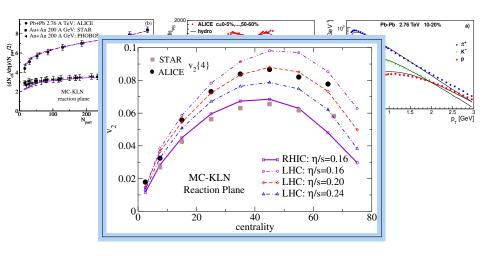


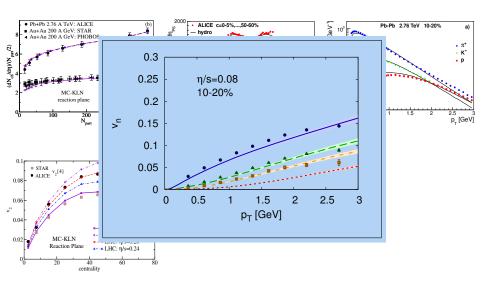


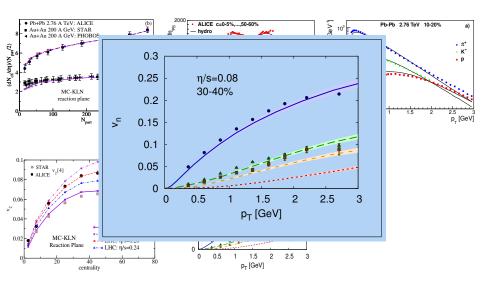
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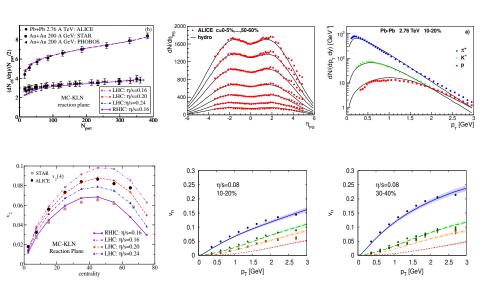


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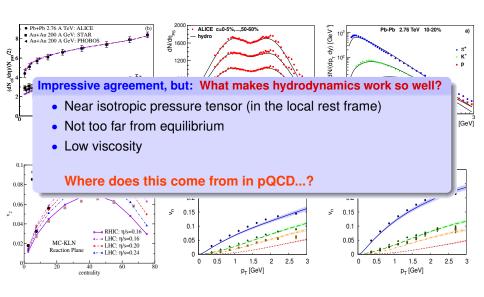








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 Hydrodynamics is a macroscopic description based on energy-momentum conservation :

$$\vartheta_\mu T^{\mu\nu}=0$$

True in any quantum field theory



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- Additional assumption : at macroscopic scales, $\mathsf{T}^{\mu\nu}$ is expressible in terms of ε (energy density), P (pressure) and \mathfrak{u}^μ (fluid velocity field)
- For a frictionless fluid : $T^{\mu\nu}_{ideal} = (\varepsilon + P) \, u^{\mu} u^{\nu} P \, g^{\mu\nu}$



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- In general : $T^{\mu\nu} = T^{\mu\nu}_{ideal} \oplus \eta \nabla^{\mu} v^{\nu} \oplus \zeta g^{\mu\nu} (\nabla_{\rho} v^{\rho}) \oplus \cdots$



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- In general : $T^{\mu\nu} = T^{\mu\nu}_{ideal} \oplus \eta \nabla^{\mu} v^{\nu} \oplus \zeta g^{\mu\nu} (\nabla_{\rho} v^{\rho}) \oplus \cdots$
- Microscopic inputs : $\epsilon = f(P)$ (EoS), η, ζ, \cdots (transport coeff.)

Why is it hard to justify in QCD?



Just after the collision, $T^{\mu\nu}$ is far from ideal

$$\mathsf{T}^{\mu\nu}_{\begin{subarray}{c} \mathbf{QCD} \\ \text{rest frame} \end{subarray}} = \begin{pmatrix} \varepsilon & & & \\ & \varepsilon & & \\ & & \varepsilon & \\ & & & -\varepsilon \end{pmatrix} \qquad \mathsf{T}^{\mu\nu}_{\begin{subarray}{c} \mathbf{ideal} \\ \text{rest frame} \end{subarray}} = \begin{pmatrix} \varepsilon & & & \\ & \frac{\varepsilon}{3} & & \\ & & \frac{\varepsilon}{3} & \\ & & & \frac{\varepsilon}{3} \end{pmatrix}$$

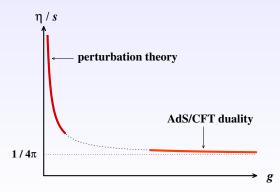
⇒ Very large viscous corrections

Why is it hard to justify in QCD?



