



# Introduction and Current Status of 2HDMs

**Yuji Omura (KMI, Nagoya Univ.)**

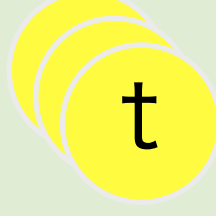
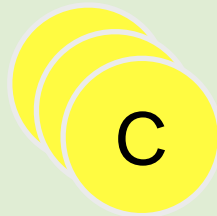
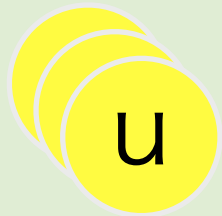
# Background and Motivation

# Standard Model $(\text{SU}(3)_c \times \text{SU}(2)_L \times \text{U}(1)_Y)$

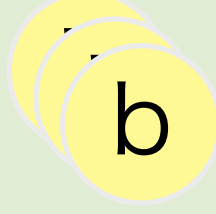
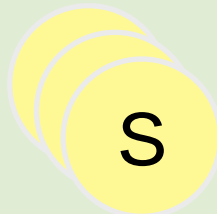
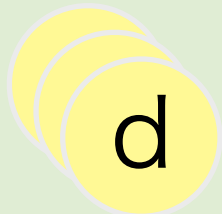
is very successful in particle physics

spin-1/2

quarks  $\text{SU}(3)_c$ -charged



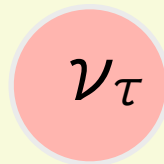
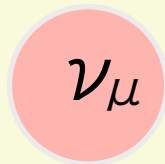
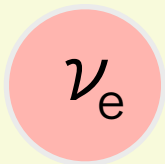
$+2/3$



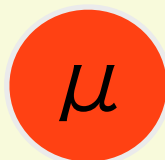
$-1/3$

EM-charge

leptons



0



-1

spin-1

$\text{SU}(3)_c$  gauge



gluon

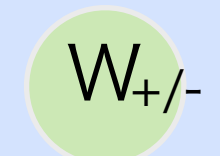
$\text{SU}(2)_L \times \text{U}(1)_Y$



massive



photon



massive

carry forces

spin-0

Higgs



breaks

$\text{SU}(2)_L \times \text{U}(1)_Y$

# Why do we study BSM?

*We expect that there are something behind*

*EW scale*

*flavor structure*

*strong CP problem*

*baryon asymmetry*

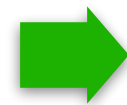
*gauge symmetry*

*dark matter*

# Possible BSMs

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*SUSY etc.*

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➡  $U(1)_{PQ}$ , LR symmetry

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*gauge symmetry* ➡ GUT

*dark matter*

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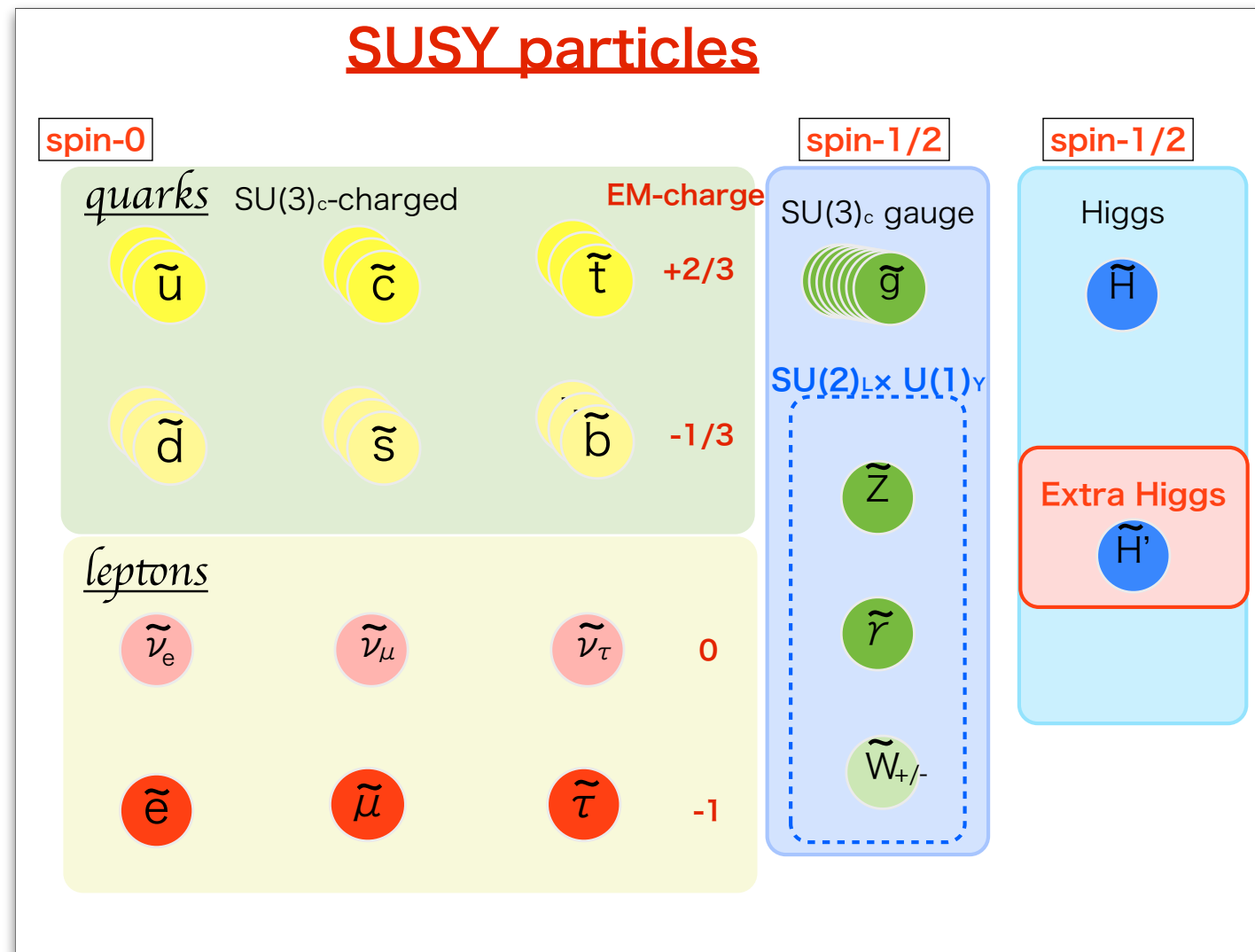
*gauge symmetry* ➡ GUT

*dark matter* ➡ extra stable particles

*There are many interesting candidates for BSMs,  
and we are looking for the evidences of new physics*

There are many interesting candidates for BSMs,  
and we are looking for the evidences of new physics

SUSY predicts so many new scalars:



*There are many interesting candidates for BSMs,  
and we are looking for the evidences of new physics*

flavor symmetry *predicts extra scalars*

quark Yukawa in SM:  $\overline{Q}_L^i H_1 y_d^i d_R^i + \overline{Q}_L^i (V^\dagger)^{ij} \tilde{H}_1 y_u^j u_R^j$



In flavor symmetric model,

$$y_d^{ij}(\Phi) \overline{Q}_L^i H_1 d_R^j + y_u^{ij}(\Phi) \overline{Q}_L^i \tilde{H}_1 u_R^j$$

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In flavor symmetric model,

or  $\overline{Q}_L^i H_1 y_d^i d_R^i + y_u^i \overline{Q}_L^i \tilde{H}_{ij} u_R^j$

There are many interesting candidates for BSMs,  
and we are looking for the evidences of new physics

$U(1)_{PQ}$  and LR symmetry *predict extra Higgs*

$$U(1)_{PQ} : H_1 \rightarrow e^{iq_1\theta} H_1, H_2 \rightarrow e^{iq_2\theta} H_2$$

$$\overline{Q}_L^i H_2 y_d^i d_R^i + \overline{Q}_L^i (V^\dagger)^{ij} \tilde{H}_1 y_u^j u_R^j$$

[Zhitnitsky SJNP 31 (1980); Dine, Fischler, Srednicki PLB 104 (1981)]

$$\text{LR symmetry: } Q_L^i \leftrightarrow Q_R^i$$

[Babu, Mohapatra, PRD41, 1286 (1990)]

$$Y_{ij} \overline{\hat{Q}}_L^i \Phi \hat{Q}_R^j$$

$$\text{where } \Phi = (\tilde{H}_u, H_d) \quad \hat{Q}_R^j = (\hat{u}_R^j, \hat{d}_R^j)^T$$

*There are many interesting candidates for BSMs,  
and we are looking for the evidences of new physics*

**GUT** *also predict extra Higgs*

$$Y_{10}^{ij} \mathbf{16}_i \mathbf{16}_j \mathbf{10} + Y_{126}^{ij} \mathbf{16}_i \mathbf{16}_j \overline{\mathbf{126}} + Y_{120}^{ij} \mathbf{16}_i \mathbf{16}_j \mathbf{120}.$$

[Georgi, Nanopoulos, NPB159, 16 (1979); Georgi, Jarlskog, Phys.Lett.86B, 297 (1979); et.al]



There are many interesting candidates for BSMs,  
and we are looking for the evidences of new physics

baryon asymmetry may also require extra Higgs

See, for instance, Dorsch, Huber, No, 1305.6610; Haarr, Kvellestad, Petersen, 1611.05757; Fuyoto, Hou, Senaha, 1705.05034,

There are many dark matter models with extra Higgs doublets.

Inert doublet model

$$H_1 = \begin{pmatrix} G^+ \\ \frac{v+h+iG}{\sqrt{2}} \end{pmatrix}, \quad H_2 = \begin{pmatrix} H^+ \\ \frac{H+iA}{\sqrt{2}} \end{pmatrix} \text{ Lighter one is DM}$$

**Extra Higgs doublet** may be a promising candidate  
that will be discovered near future.

There are actually so many works on the physics of the extra Higgs doublets.

From phenomenological point of view, the extension is “very easy”, but quite attractive and interesting:

- simple but have rich phenomenology: Higgs physics, flavor physics, etc.
- do not break the gauge anomaly-free conditions,
- do not break the SM prediction so much: ( $\rho$  para. = 1 @tree)
- may explain the discrepancies in flavor physics.

# Contents

- Setup of 2HDM
- Hunting for the scalars
- About the discrepancies in 2HDMs
- What can we expect from the underlying theory?
- Summary

# Setup of 2HDM

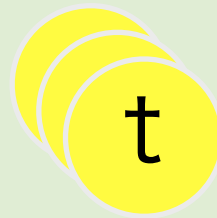
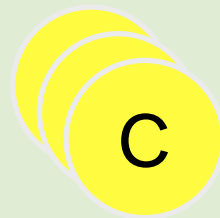
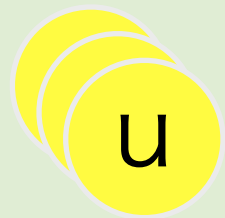
# Let me start from the SM.

spin-1/2

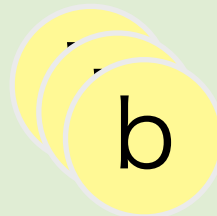
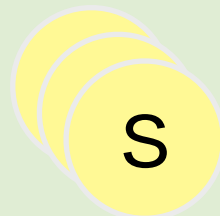
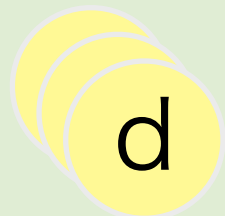
quarks

$SU(3)_c$ -charged

EM-charge

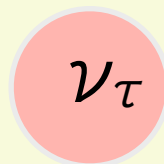
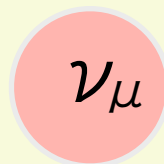
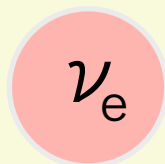


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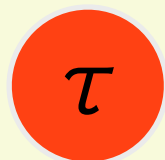


$-1/3$

leptons



$0$



$-1$

spin-1

$SU(3)_c$  gauge



gluon

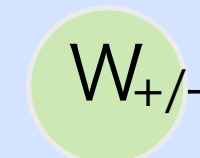
$SU(2)_L \times U(1)_Y$



massive



photon



massive

carry forces

spin-0

Higgs



breaks

$SU(2)_L \times U(1)_Y$

# Add just one extra Higgs doublet.

spin-1/2

quarks SU(3)<sub>c</sub>-charged

EM-charge

u c t +2/3

d s b -1/3

leptons

$\nu_e$   $\nu_\mu$   $\nu_\tau$  0

e  $\mu$   $\tau$  -1

spin-1

SU(3)<sub>c</sub> gauge

g  
gluon

SU(2)<sub>L</sub> × U(1)<sub>Y</sub>

Z  
massive

$\gamma$   
photon

W<sub>±</sub>  
massive

carry forces

spin-0

Higgs

H

Extra Higgs

H'

breaks  
SU(2)<sub>L</sub> × U(1)<sub>Y</sub>

*Both Higgs doublets generally obtain non-vanishing VEVs.*

$$\langle H_1 \rangle = \begin{pmatrix} 0 \\ \frac{1}{\sqrt{2}}v_1 \end{pmatrix}, \quad \langle H_2 \rangle = \begin{pmatrix} 0 \\ \frac{1}{\sqrt{2}}v_2 \end{pmatrix}$$

*Both VEVs contribute to  $Z$  and  $W$  masses.*

$$M_Z^2 = \frac{1}{4}g_Z^2(v_1^2 + v_2^2)$$

$$M_W^2 = \frac{1}{4}g^2(v_1^2 + v_2^2)$$

The 2 doublets predict extra scalars:

$$H_1 = \begin{pmatrix} H_1^+ \\ \frac{v_1 + h_1 + iA_1}{\sqrt{2}} \end{pmatrix}, \quad H_2 = \begin{pmatrix} H_2^+ \\ \frac{v_2 + h_2 + iA_2}{\sqrt{2}} \end{pmatrix}$$

Mass eigenstates



$$\hat{H}_1 = \begin{pmatrix} \text{Goldstones } G^+ \\ \frac{v + \phi_1 + iG}{\sqrt{2}} \end{pmatrix}, \quad \hat{H}_2 = \begin{pmatrix} \text{charged Higgs } H^+ \\ \frac{\phi_2 + iA}{\sqrt{2}} \end{pmatrix}$$

2 CP-even scalars      pseudo-scalar

$v = \sqrt{v_1^2 + v_2^2}$

Mass base of CP-even is different from  $H^+$  and  $A$ .