Shear viscosity at weak coupling in QCD

$$\frac{\eta}{s} = \frac{5.12}{g^4 \ln\left(\frac{2.42}{g}\right)}$$



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OM Quantum chace

- QM, Quantum chaos
 - Berry's conjecture

Formulation of QM in the classical phase-space



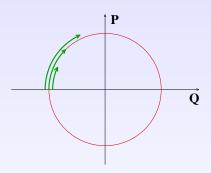
- Quantum Mechanics introduces a natural smearing due to the uncertainty principle. To make the connection with classical mechanics, it is useful to use Moyal's formulation of QM in terms of classical variables
- Dual formulation of QM :

Density	ρ̂		W(Q, P)
Evolution	$\partial_t \widehat{\rho} + i[\widehat{H}, \widehat{\rho}] = 0$	Weyl-Wigner trans.	$\partial_t \frac{\mathbf{W}}{\mathbf{W}} + \{\{\mathbf{W}, \mathbf{H}\}\} = 0$
Initial condition	coherent state		Gaussian in Q, P

• Moyal bracket :
$$\{\{\cdot,\cdot\}\}=\underbrace{\{\cdot,\cdot\}}_{\text{Poisson bracket}}+\mathcal{O}(\hbar^2)$$

The Moyal equation becomes the Liouville equation in the classical limit $\hbar \to 0$

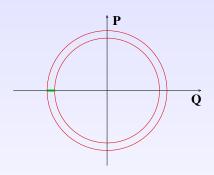




• The oscillation frequency depends on the initial condition

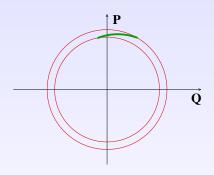
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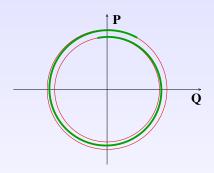
- The oscillation frequency depends on the initial condition
- Because of QM, the initial ensemble is a set of width ~ ħ





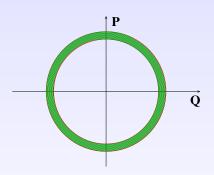
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- The oscillation frequency depends on the initial condition
- Because of QM, the initial ensemble is a set of width $\sim \hbar$
- This ensemble of initial configurations spreads in time
- At large times, the ensemble fills densely all the region allowed by energy conservation ⇒ microcanonical equilibrium

Quantum chaos



- Central issue: consider a Hamiltonian that leads to chaotic classical behavior; What happens when this system is quantized?
- Schrodinger's equation is linear :

$$i\vartheta_t\Psi=\widehat{H}\,\Psi$$

• Once we know the spectrum of the Hamiltonian $\{E_n, \Psi_n\}$, any wavefunction evolves as:

$$\Psi(t) = \sum_n c_n \, e^{i E_n \, t} \, \Psi_n$$

 $E_n \in \mathbb{R} \Rightarrow$ nothing is unstable. Where is the chaos?

Berry's conjecture (1977)



- The complexity of the classical dynamics translates in the complexity of the high lying eigenfunctions
- Berry's conjecture: for most practical purposes, high lying eigenfunctions of classically chaotic systems behave as Gaussian random functions with 2-point correlations given by

$$\left\langle \Psi^*(X - \frac{s}{2})\Psi(X + \frac{s}{2}) \right\rangle = \int dP \; e^{iP \cdot s/\hbar} \; \delta \big[E - H(X, P) \big]$$

 Then, the Wigner distribution associated with the eigenfunction $\Psi_{\scriptscriptstyle F}$ is

$$W(\mathbf{X}, \mathbf{P}) = \int d\mathbf{s} \ e^{-i\mathbf{P}\cdot\mathbf{s}/\hbar} \ \Psi_{\mathrm{E}}^*(\mathbf{X} - \frac{\mathbf{s}}{2}) \Psi_{\mathrm{E}}(\mathbf{X} + \frac{\mathbf{s}}{2})$$
$$\sim \delta[\mathbf{E} - \mathbf{H}(\mathbf{X}, \mathbf{P})]$$