$$\begin{pmatrix} \phi_1 \\ \phi_2 \end{pmatrix} = \begin{pmatrix} \cos \theta_{\beta\alpha} & \sin \theta_{\beta\alpha} \\ -\sin \theta_{\beta\alpha} & \cos \theta_{\beta\alpha} \end{pmatrix} \begin{pmatrix} H \\ h \end{pmatrix}$$

heavy Higgs

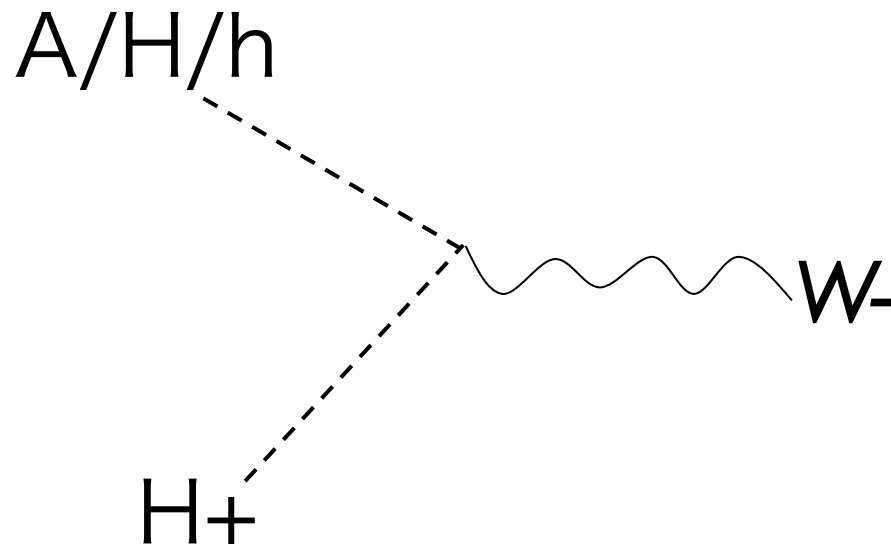
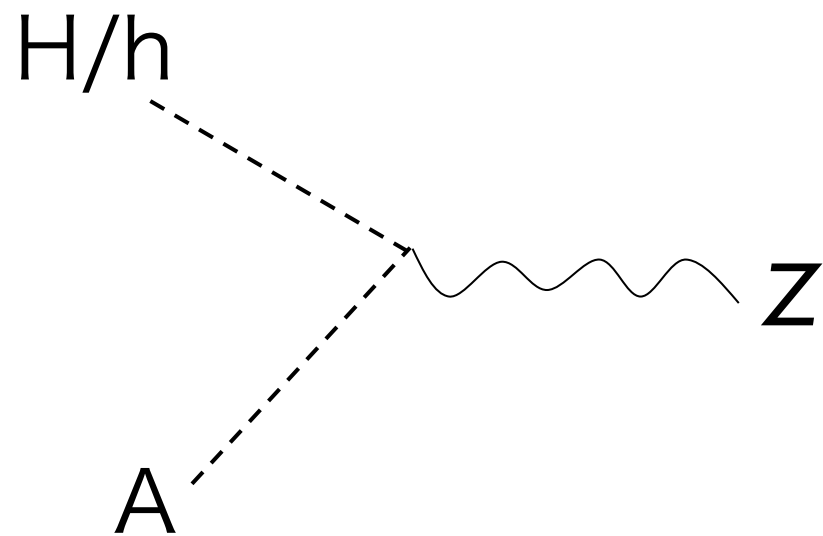
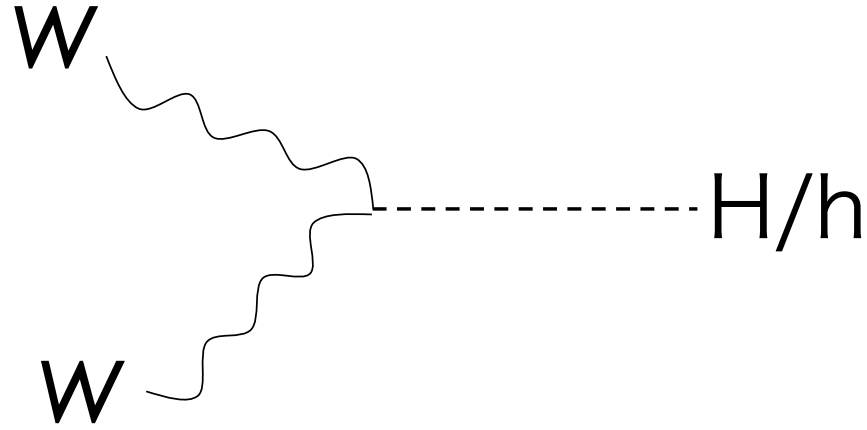
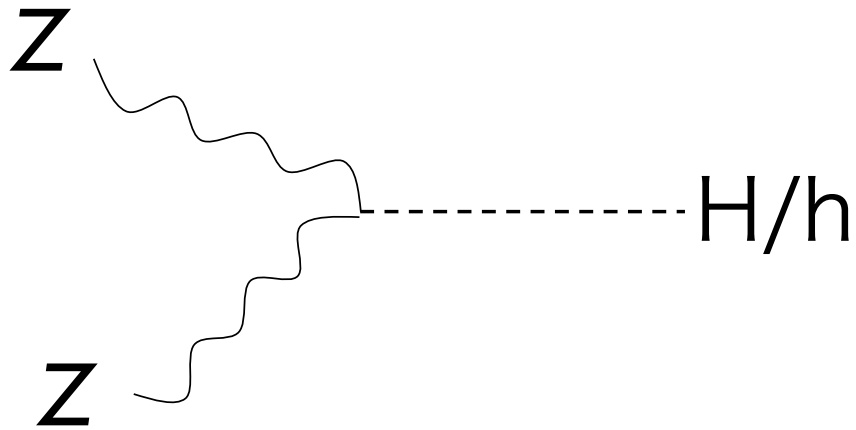
$\sin \theta_{\beta\alpha} \rightarrow 1$  (SM limit)

125-GeV Higgs



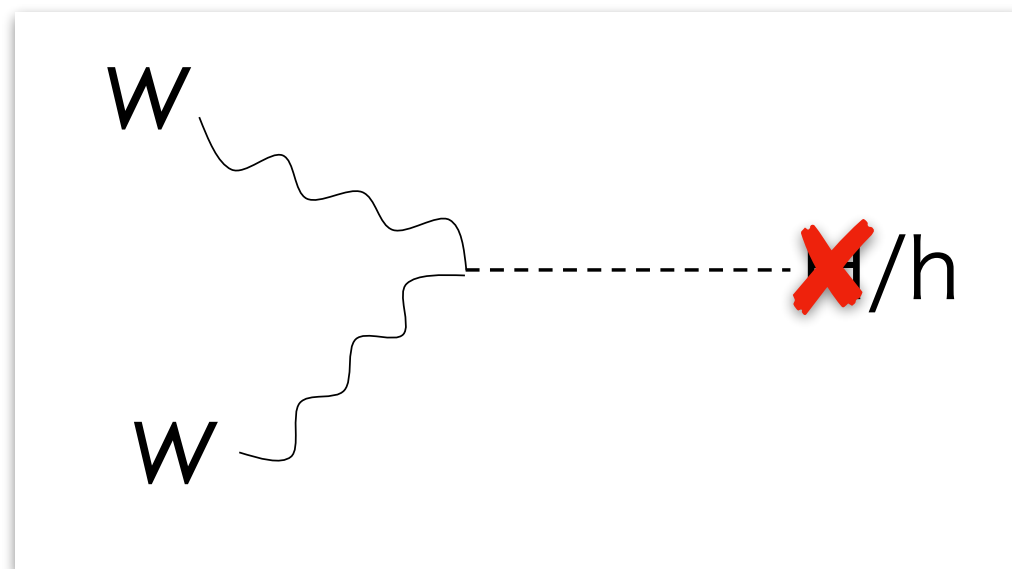
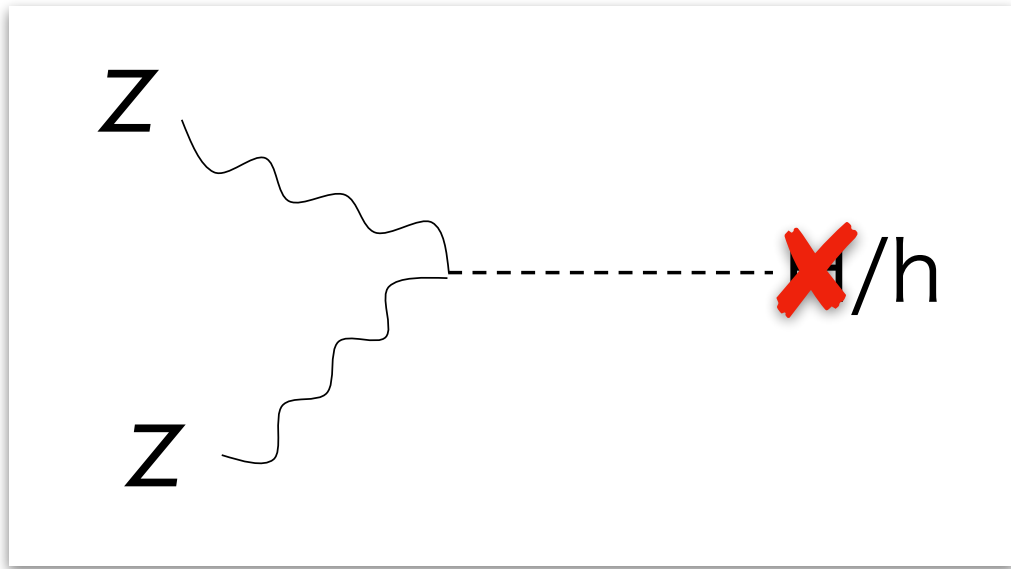
# How the scalars interact with the SM?

## EW gauge interaction @tree

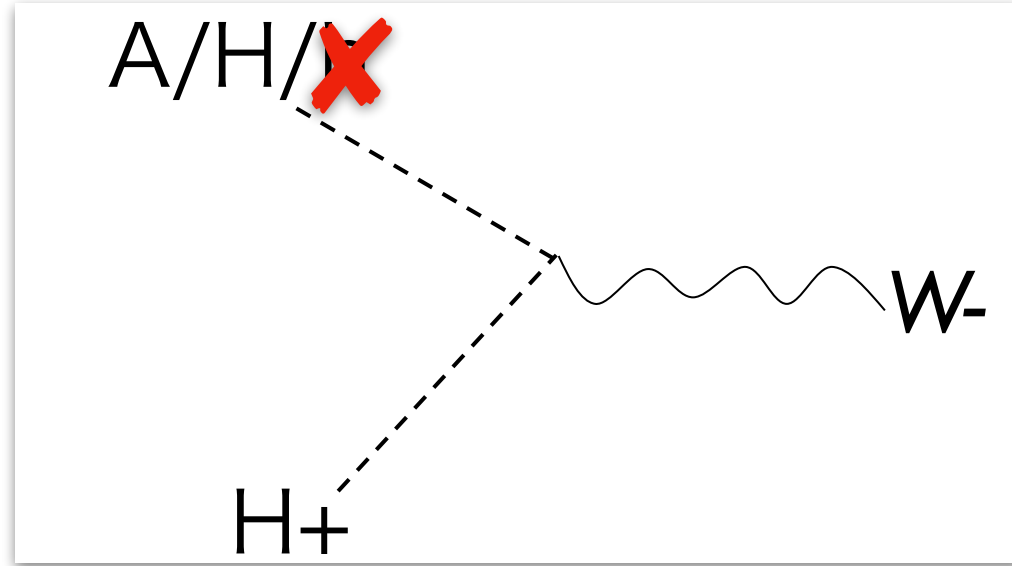
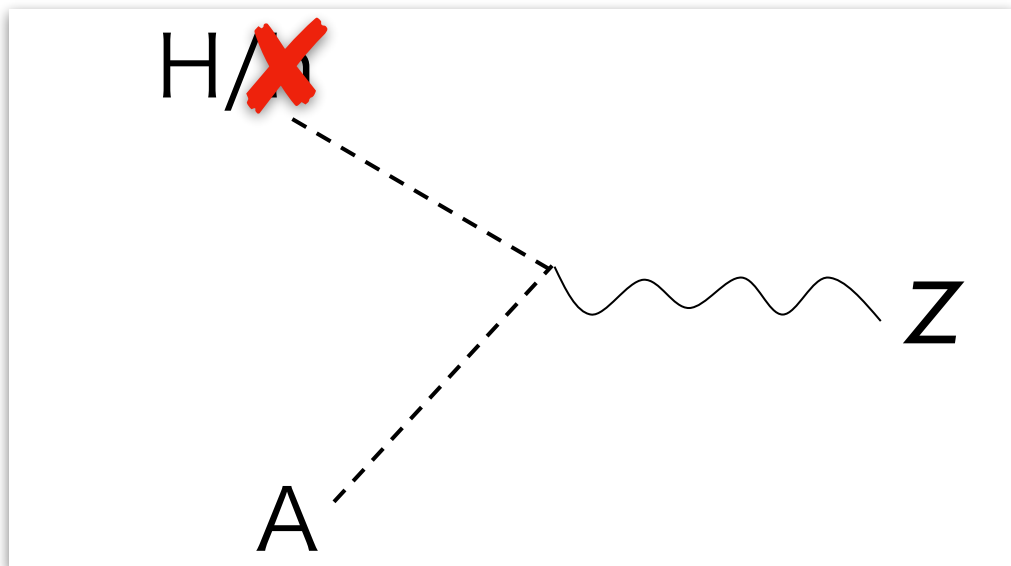


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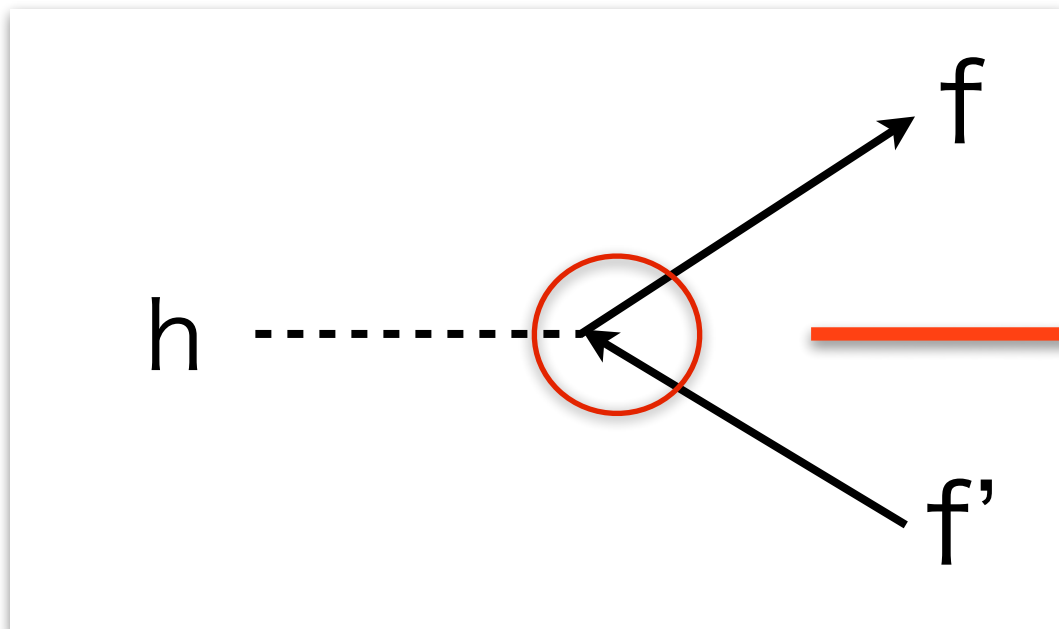
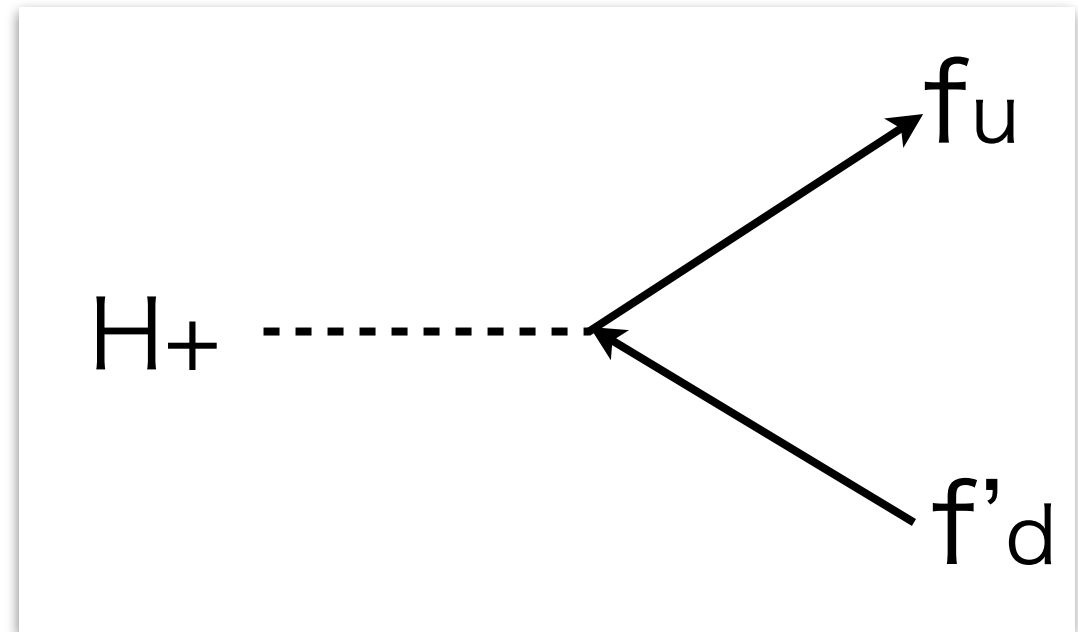
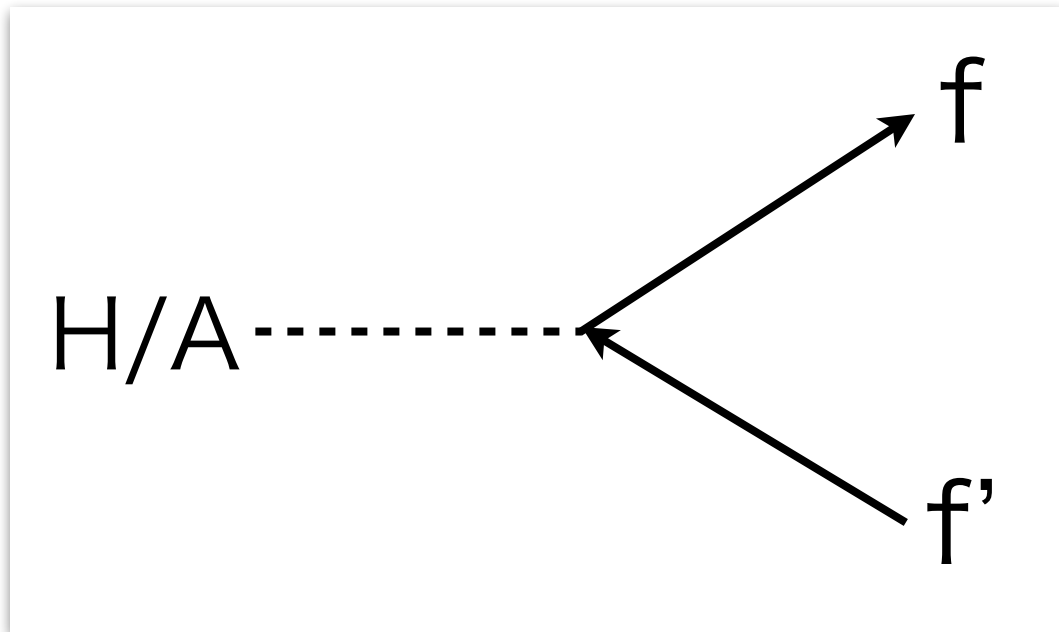


**In the SM limit**

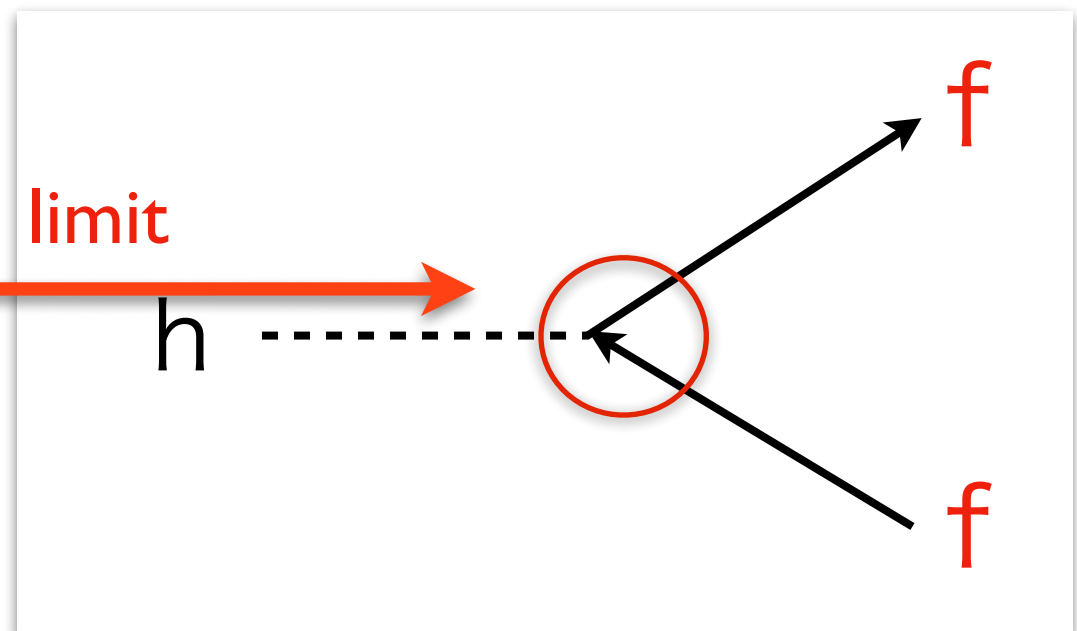


# How the scalars interact with the SM?

## Yukawa interaction



SM limit



# How heavy are they?

*There is a source to shift the extra scalar masses:*

$$V \supset M^2 H_1^\dagger H_2$$

*They can be heavy as much as we want, in principle.*

*The mass difference is at most EW scale:*

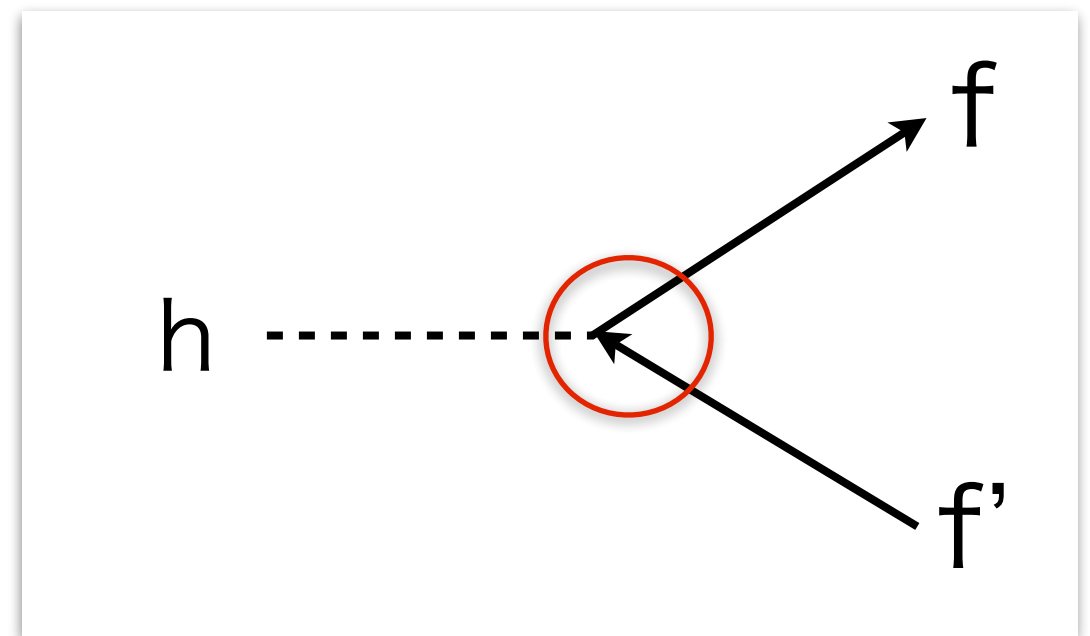
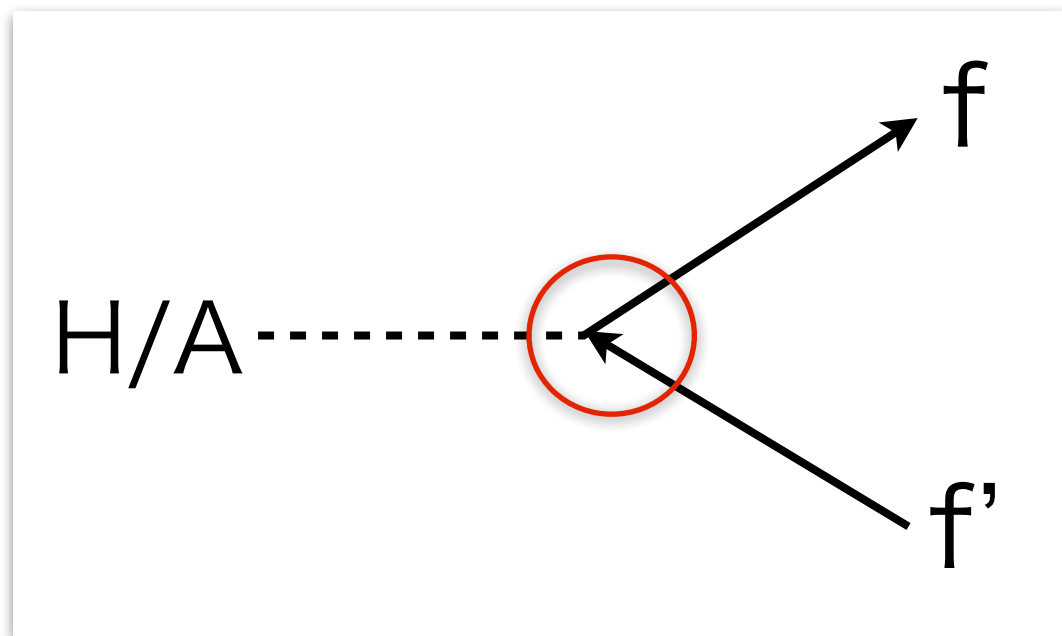
$$m_H^2 \simeq m_A^2 + \mathcal{O}(v^2)$$

$$m_{H_\pm}^2 \simeq m_A^2 + \mathcal{O}(v^2)$$

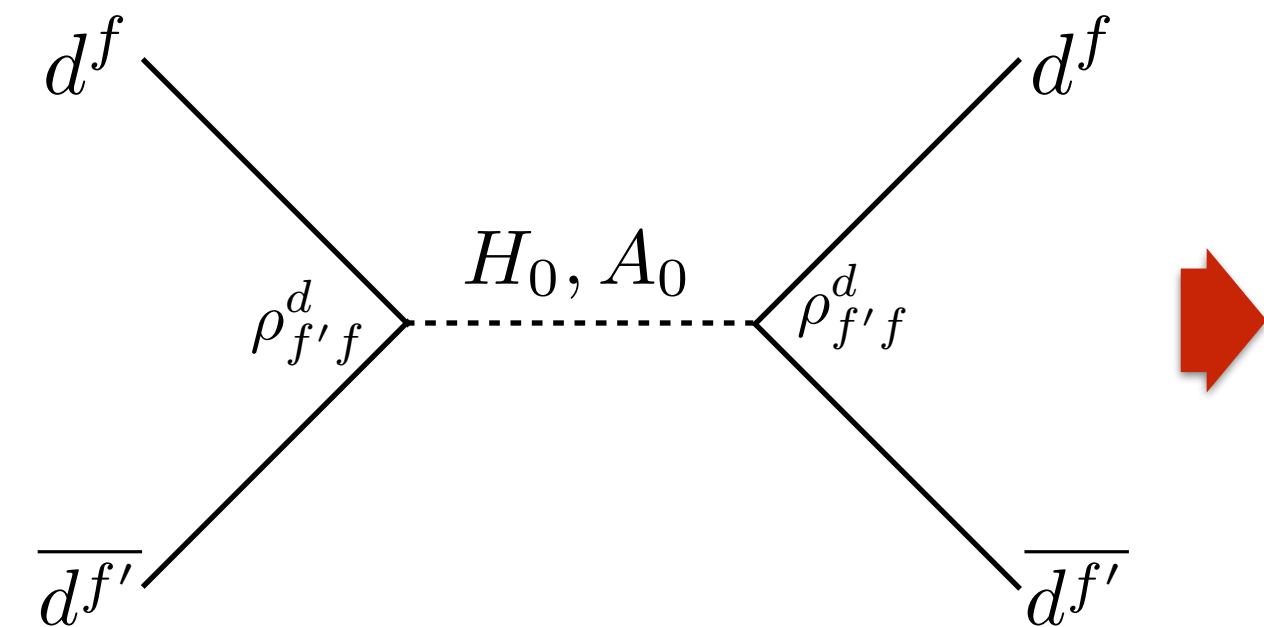
Hunt for the scalars

Strong constrain comes from flavor physics

In general, there are new flavor violating couplings



# Strong constrain comes from flavor physics



## Constraints from $\Delta F=2$ processes

$$\mathcal{H}_{eff} = C_4(\bar{q}_L q'_R)(\bar{q}_R q'_L)$$

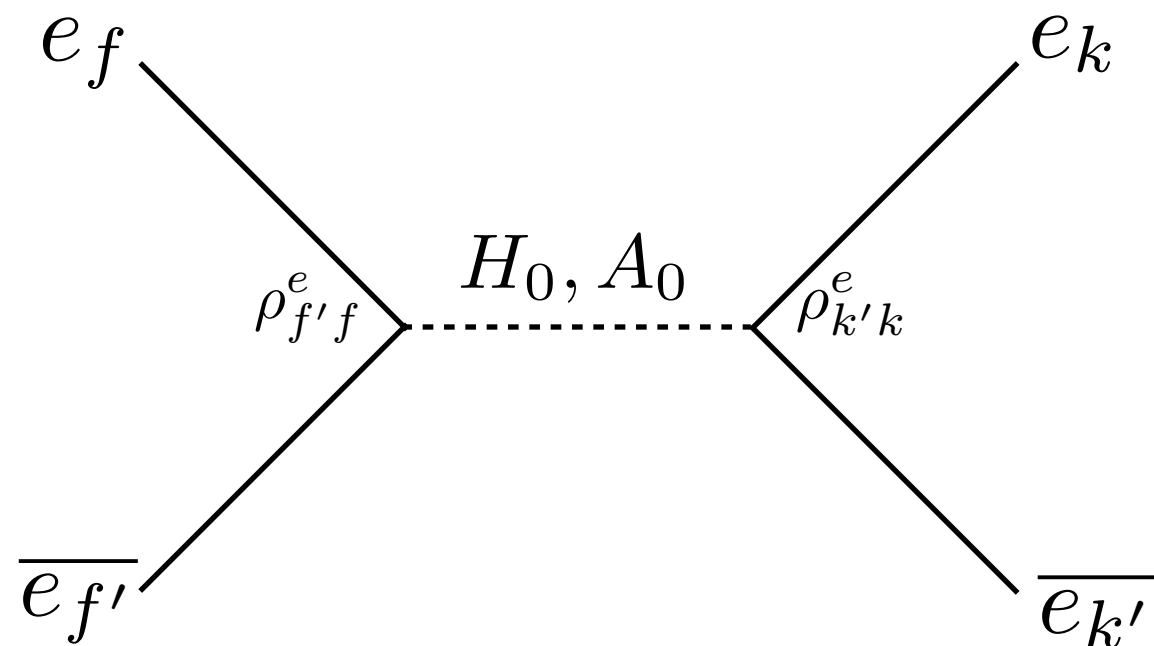
$$K - \bar{K} : m_H / \sqrt{\rho_{sd}^d \rho_{ds}^{d*}} \gtrsim \mathcal{O}(10^5) \text{ TeV}$$

$$B_d - \bar{B}_d : m_H / \sqrt{\rho_{bd}^d \rho_{db}^{d*}} \gtrsim \mathcal{O}(10^3) \text{ TeV}$$

$$B_s - \bar{B}_s : m_H / \sqrt{\rho_{bs}^d \rho_{sb}^{d*}} \gtrsim \mathcal{O}(10^2) \text{ TeV}$$

$$D - \bar{D} : m_H / \sqrt{\rho_{cu}^u \rho_{uc}^{u*}} \gtrsim \mathcal{O}(10^3) \text{ TeV}$$