⇒ micro-canonical equilibrium for a single eigenstate

Eigenstate thermalization hypothesis (Srednicki, 1994)



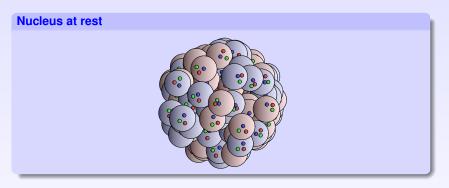
- If an energy eigenstate obeys Berry's conjecture, then a measurement performed on that state will lead to the Bose-Einstein (or Fermi-Dirac) distribution for the single particle distribution
- Generic states approach equilibrium via decoherence of their individual energy eigenstate components

CGC Description

of Heavy Ion Collisions

What do we need to know about nuclei?





· At low energy: valence quarks

What do we need to know about nuclei?



Slightly boosted nucleus



- · At low energy: valence quarks
- · At higher energy:
 - Lorenz contraction of longitudinal sizes
 - Time dilation > slowing down of the internal dynamics
 - · Gluons start becoming important

What do we need to know about nuclei?



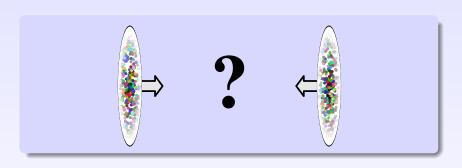
High energy nucleus



- · At low energy: valence quarks
- At higher energy :
 - Lorenz contraction of longitudinal sizes
 - Time dilation > slowing down of the internal dynamics
 - Gluons start becoming important
- · At very high energy : gluons dominate

Multiple scatterings and gluon recombination

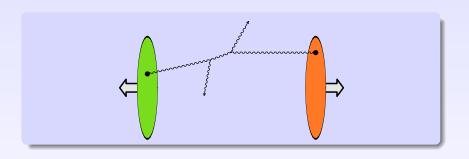




 Main difficulty: How to treat collisions involving a large number of partons?

Multiple scatterings and gluon recombination

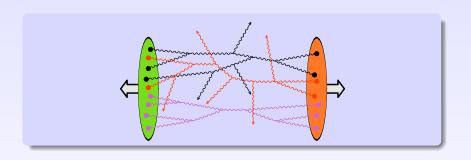




Dilute regime : one parton in each projectile interact
 single parton distributions, standard perturbation theory

Multiple scatterings and gluon recombination





- Dense regime : multiparton processes become crucial

 - > multi-parton distributions
 - $\,\rhd\,$ alternative approach : treat the gluons in the projectiles as external currents

$$\mathcal{L} = -\frac{1}{4}\mathsf{F}^2 + \mathsf{A}\cdot(\mathsf{J}_1+\mathsf{J}_2)$$

(gluons only, field A for $k^+ < \Lambda$, classical source J for $k^+ > \Lambda$)

Color Glass Condensate



CGC = effective theory of small x gluons

• The fast partons $(k^+ > \Lambda^+)$ are frozen by time dilation > described as static color sources on the light-cone :

$$J^{\mu} = \delta^{\mu +} \rho(x^{-}, \vec{x}_{\perp})$$
 $(0 < x^{-} < 1/\Lambda^{+})$

- The color sources ρ are random, and described by a probability distribution $W_{\Lambda^+}[\rho]$
- Slow partons ($k^+ < \Lambda^+$) cannot be considered static over the time-scales of the collision process
 - > must be treated as standard gauge fields
 - \triangleright eikonal coupling to the current $J^{\mu}: A_{\mu}J^{\mu}$

Terminology

Weakly coupled: g ≪ 1

• Weakly interacting : $gA \ll 1$ $g^2f(p) \ll 1$ $(2 \rightarrow 2) \gg (2 \rightarrow 3), (3 \rightarrow 2), \cdots$

• Strongly interacting : $g\mathcal{A}\sim 1$ $g^2f(p)\sim 1$ $(2\to 2)\sim (2\to 3)\sim (3\to 2)\sim \cdots$

No well defined quasi-particles

CGC = weakly coupled, but strongly interacting effective theory

Power counting



• CGC effective theory with cutoff at the scale Λ_0 :