# Strong constrain comes from flavor physics



## LFV constraints

$$\mathcal{H}_{eff} = C_4 (\overline{l'_L} l_R) (\overline{l_R^1} l_L^2)$$

From  $\mu \rightarrow 3e$

$$m_H / (\sqrt{\rho_{e\mu}^e \rho_{ee}^e}) \gtrsim 150 \text{ TeV}$$

From  $\tau \rightarrow l' l l$  ( $l', l = \mu, e$ )

$$m_H / (\sqrt{\rho_{\tau l'}^e \rho_{ll}^e}) \gtrsim 10 \text{ TeV}$$



If we assign the symmetry to distinguish the two doublets, we can evade the strong bounds.

$$U(1)_{PQ} : H_1 \rightarrow e^{iq_1\theta} H_1, H_2 \rightarrow e^{iq_2\theta} H_2$$

$$Z_2 : H_1 \rightarrow +H_1, H_2 \rightarrow -H_2$$

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► no new flavor violating coupling

$$\overline{Q}_L^i H_2 y_d^i d_R^i + \overline{Q}_L^i (V^\dagger)^{ij} \tilde{H}_1 y_u^j u_R^j$$

but this symmetry causes a problem:

$$V \supset \cancel{M}^2 H_1^\dagger H_2$$

scalar masses become EW-scale!

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$$\overline{Q}_L^i H_2 y_d^i d_R^i + \overline{Q}_L^i (V^\dagger)^{ij} \tilde{H}_1 y_u^j u_R^j$$

people often say that

“ $Z_2$  is broken in only mass para.,” and allow

$$V \supset M^2 H_1^\dagger H_2 \text{ to shift the scalar masses}$$

Then, the Yukawa couplings are controlled by the symmetry.

There are several choices for the charge assignment:

**Type-I 2HDM:**  $Y_{ff'}^u \overline{Q}_L^f \widetilde{H}_1 u_R^{f'} + Y_{ff'}^d \overline{Q}_L^f H_1 d_R^{f'} + Y_{ff'}^e \overline{l}_L^f H_1 e_R^{f'}$

**Type-II 2HDM:**  $Y_{ff'}^u \overline{Q}_L^f \widetilde{H}_1 u_R^{f'} + Y_{ff'}^d \overline{Q}_L^f H_2 d_R^{f'} + Y_{ff'}^e \overline{l}_L^f H_2 e_R^{f'}$

*(inspired by MSSM)*

**Type-X 2HDM:**  $Y_{ff'}^u \overline{Q}_L^f \widetilde{H}_1 u_R^{f'} + Y_{ff'}^d \overline{Q}_L^f H_1 d_R^{f'} + Y_{ff'}^e \overline{l}_L^f H_2 e_R^{f'}$

**Type-Y 2HDM:**  $Y_{ff'}^u \overline{Q}_L^f \widetilde{H}_1 u_R^{f'} + Y_{ff'}^d \overline{Q}_L^f H_2 d_R^{f'} + Y_{ff'}^e \overline{l}_L^f H_1 e_R^{f'}$

(flipped)

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There are several choices for the charge assignment:

**Let me focus on type-II**

**Type-II 2HDM:**  $Y_{ff'}^u \overline{Q}_L^f \widetilde{H}_1 u_R^{f'} + Y_{ff'}^d \overline{Q}_L^f H_2 d_R^{f'} + Y_{ff'}^e \overline{l}_L^f H_2 e_R^{f'}$

*(inspired by MSSM)*

Type-X 2HDM:  $Y_{ff'}^u \overline{Q}_L^f \widetilde{H}_1 u_R^{f'} + Y_{ff'}^d \overline{Q}_L^f H_1 d_R^{f'} + Y_{ff'}^e \overline{l}_L^f H_2 e_R^{f'}$

Type-Y 2HDM:  $Y_{ff'}^u \overline{Q}_L^f \widetilde{H}_1 u_R^{f'} + Y_{ff'}^d \overline{Q}_L^f H_2 d_R^{f'} + Y_{ff'}^e \overline{l}_L^f H_1 e_R^{f'}$   
(flipped)

Even if there is no new flavor violating couplings,  
the CKM predicts enough large deviations in flavor physics:

**Experiment:** ( $E_\gamma > 1.6\text{GeV}$ )

$$Br(b \rightarrow s\gamma)_{\text{exp}} = (3.32 \pm 0.16) \times 10^{-4}$$

(Y.Amhis, arXiv:1612.07233)

**SM prediction (NNLO):** ( $E_\gamma > 1.6\text{GeV}$ )

$$Br(b \rightarrow s\gamma)_{\text{SM}} = (3.36 \pm 0.23) \times 10^{-4}$$

(Misiak, et.al., arXiv:1208.2788;1503.01789;1702.04571)

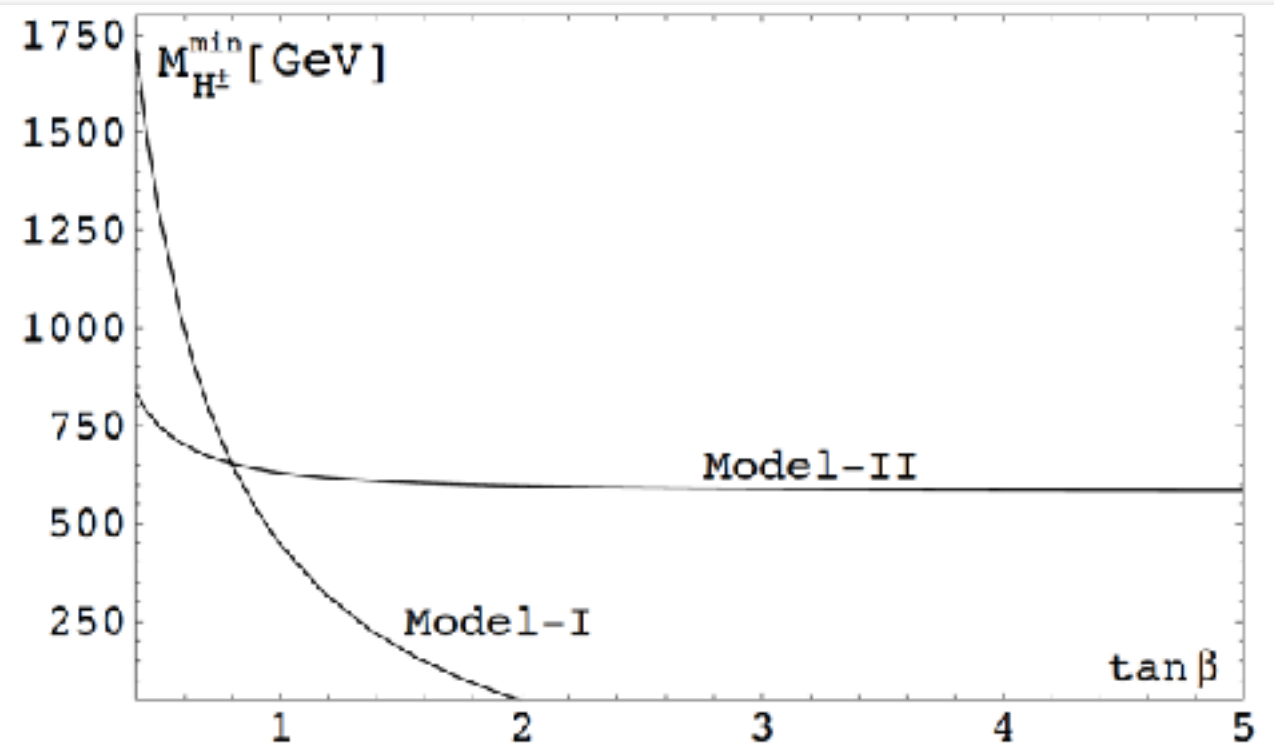
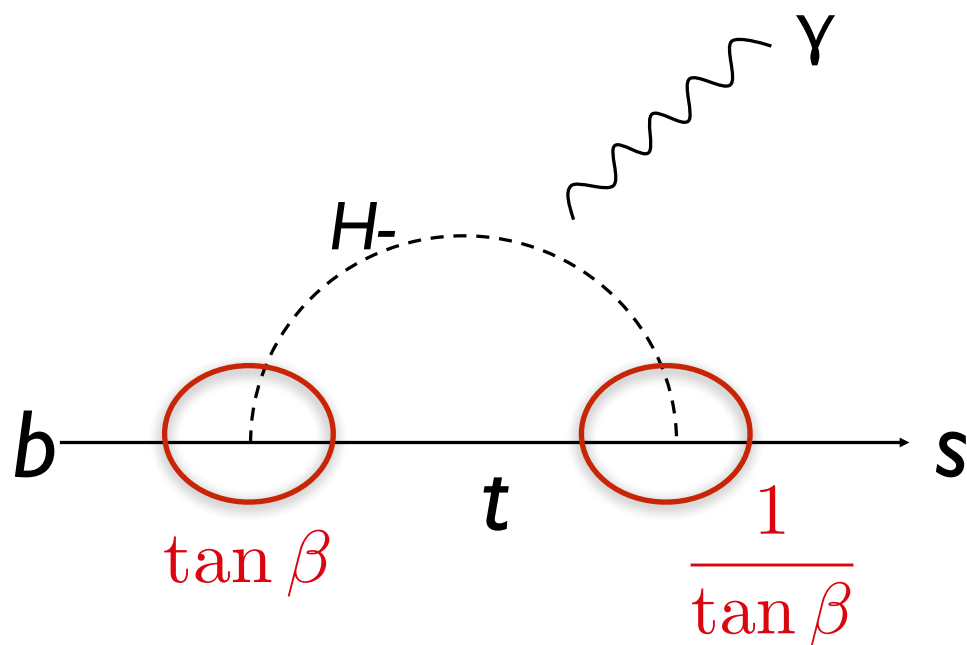


Figure 4: 95% C.L. lower bounds on  $M_{H^\pm}$  as functions of  $\tan \beta$ .

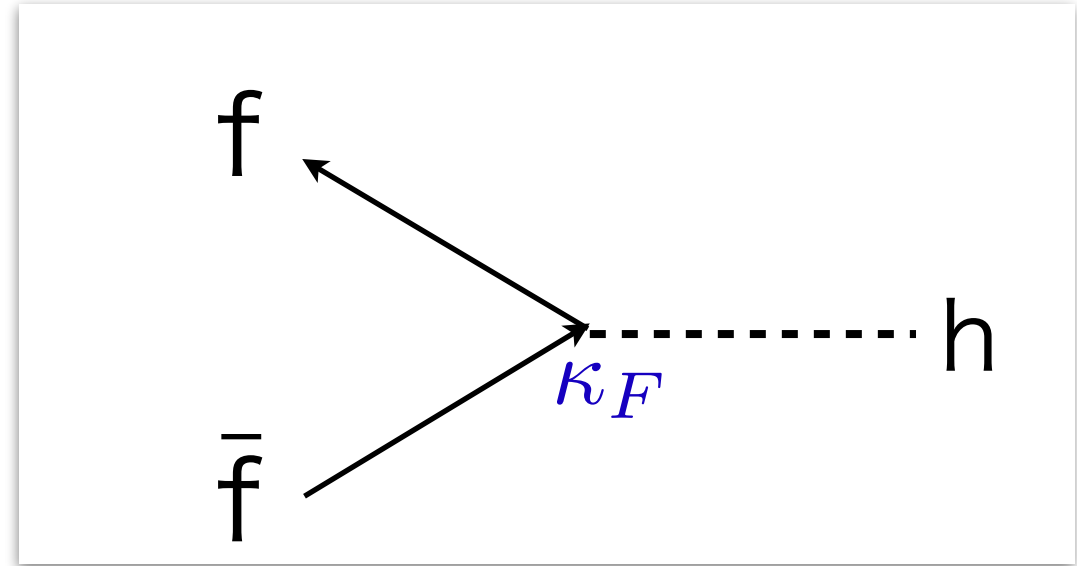
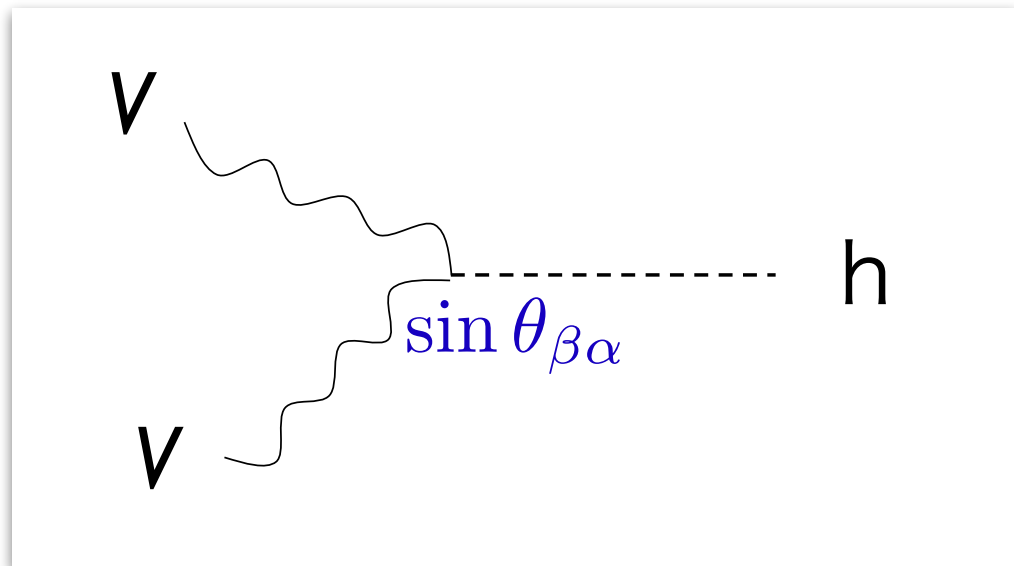
➡  $Q_7 = \frac{e}{16\pi^2} m_b (s_L \sigma^{\mu\nu} b_R) F_{\mu\nu}$

$$m_{H^\pm} \gtrsim 580\text{GeV}$$

(Belle, arXiv:1608.02344; Misiak, et.al., arXiv:1702.04571)

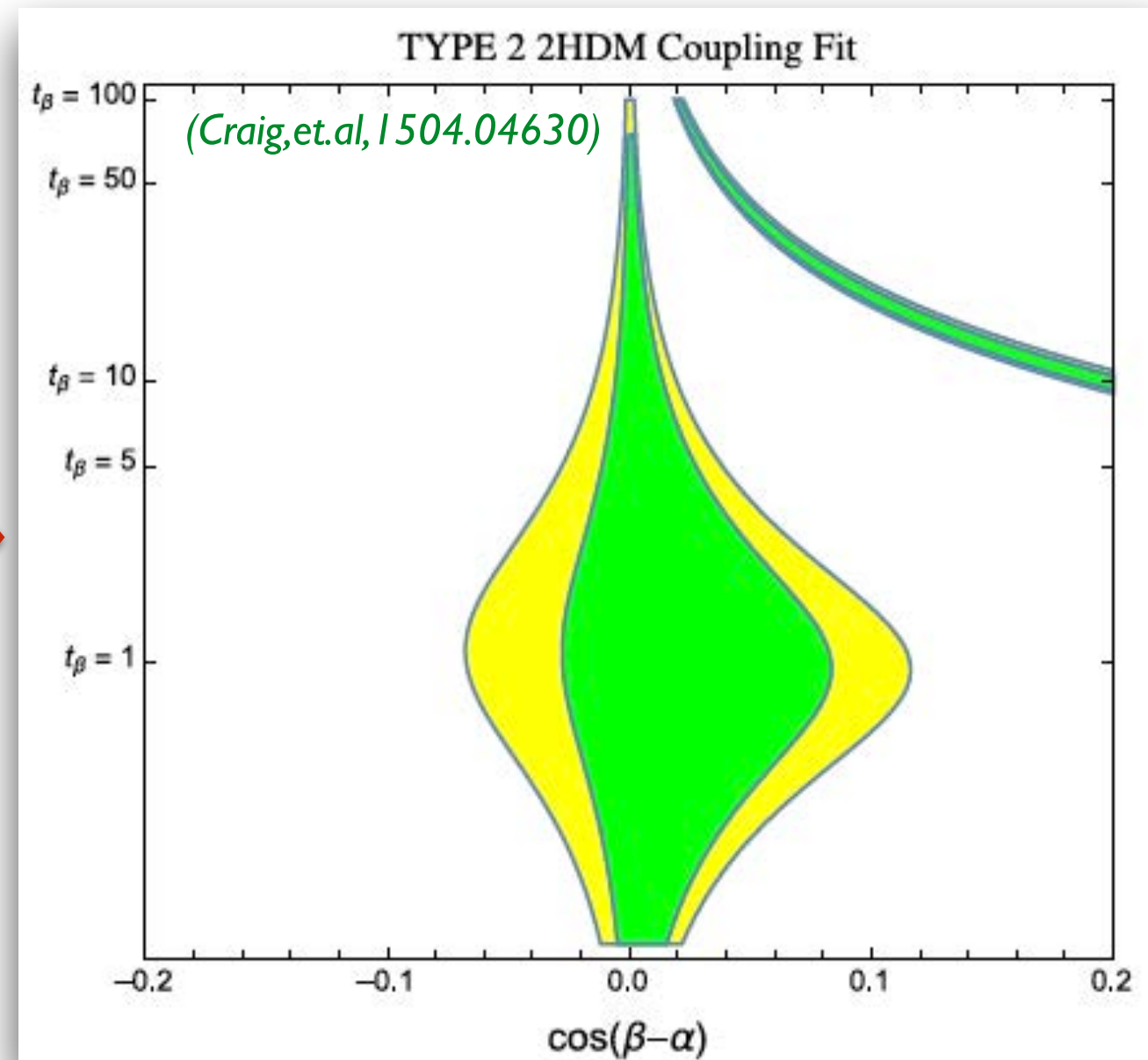
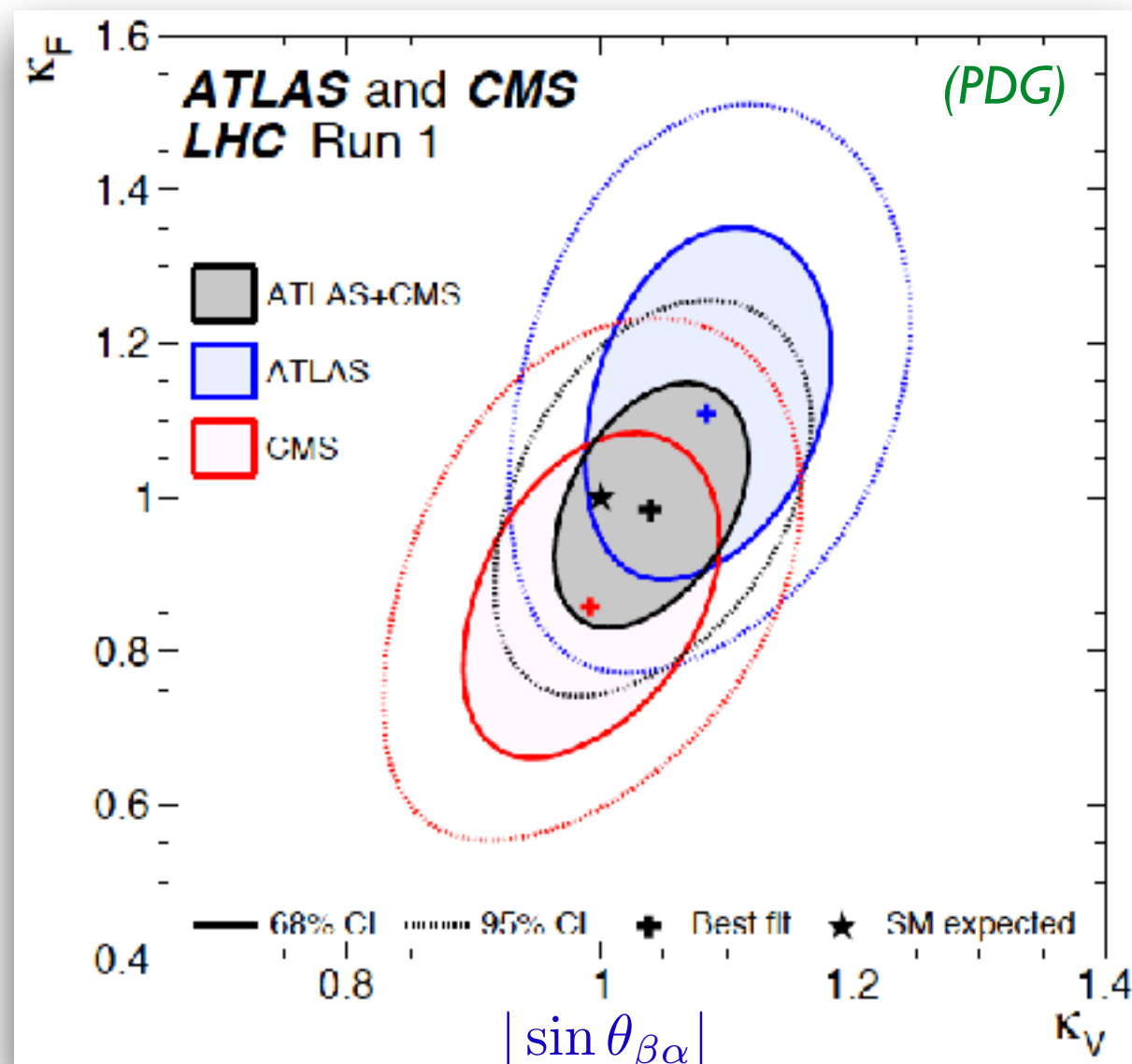
# We could see the scalars directly/indirectly at the LHC

*The SM prediction about the 125 GeV Higgs is deviated.*



# We could see the scalars directly/indirectly at the LHC

*The SM prediction about the 125 GeV Higgs is deviated.*

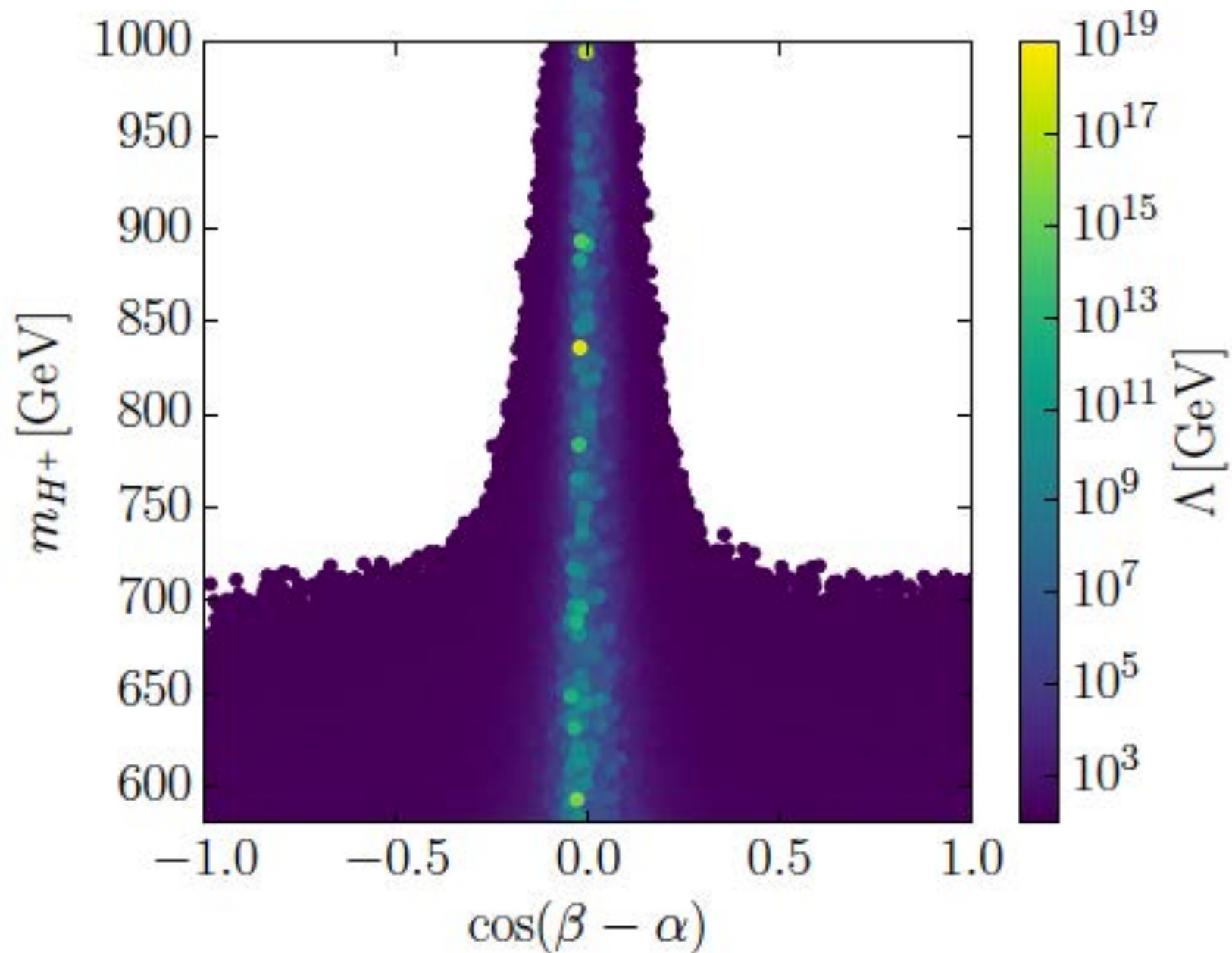


*Less than 10 % deviation is only allowed.*



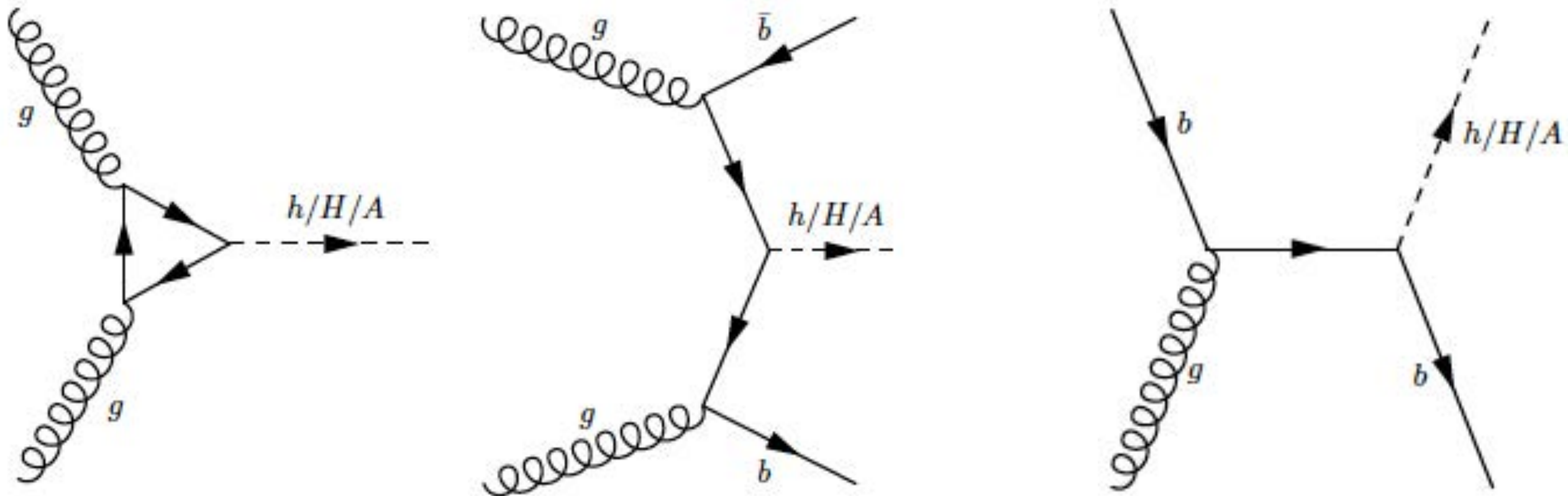
The Landau pole appears,  
if the Higgs signal is deviated from the SM one.