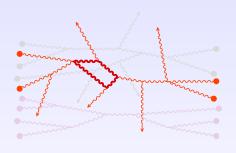
$$\mathcal{S} = \underbrace{-\frac{1}{4} \int F_{\mu\nu} F^{\mu\nu}}_{\mathcal{S}_{YM}} + \int \underbrace{(J_1^{\mu} + J_2^{\mu})}_{\text{fast partons}} A_{\mu}$$

• Expansion in g² in the saturated regime:

$$T^{\mu\nu} \sim \frac{1}{g^2} \left[c_0 + c_1 \; g^2 + c_2 \; g^4 + \cdots \right]$$

Power counting





In the saturated regime: $J^{\mu} \sim g^{-1}$

$$g^{-2}$$
 $g^{\text{\#}}$ of external legs $g^{2\times(\text{\# of loops})}$

No dependence on the number of sources J^µ
 ▷ infinite number of graphs at each order

Leading Order in g^2 : tree diagrams



 The Leading Order is the sum of all the tree diagrams
 Observables can be expressed in terms of classical solutions of Yang-Mills equations :

$$\mathfrak{D}_{\mu}\mathfrak{F}^{\mu\nu}=J_{1}^{\nu}+J_{2}^{\nu}$$

Boundary conditions for inclusive observables :

$$\lim_{x^0\to -\infty}\mathcal{A}^{\mu}(x)=0$$

Example: 00 component of the energy-momentum tensor

$$T_{LO}^{00} = \frac{1}{2} \left[\underbrace{\mathcal{E}^2 + \mathcal{B}^2}_{\text{class. fields}} \right]$$

Next to Leading Order in q^2 : 1-loop diagrams



Getting the NLO from tree graphs...

$$\mathcal{O}_{\text{NLO}} = \left[\frac{1}{2} \int_{\mathbf{u}, \mathbf{v}} \mathbf{\Gamma}_{2}(\mathbf{u}, \mathbf{v}) \, \mathbb{T}_{\mathbf{u}} \mathbb{T}_{\mathbf{v}} + \int_{\mathbf{u}} \alpha(\mathbf{u}) \, \mathbb{T}_{\mathbf{u}} \right] \, \mathcal{O}_{\text{LO}}$$

ullet I is the generator of the shifts of the initial value of the field :

$$\mathbb{T}_{\mathbf{u}} \sim \frac{\partial}{\partial \mathcal{A}_{init}}$$

$$exp\left[\int_{\mathbf{u}} \frac{\alpha_{\mathbf{u}}}{\alpha_{\mathbf{u}}} \, \mathbb{T}_{\mathbf{u}}\right] \, \underbrace{0 \left[\underbrace{\mathcal{A}_{\tau}(\underbrace{\mathcal{A}_{\text{init}}})}_{\text{init. value}} \right]}_{\text{init. value}} = \underbrace{0 \left[\mathcal{A}_{\tau}(\underbrace{\mathcal{A}_{\text{init}} + \boldsymbol{\alpha}}_{\text{shifted init. value}}) \right]}_{\text{shifted init. value}}$$

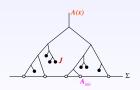
Shift operator \mathbb{T} – Definition



Equations of motion for a field ${\cal A}$ and a small perturbation α

$$\Box \mathcal{A} + V'(\mathcal{A}) = J$$

$$[\Box + V''(\mathcal{A})] \alpha = 0$$



• Getting the perturbation by shifting the initial condition of \mathcal{A} at one point :

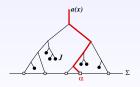
$$\alpha(\mathbf{x}) = \int_{\mathbf{u}} \alpha_{\mathbf{u}} \, \mathbb{T}_{\mathbf{u}} \, \mathcal{A}(\mathbf{x})$$

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Equations of motion for a field ${\cal A}$ and a small perturbation α

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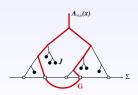
$$\alpha(\mathbf{x}) = \int_{\mathbf{u}} \alpha_{\mathbf{u}} \, \mathbb{T}_{\mathbf{u}} \, \mathcal{A}(\mathbf{x})$$

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Equations of motion for a field A and a small perturbation α

$$\begin{array}{rcl} \square \mathcal{A} + V'(\mathcal{A}) & = & J \\ [\square + V''(\mathcal{A})] & \pmb{\alpha} & = & 0 \end{array}$$