*(Philipp Basler , Pedro M. Ferreira, Margarete Muhlleitner, Rui Santos, 1710.10410)*



We could see the scalars directly at the LHC

*Then, the heavy scalars are produced at the LHC:*

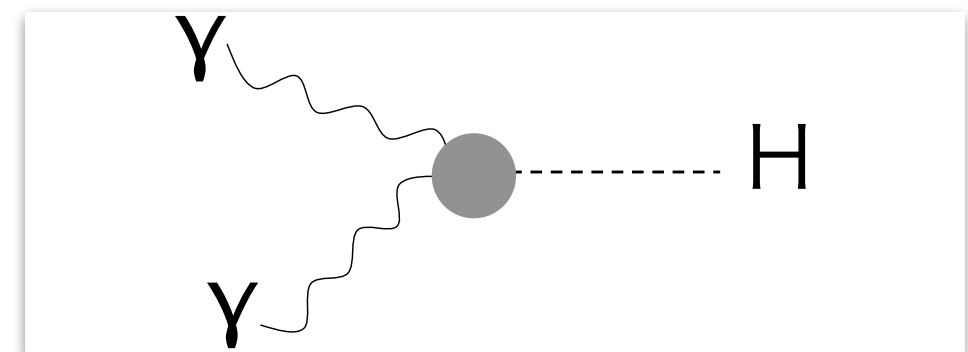
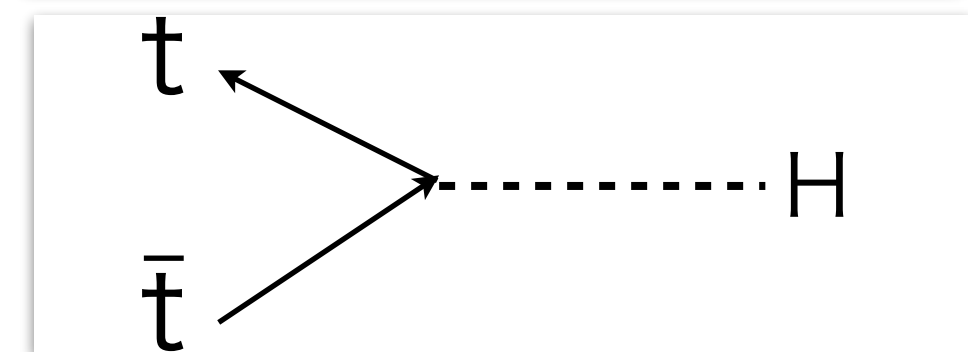
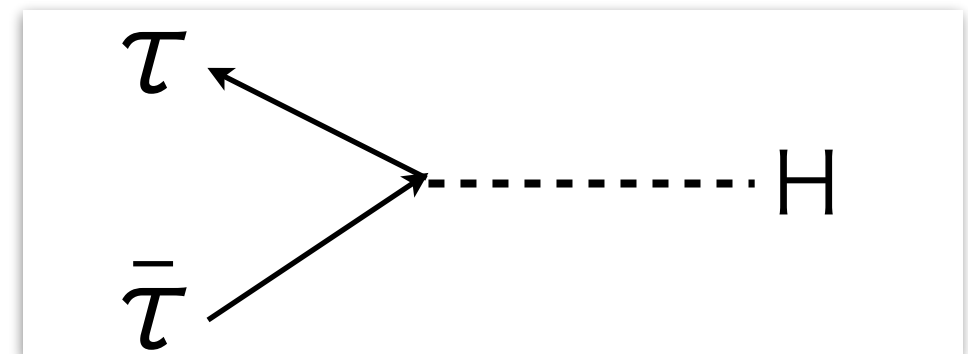
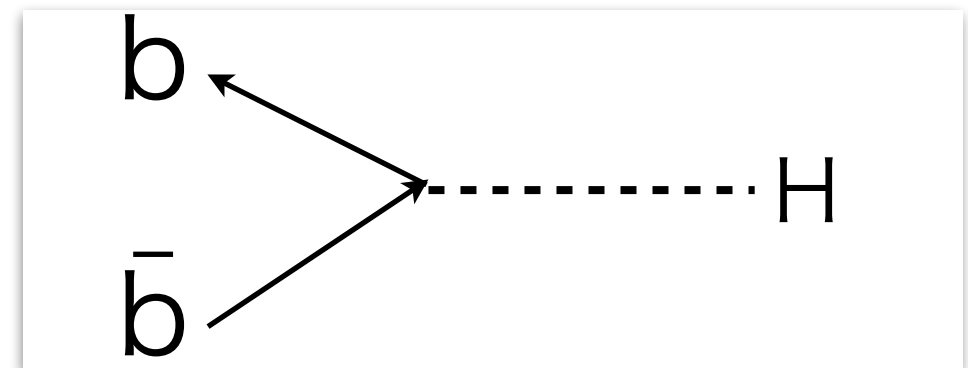
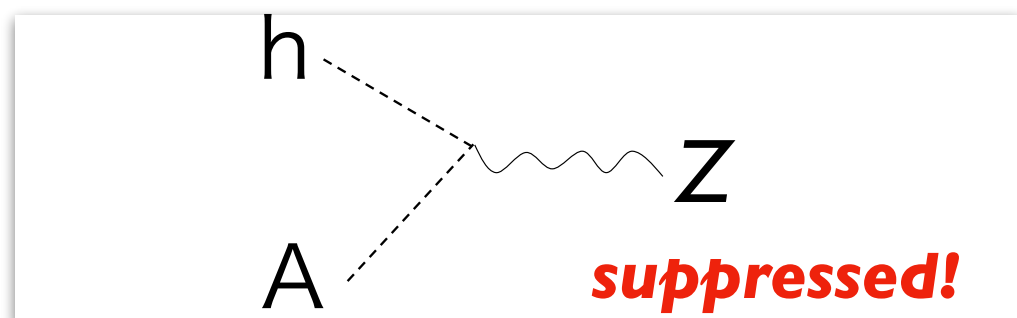
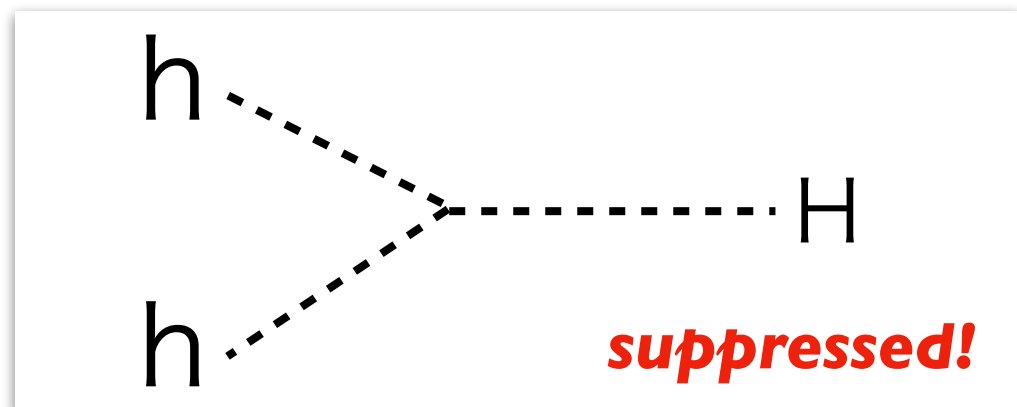
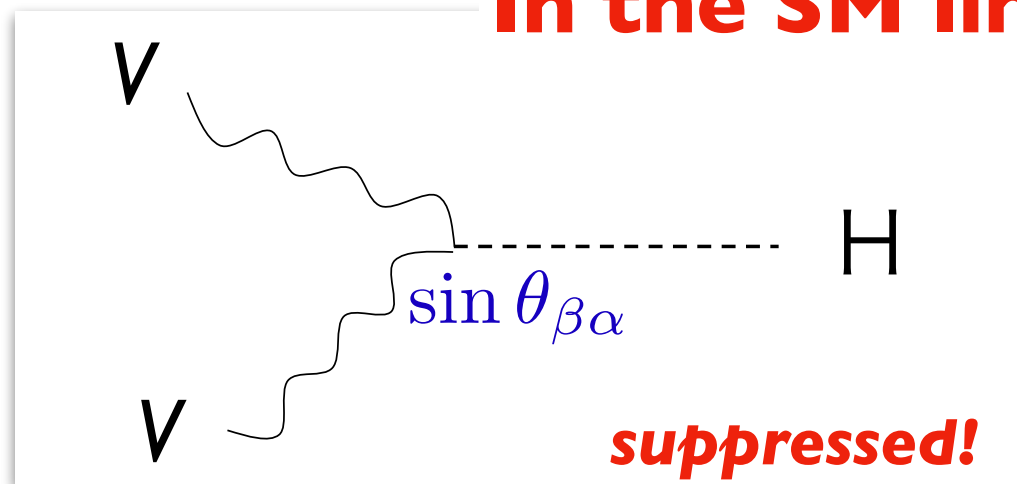