Getting the perturbation by shifting the initial condition of A at one point :

$$\alpha(\mathbf{x}) = \int_{\mathbf{u}} \alpha_{\mathbf{u}} \, \mathbb{T}_{\mathbf{u}} \, \mathcal{A}(\mathbf{x})$$

A loop is obtained by shifting the initial condition of A at two points

Initial state logarithms



• In the CGC, upper cutoff on the loop momentum: $k^{\pm} < \Lambda$, to avoid double counting with the sources $J_{1,2}^{\gamma}$ ⊳ logarithms of the cutoff

Central result for factorization at Leading Log

$$\begin{split} &\frac{1}{2}\int\limits_{u,v}^{} &\Gamma_{2}(u,v) \, \mathbb{T}_{u} \mathbb{T}_{v} + \int\limits_{u}^{} \alpha(u) \, \mathbb{T}_{u} = \\ &= \log \left(\Lambda^{+}\right) \, \mathfrak{H}_{1} + \log \left(\Lambda^{-}\right) \, \mathfrak{H}_{2} + \text{terms w/o logs} \end{split}$$

 $\mathfrak{H}_{1,2} = \text{JIMWLK Hamiltonians of the two nuclei}$

- No mixing between the logs of the two nuclei
- observables, these logs have a universal structure

Factorization of the logarithms



Inclusive observables at Leading Log accuracy

$$\left\langle 0\right\rangle_{\text{Leading Log}} = \int \left[D\rho_1 \ D\rho_2\right] \ W_1\left[\rho_1\right] \ W_2\left[\rho_2\right] \ \underbrace{0_{\text{LO}}[\rho_1,\rho_2]}_{\text{fixed }\rho_{1,2}}$$

Logs absorbed into the scale evolution of W_{1,2}

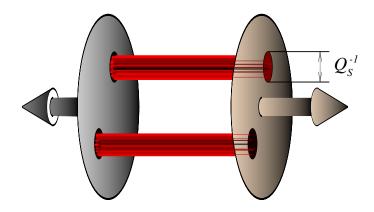
$$\Lambda \frac{\partial W}{\partial \Lambda} = \mathcal{H} W$$
 (JIMWLK equation)

• Universality: the same W's for all inclusive observables

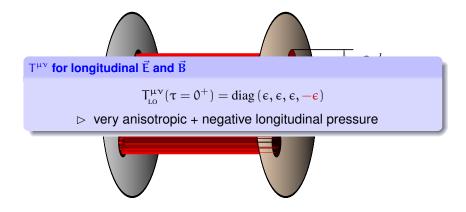
Heavy Ion Collisions

Isotropization in

Energy momentum tensor of the initial classical field

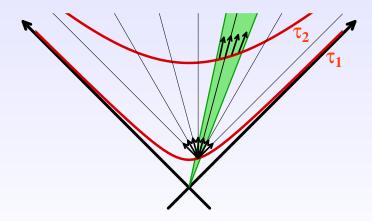


Energy momentum tensor of the initial classical field



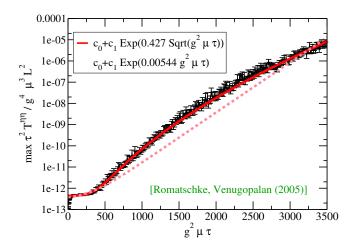
Competition between Expansion and Isotropization





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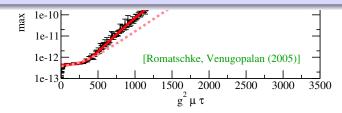
Weibel instabilities for small perturbations



Weibel instabilities for small perturbations

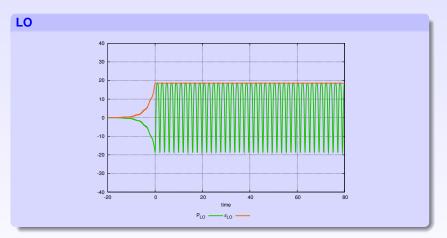


- The perturbations that alter the classical field in loop corrections diverge with time, like $\exp \sqrt{\mu \tau}$ $(\mu \sim Q_s)$
- Some components of $T^{\mu\nu}$ have secular divergences when evaluated beyond tree level



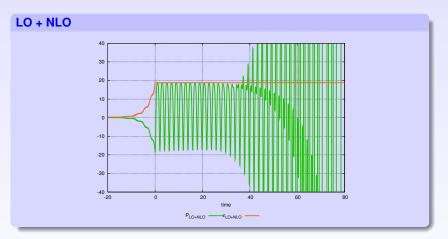
Example of pathologies in fixed order calculations (scalar theory)





Example of pathologies in fixed order calculations (scalar theory)



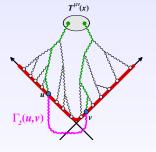


- Small correction to the energy density (protected by energy conservation)
- Secular divergence in the pressure

Improved power counting and resummation



$$\mathsf{Loop} \sim \mathsf{g}^2 \qquad , \qquad \mathbb{T} \sim \mathsf{e}^{\sqrt{\mu\tau}}$$



• 1 loop : $(ge^{\sqrt{\mu\tau}})^2$

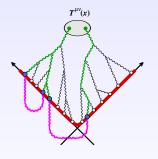
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Improved power counting and resummation



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$$\mathbb{T} \sim e^{\sqrt{\mu \tau}}$$



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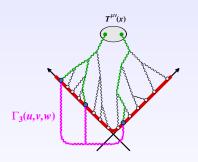
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• 2 disconnected loops : $(ge^{\sqrt{\mu\tau}})^4$

Improved power counting and resummation



Loop
$$\sim g^2$$
 , $\mathbb{T} \sim e^{\sqrt{\mu \tau}}$



 1 loop : $(ae^{\sqrt{\mu\tau}})^2$

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- 2 disconnected loops : $(ae^{\sqrt{\mu\tau}})^4$
- 2 entangled loops : $q(qe^{\sqrt{\mu\tau}})^3 > \text{subleading}$

Leading terms

- All disconnected loops to all orders
 - > exponentiation of the 1-loop result

Resummation of the leading secular terms



$$\begin{array}{lcl} \overline{T}^{\mu\nu}_{\text{resummed}} & = & \exp\left[\frac{1}{2}\int\limits_{\mathbf{u},\nu} \underline{\Gamma}_2(\mathbf{u},\nu) \mathbb{T}_{\mathbf{u}} \mathbb{T}_{\nu}\right] T^{\mu\nu}_{\text{LO}}[\mathcal{A}_{\text{init}}] \\ \\ & = & \underbrace{T^{\mu\nu}_{\text{LO}} + T^{\mu\nu}_{\text{NLO}}}_{\text{in full}} + \underbrace{T^{\mu\nu}_{\text{NNLO}} + \cdots}_{\text{partially}} \end{array}$$

 The exponentiation of the 1-loop result collects all the terms with the worst time behavior

Resummation of the leading secular terms



- The exponentiation of the 1-loop result collects all the terms with the worst time behavior
- Equivalent to Gaussian fluctuations of the initial field + classical time evolution
- At $Q_s \tau_0 \ll 1$: $\mathcal{A}_{init} \sim Q_s/g$, $a \sim Q_s$

$$e^{\frac{\alpha}{2}\partial_x^2} f(x) = \int_0^{+\infty} dz \frac{e^{-z^2/2\alpha}}{\sqrt{2\pi\alpha}} f(x+z)$$

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Note: Classical field + Fluctuations = Coherent state



 This Gaussian distribution of initial fields is the Wigner distribution of a **coherent state** $|A\rangle$

Coherent states are the "most classical quantum states"

Their Wigner distribution has the minimal support permitted by the uncertainty principle (O(ħ) for each mode)

• $|A\rangle$ is not an eigenstate of the full Hamiltonian > decoherence via interactions

What needs to be done?

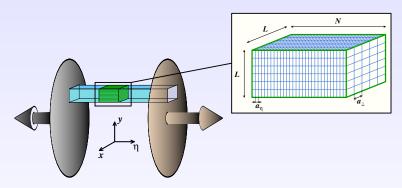


Main steps

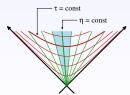
- 1. Determine the 2-point function $\Gamma_2(\mathbf{u}, \mathbf{v})$ that defines the Gaussian fluctuations, for the initial time $Q_s\tau_0$ of interest Note : this is an initial value problem, whose outcome is uniquely determined by the state of the system at $\mathbf{x}^0 = -\infty$, and depends on the history of the system from $\mathbf{x}^0 = -\infty$ to $\mathbf{\tau} = \mathbf{\tau}_0$ Problem solvable only if the fluctuations are weak, $\mathbf{a}^\mu \ll Q_s/g$ $Q_s\tau_0 \ll 1$ necessary for the fluctuations to be Gaussian
- 2. Solve the classical Yang-Mills equations from τ_0 to τ_f Note: the problem as a whole is boost invariant, but individual field configurations are not \implies 3+1 dimensions necessary
- 3. Do a Monte-Carlo sampling of the fluctuating initial conditions