*The scalars decays to heavy fermions and di-bosons*

# We could see the scalars directly at the LHC

*The scalars decays to heavy fermions and di-bosons*

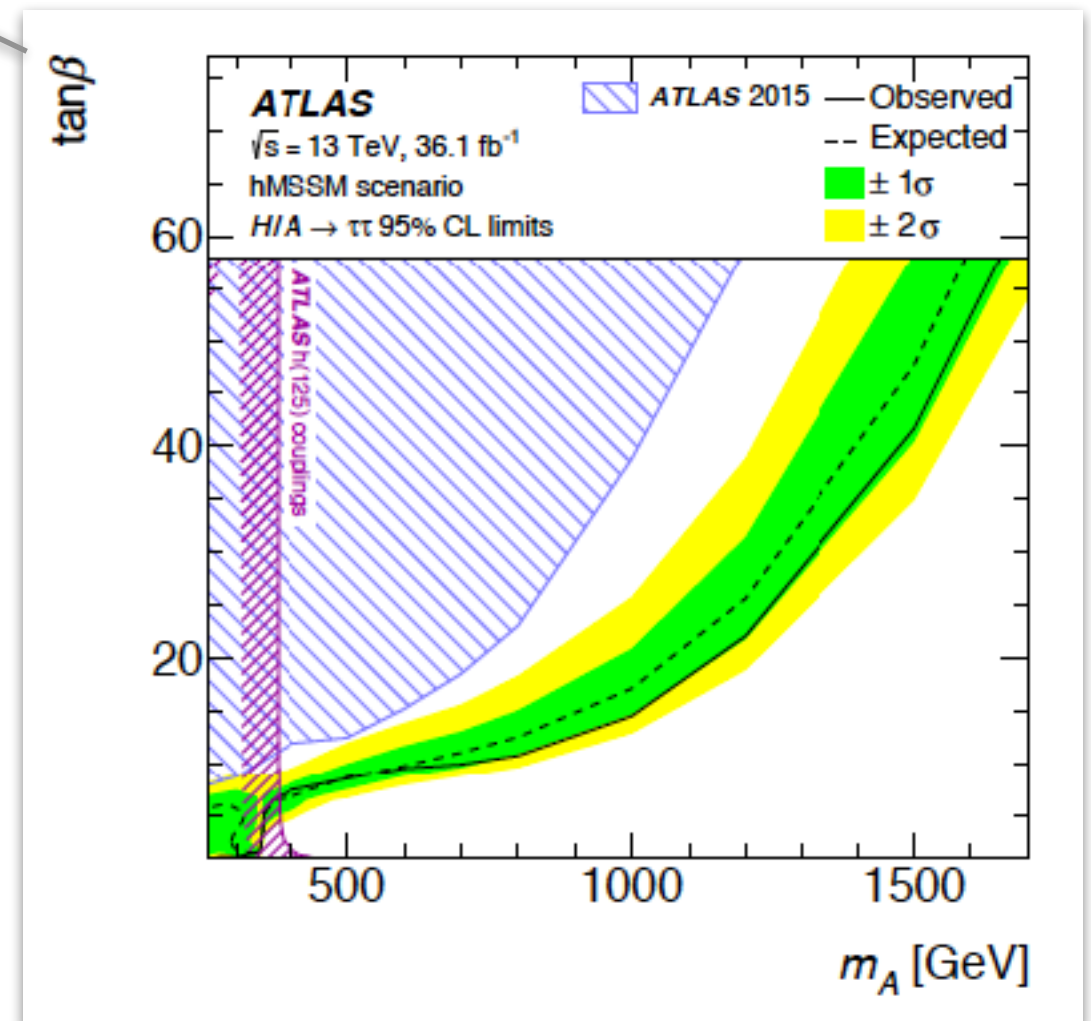
**In the SM limit,**



# Actually, many modes have been already studied.

Label	Channel	Experiment	Mass range [GeV]	$\mathcal{L}$ [fb <sup>-1</sup> ]	Channel	Experiment	Mass range [TeV]	$\mathcal{L}$ [fb <sup>-1</sup> ]
$C_{13}^{bb}$	$pp \rightarrow H/A \rightarrow bb$	CMS [75]	[0.55;1.2]	2.69	$pp \rightarrow H^\pm \rightarrow \tau^\pm \nu$	ATLAS [100] CMS [101]	[0.2;2] [0.18;3]	14.7 12.9
$A_{13}^{\tau\tau}$ $C_{13}^{\tau\tau}$	$gg \rightarrow H/A \rightarrow \tau\tau$	ATLAS [76] CMS [77]	[0.2;2.25] [0.09;3.2]	36.1 12.9	$pp \rightarrow H^\pm \rightarrow t\bar{b}$	ATLAS [102] ATLAS [103]	[0.3;1] [0.2;0.3] $\cup$ [1;2]	13.2 13.2
$A_{13b}^{\tau\tau}$ $C_{13b}^{\tau\tau}$	$bb \rightarrow H/A \rightarrow \tau\tau$	ATLAS [76] CMS [77]	[0.2;2.25] [0.09;3.2]	36.1 12.9				
$A_{13}^{\gamma\gamma}$ $C_{13}^{\gamma\gamma}$	$pp \rightarrow H/A \rightarrow \gamma\gamma$	ATLAS [78] CMS [79]	[0.2;2.7] [0.5;4]	36.7 35.9				
$A_{13}^{Z\gamma}$ $C_{13}^{Z\gamma}$	$gg \rightarrow H/A \rightarrow Z\gamma [\rightarrow (\ell\ell)\gamma]$	ATLAS [45] CMS [80]	[0.25;2.4] [0.35;4]	36.1 35.9				
$A_{13}^{2\ell 2\ell}$ $A_{13V}^{2\ell 2\ell}$	$gg \rightarrow H \rightarrow ZZ [\rightarrow (\ell\ell)(\ell\ell, \nu\nu)]$	ATLAS [81]	[0.2;1.2]	36.1				
$C_{13}^{2\ell 2\nu}$ $C_{13V}^{2\ell 2\nu}$	$VV \rightarrow H \rightarrow ZZ [\rightarrow (\ell\ell)(\ell\ell, \nu\nu)]$	ATLAS [81]	[0.2;1.2]	36.1				
$C_{13}^{2\ell 2\nu}$ $C_{13V}^{2\ell 2\nu}$	$pp \rightarrow H \rightarrow ZZ [\rightarrow (\ell\ell)(\nu\nu)]$	CMS [82]	[0.6;2.5]	35.9				
$C_{13}^{2\ell 2\nu}$ $C_{13V}^{2\ell 2\nu}$	$gg \rightarrow H \rightarrow ZZ [\rightarrow (\ell\ell)(\nu\nu)]$	CMS [83]	[0.2;0.6]	2.3				
$C_{13V}^{2\ell 2\nu}$ $C_{13V}^{4\ell}$	$VV \rightarrow H \rightarrow ZZ [\rightarrow (\ell\ell)(\nu\nu)]$	CMS [83]	[0.2;0.6]	2.3				
$C_{13V}^{4\ell}$ $C_{13}^{2\ell 2q}$	$(VV + VH) \rightarrow H \rightarrow ZZ \rightarrow (\ell\ell)(\ell\ell)$	CMS [84]	[0.13;2.53]	12.9				
$A_{13}^{2\ell 2q}$ $A_{13V}^{2\ell 2q}$	$pp \rightarrow H \rightarrow ZZ [\rightarrow (\ell\ell)(qq)]$	CMS [85]	[0.5;2]	12.9				
$A_{13}^{2\ell 2q}$ $A_{13V}^{2\ell 2q}$	$gg \rightarrow H \rightarrow ZZ [\rightarrow (\ell\ell, \nu\nu)(qq)]$	ATLAS [86]	[0.3;3]	36.1				
$A_{13}^{2\ell 2q}$ $A_{13V}^{2\ell 2q}$	$VV \rightarrow H \rightarrow ZZ [\rightarrow (\ell\ell, \nu\nu)(qq)]$	ATLAS [86]	[0.3;3]	36.1				
$A_{13}^{2(\ell\nu)}$ $A_{13V}^{2(\ell\nu)}$	$gg \rightarrow H \rightarrow WW [\rightarrow (e\nu)(\mu\nu)]$	ATLAS [87]	[0.25;4]	36.1				
$C_{13}^{2(\ell\nu)}$ $A_{13}^{4\nu 2q}$	$VV \rightarrow H \rightarrow WW [\rightarrow (e\nu)(\mu\nu)]$	ATLAS [87]	[0.25;3]	36.1				
$A_{13}^{4\nu 2q}$ $A_{13V}^{4\nu 2q}$	$(gg + VV) \rightarrow H \rightarrow WW \rightarrow (\ell\nu)(\ell\nu)$	CMS [88]	[0.2;1]	2.3				
$A_{13}^{4\nu 2q}$ $A_{13V}^{4\nu 2q}$	$gg \rightarrow H \rightarrow WW [\rightarrow (\ell\nu)(qq)]$	ATLAS [89]	[0.3;3]	36.1				
$A_{13}^{4\nu 2q}$ $A_{13V}^{4\nu 2q}$	$VV \rightarrow H \rightarrow WW [\rightarrow (\ell\nu)(qq)]$	ATLAS [89]	[0.3;3]	36.1				
$A_{13}^{4q}$	$pp \rightarrow H \rightarrow VV [\rightarrow (qq)(qq)]$	ATLAS [90]	[1.2;3]	36.7				
$A_{13}^{4b}$ $C_{13}^{4b}$	$pp \rightarrow H \rightarrow hh \rightarrow (bb)(bb)$	ATLAS [91] CMS [92]	[0.3;3] [0.26;1.2]	13.3 35.9				
$C_{13g}^{4b}$ $A_{13}^{2\gamma 2h}$	$gg \rightarrow H \rightarrow hh \rightarrow (bb)(bb)$	CMS [93]	[1.2;3]	35.9				
$C_{13}^{2\gamma 2h}$ $C_{13}^{2\gamma 2h}$	$pp \rightarrow H \rightarrow hh [\rightarrow (\gamma\gamma)(bb)]$	ATLAS [94] CMS [95]	[0.275;0.4] [0.25;0.9]	3.2 35.9				
$C_{13}^{2b 2\tau}$ $C_{13}^{2b 2V}$	$pp \rightarrow H \rightarrow hh \rightarrow (bb)(\tau\tau)$	CMS [96]	[0.25;0.9]	35.9				
$A_{13}^{2\gamma 2W}$ $A_{13}^{4bZ}$	$pp \rightarrow H \rightarrow hh \rightarrow (bb)(VV \rightarrow \ell\nu\ell\nu)$	CMS [97]	[0.26;0.9]	36				
$A_{13}^{4bZ}$ $A_{13b}^{4bZ}$	$gg \rightarrow H \rightarrow hh [\rightarrow (\gamma\gamma)(WW)]$	ATLAS [98]	[0.25;0.5]	13.3				
$A_{13}^{4bZ}$ $A_{13b}^{4bZ}$	$gg \rightarrow A \rightarrow hZ \rightarrow (bb)Z$	ATLAS [99]	[0.2;2]	36.1				
$A_{13b}^{4bZ}$ $A_{13b}^{4bZ}$	$bb \rightarrow A \rightarrow hZ \rightarrow (bb)Z$	ATLAS [99]	[0.2;2]	36.1				

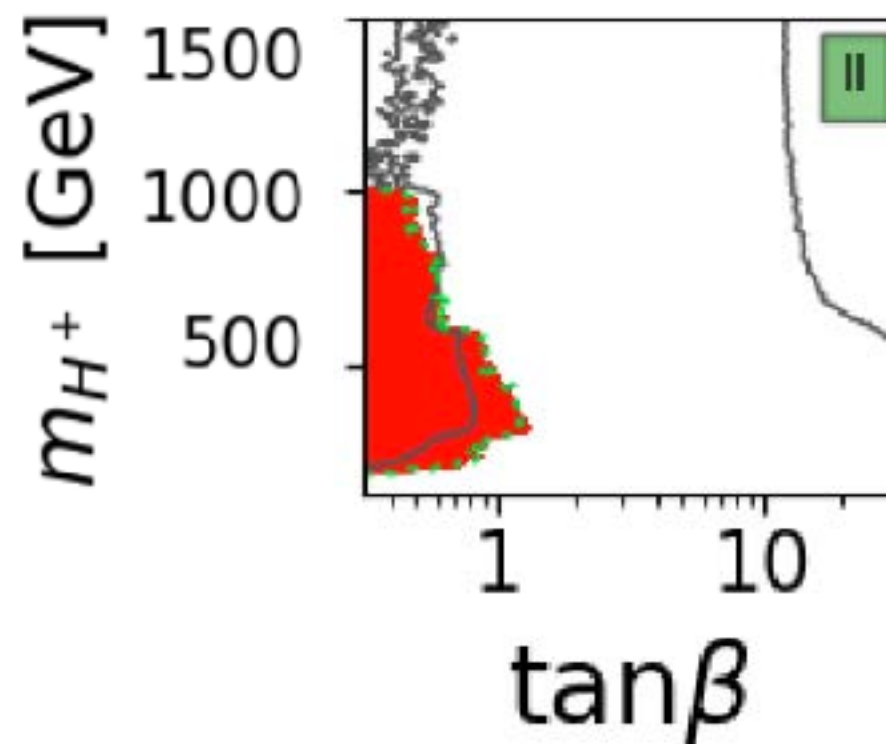
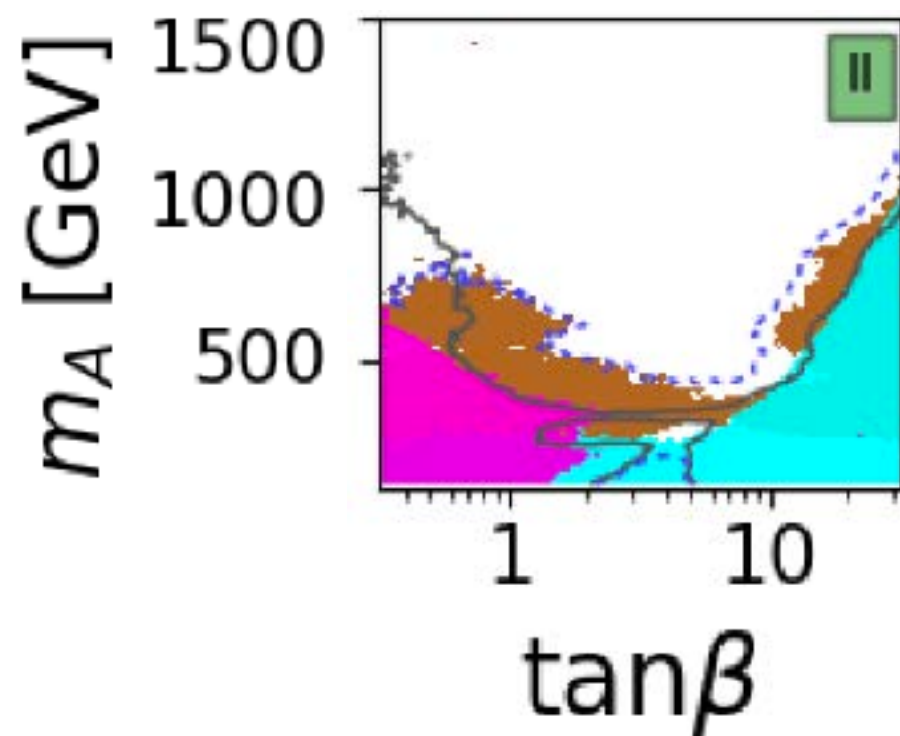
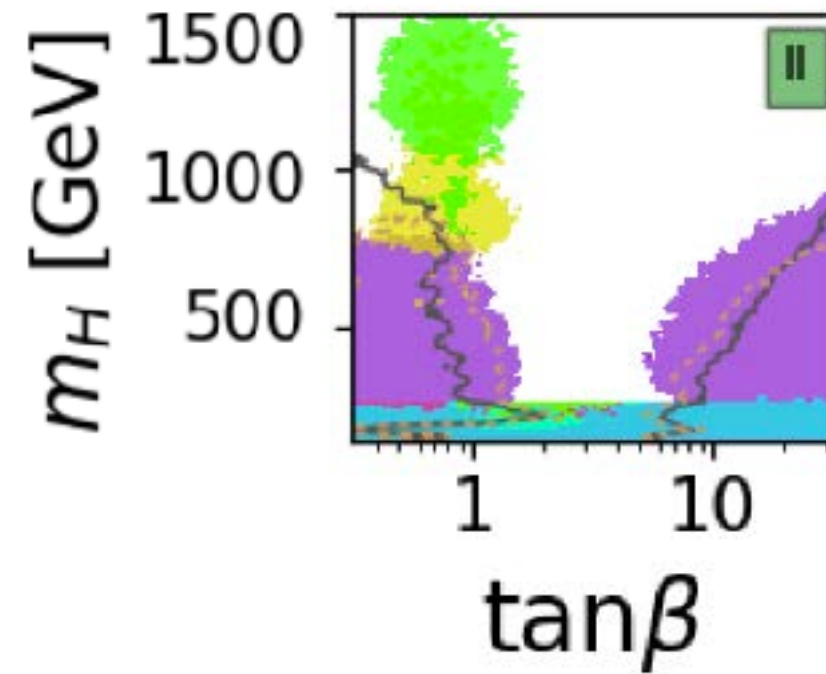
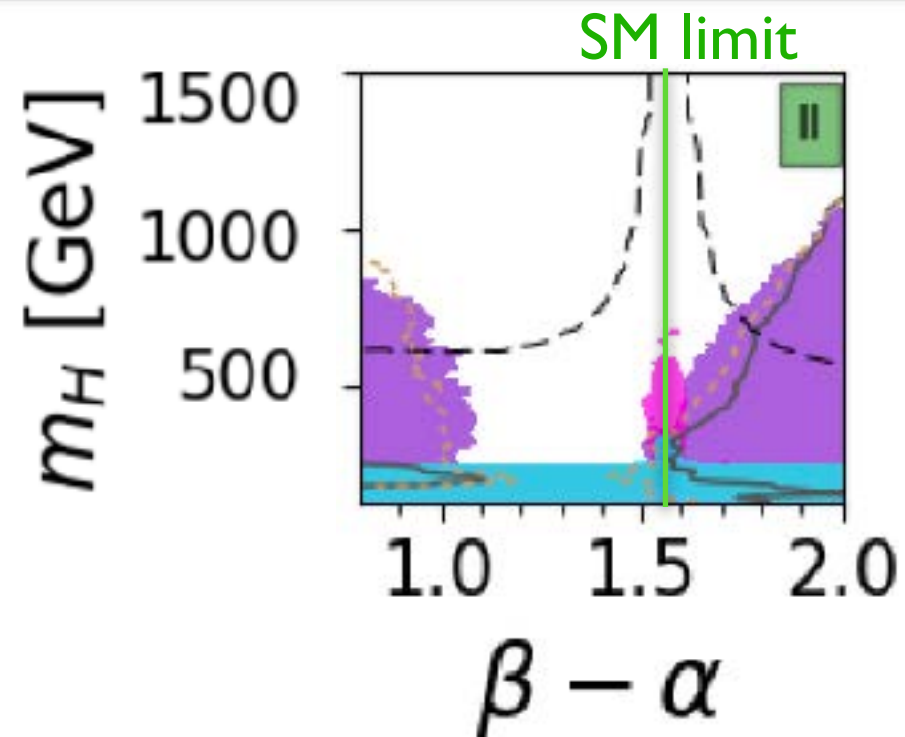
(Chowdhury, Eberhardt, 1711.02095)