Discretization of the expanding volume





- Comoving coordinates : τ, η, x_⊥
- Only a sub-volume is simulated + periodic boundary conditions
- L² × N lattice

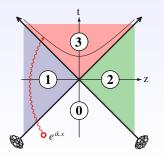


Gaussian spectrum of fluctuations



Expression of the variance (from 1-loop considerations)

$$\begin{array}{rcl} \Gamma_2(u,\nu) & = & \int\limits_{\text{modes }k} \alpha_k(u) \alpha_k^*(\nu) \\ \\ \left[\mathcal{D}_\rho \mathcal{D}^\rho \delta_\mu^\nu - \mathcal{D}_\mu \mathcal{D}^\nu + ig \, \mathfrak{F}_\mu^{\ \nu} \right] \! \alpha_k^\mu & = & 0 \quad , \quad \lim\limits_{x^0 \to -\infty} \alpha_k(x) \sim e^{ik \cdot x} \end{array}$$



- **0.** $\mathcal{A}^{\mu} = 0$. trivial
- **1,2**. A^{μ} = pure gauge, analytical solution
 - 3. A^{μ} non-perturbative \Rightarrow expansion in $Q_s \tau$
 - We need the fluctuations in Fock-Schwinger gauge $x^{+}a^{-} + x^{-}a^{+} = 0$

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 Delicate light-cone crossings, since $\mathcal{F}^{\mu\nu} = \infty$ there

Mode functions for given quantum numbers : v, k_{\perp}, λ, c

$$\begin{split} \alpha^i &= \beta^{+i} + \beta^{-i} & \qquad \qquad \alpha^\eta &= \mathcal{D}^i \Big(\frac{\beta^{+i}}{2+i\nu} - \frac{\beta^{-i}}{2-i\nu} \Big) \\ e^i &= -i\nu \Big(\beta^{+i} - \beta^{-i} \Big) & \qquad \qquad e^\eta &= -\mathcal{D}^i \Big(\beta^{+i} - \beta^{-i} \Big) \end{split}$$

$$\begin{split} \beta^{+i} &\equiv e^{\frac{\pi \nu}{2}} \Gamma(-i\nu) e^{i\nu\eta} \, \mathfrak{U}_{1}^{\dagger}(\boldsymbol{x}_{\perp}) \int\limits_{\boldsymbol{p}_{\perp}} e^{i\boldsymbol{p}_{\perp} \cdot \boldsymbol{x}_{\perp}} \, \widetilde{\mathfrak{U}}_{1}(\boldsymbol{p}_{\perp} + \boldsymbol{k}_{\perp}) \left(\frac{p_{\perp}^{2} \tau}{2 \boldsymbol{k}_{\perp}} \right)^{i\nu} \left(\delta^{ij} - 2 \frac{p_{\perp}^{i} p_{\perp}^{j}}{p_{\perp}^{2}} \right) \boldsymbol{\varepsilon}_{\lambda}^{j} \\ \beta^{-i} &\equiv e^{-\frac{\pi \nu}{2}} \Gamma(i\nu) e^{i\nu\eta} \, \mathfrak{U}_{2}^{\dagger}(\boldsymbol{x}_{\perp}) \int\limits_{\boldsymbol{p}_{\perp}} e^{i\boldsymbol{p}_{\perp} \cdot \boldsymbol{x}_{\perp}} \, \widetilde{\mathfrak{U}}_{2}(\boldsymbol{p}_{\perp} + \boldsymbol{k}_{\perp}) \left(\frac{p_{\perp}^{2} \tau}{2 \boldsymbol{k}_{\perp}} \right)^{-i\nu} \left(\delta^{ij} - 2 \frac{p_{\perp}^{i} p_{\perp}^{j}}{p_{\perp}^{2}} \right) \boldsymbol{\varepsilon}_{\lambda}^{j} \end{split}$$

- Linearized EOM and Gauss' law satisfied up to terms of order (Q_sτ)²
- Fock-Schwinger gauge condition ($a^{\tau} = e^{\tau} = 0$)
- Evolved from plane waves in the remote past

Francois Gelis

Computational cost



Initial Conditions

Naive :

$$N \log(N) \times L^4 \log(L) \times N_{confs}$$

Better algorithm :

$$N \log(N) \times L^4 \times (\log(L) + N_{confs})$$

Time evolution

$$N \times L^2 \times N_{confs} \times N_{time\ steps}$$

Useful statistics (at fixed volume)

$$\sqrt{N_{\text{confs}}} ~\sim ~ \frac{g^2}{(\alpha_\perp \alpha_n)^2}$$

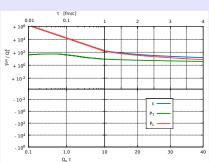
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Ultraviolet subtractions



• Fixed spacing in $\eta \iff \Lambda_z \sim \tau^{-1}$

Bare ε and $P_{_L}$ diverge as τ^{-2} when $\tau \to 0^+$



Ultraviolet subtractions

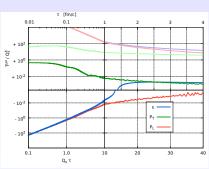


• Fixed spacing in η $\iff \Lambda_z \sim \tau^{-1}$

Bare ε and $P_{_L}$ diverge as τ^{-2} when $\tau \to 0^+$

• Zero point energy $\sim \Lambda_{\perp}^2 \Lambda_z^2$:

Subtracted by redoing the calculation with the sources turned off



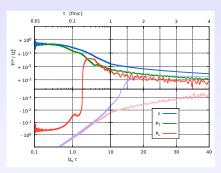
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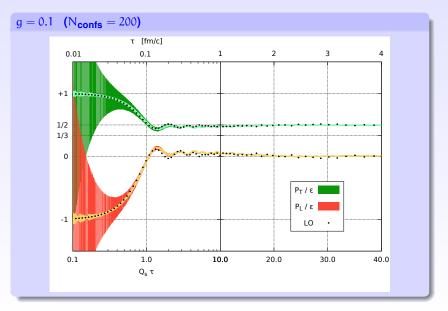
Zero point energy ~ Λ²_⊥Λ²_z :
 Subtracted by redoing the calculation with the sources turned off



• Subleading divergences $\sim \Lambda_z^2$ in ϵ and P_L :
Exist only at finite \perp lattice spacing (not in the continuum)
Same counterterm in ϵ and P_L to preserve $T^\mu{}_\mu = 0$ Must be of the form $A \times \tau^{-2}$ to preserve Bjorken's law
At the moment, not calculated from first principles $\Rightarrow A$ fitted

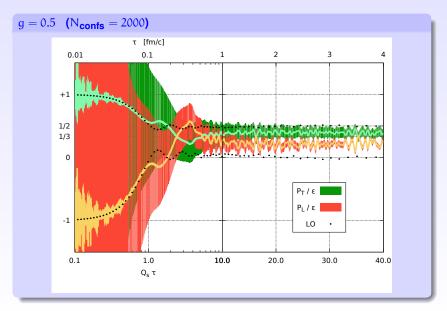
Time evolution of $P_{_{\rm T}}/\varepsilon$ and $P_{_{\rm T}}/\varepsilon$ (64 \times 64 \times 128 lattice)





Time evolution of $P_{_{\rm T}}/\varepsilon$ and $P_{_{\rm T}}/\varepsilon$ (64 \times 64 \times 128 lattice)





Summary

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- CGC calculation of the energy momentum tensor in a nucleus-nucleus collision, up to $Q_s \tau \lesssim 20$
- · Method:
 - · Classical statistical method
 - Initial Gaussian fluctuations : analytical, from a 1-loop calculation
 - Time evolution: numerical, 3+1d Yang-Mills equations on a lattice
- Accuracy :

$$\left< 0_{in} \middle| \mathsf{T}^{\mu \nu}(\tau, x) \middle| 0_{in} \right> \ \, \text{at LO + NLO + leading secular terms}$$

- · Results:
 - Sizable longitudinal pressure $(P_{_{\rm T}}/P_{_{\rm T}} \sim 60\% \text{ for } q = 0.5)$
 - Typical timescale : $Q_s \tau \sim 2 3$