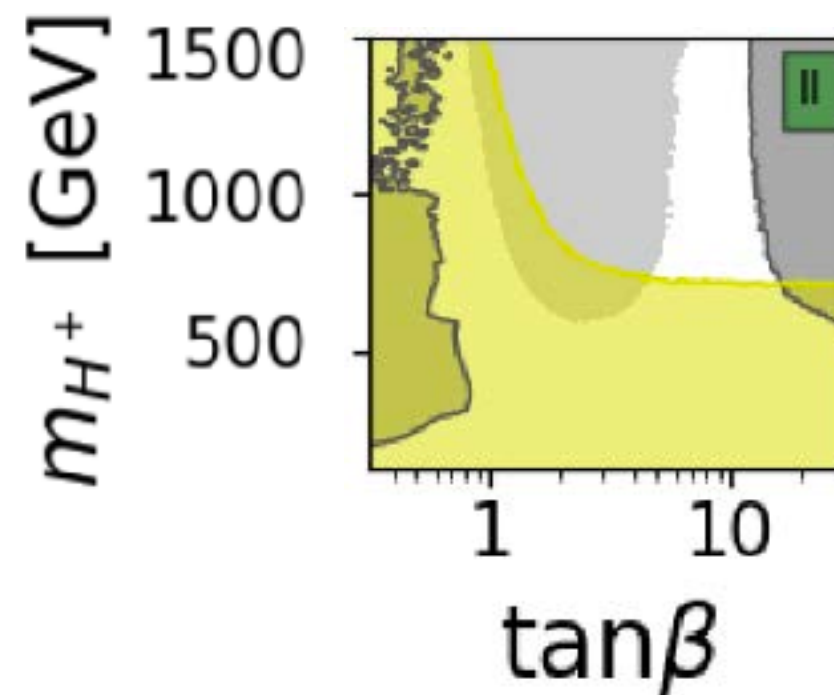
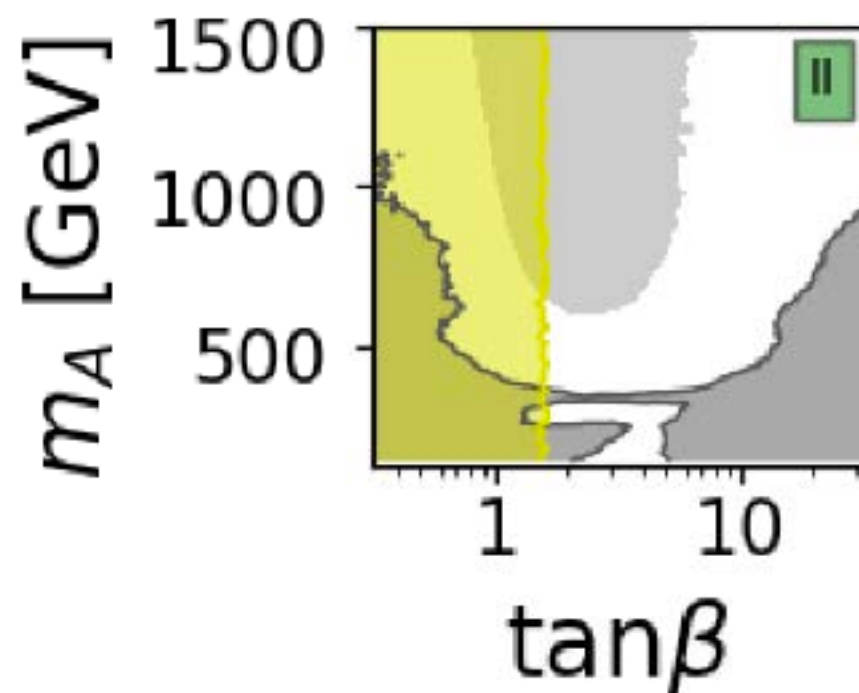
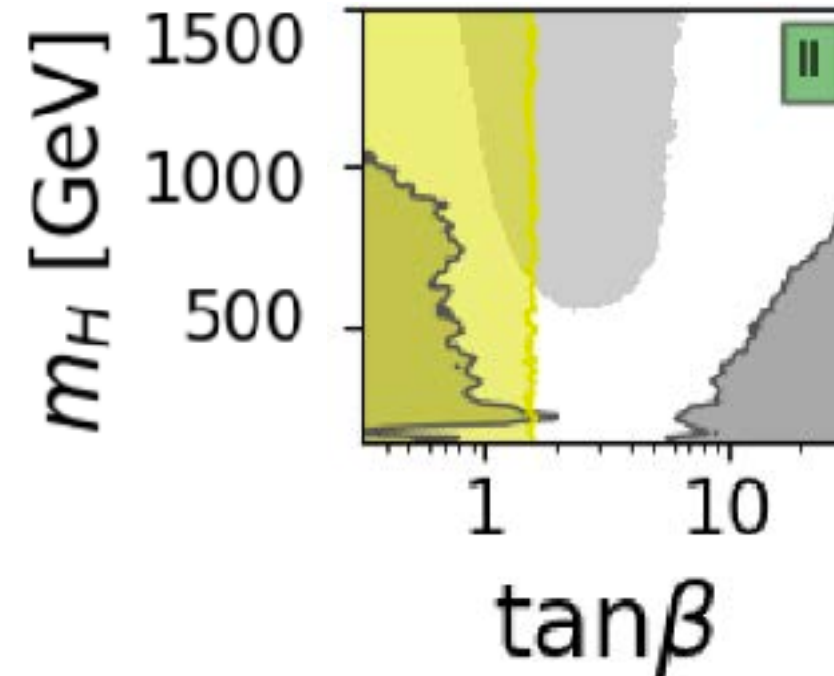
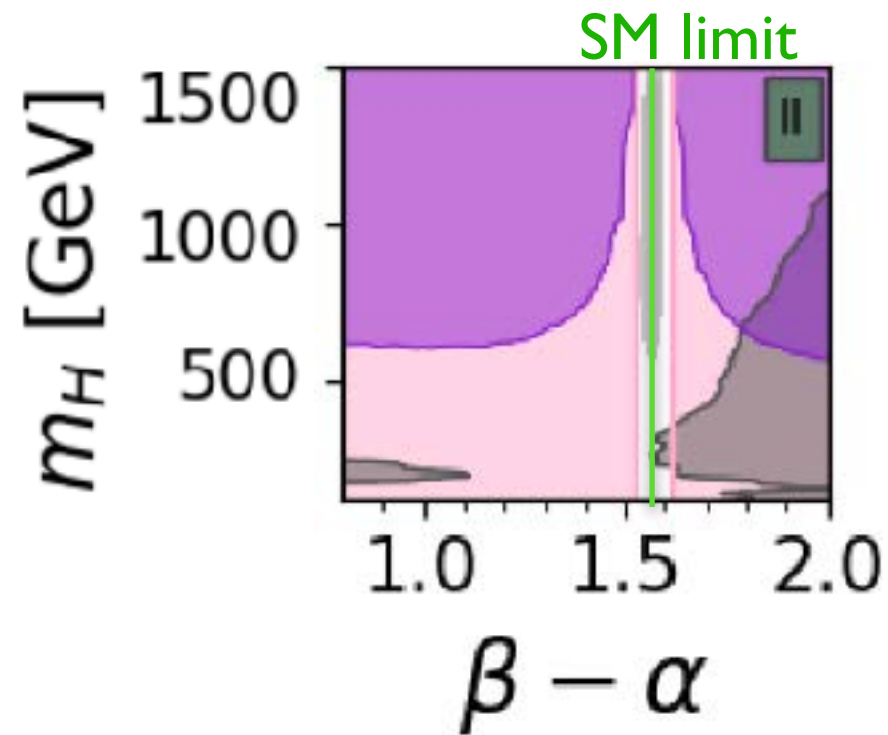
# Summary plots

(Chowdhury, Eberhardt, 1711.02095)



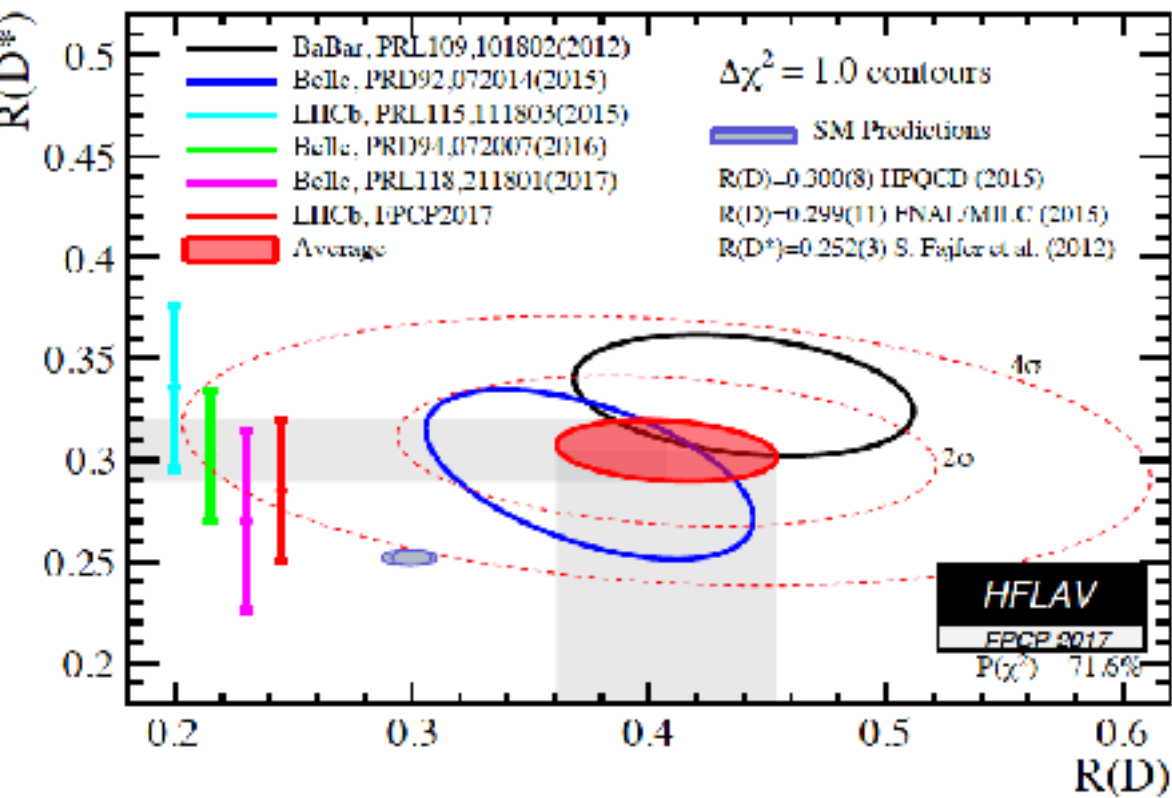
# Compare with others

(Chowdhury, Eberhardt, 1711.02095)



# Impact of the 2HDM on the excesses in the flavor physics

# The excess in Lepton Flavor Universality (LFU) in $B \rightarrow D^{(*)} \ell \nu$

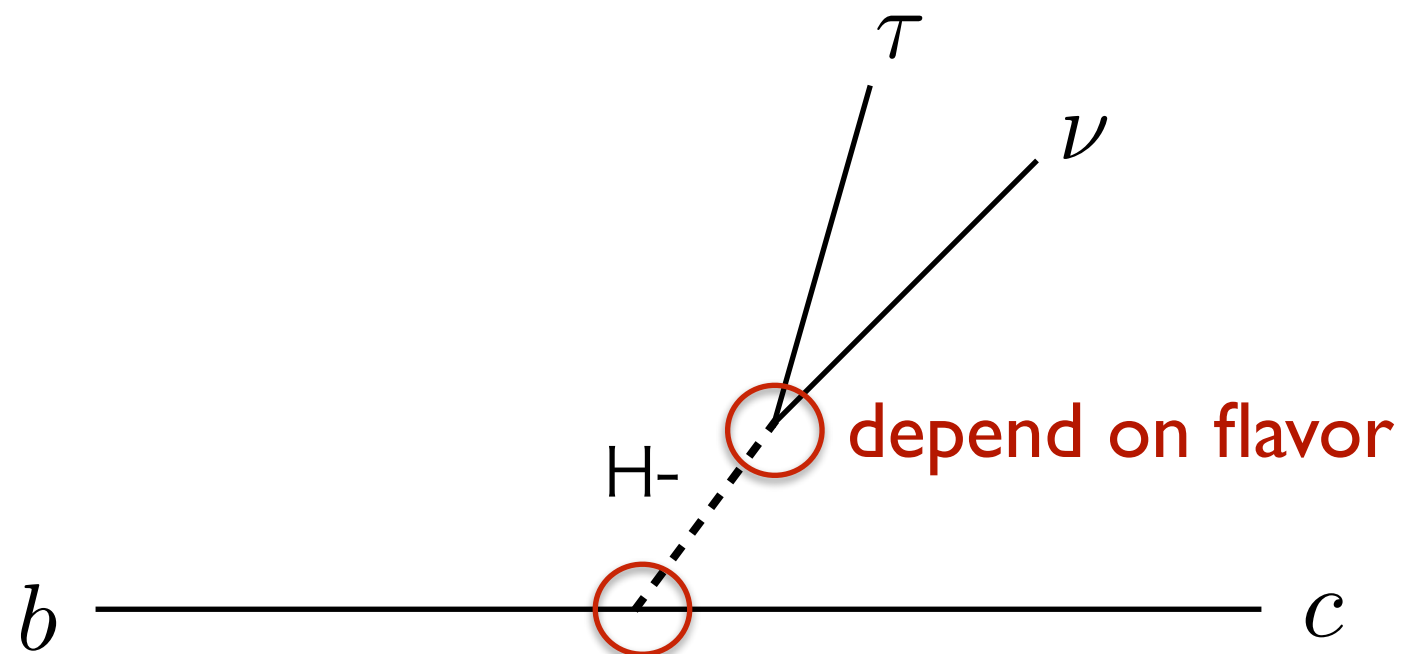


lepton universality of  $B \rightarrow D^{(*)} \tau \nu$

$$R(D^{(*)}) \equiv \frac{\Gamma(B \rightarrow D^{(*)} \tau \nu)}{\Gamma(B \rightarrow D^{(*)} \ell \nu)}$$

where  $\ell = e, \mu$

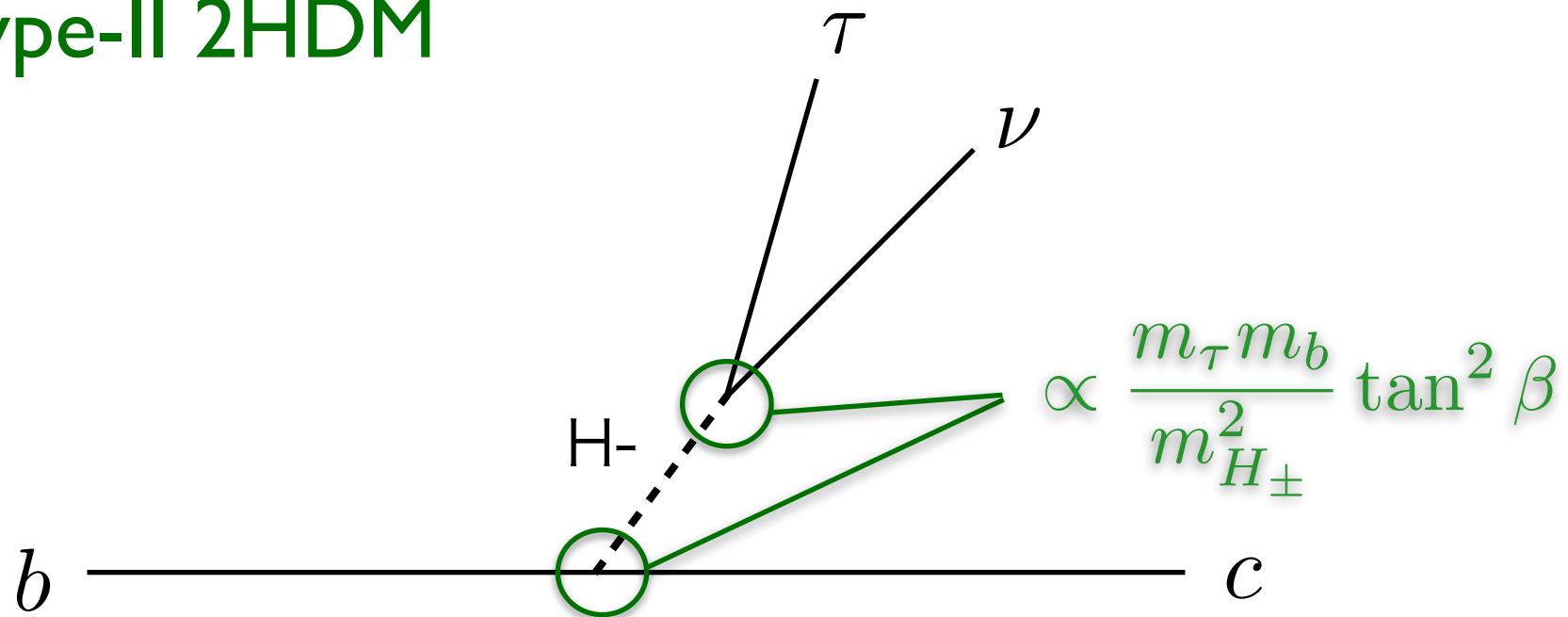
Charged Higgs is a very good candidate!





# The type-II 2HDM is strongly constrained.

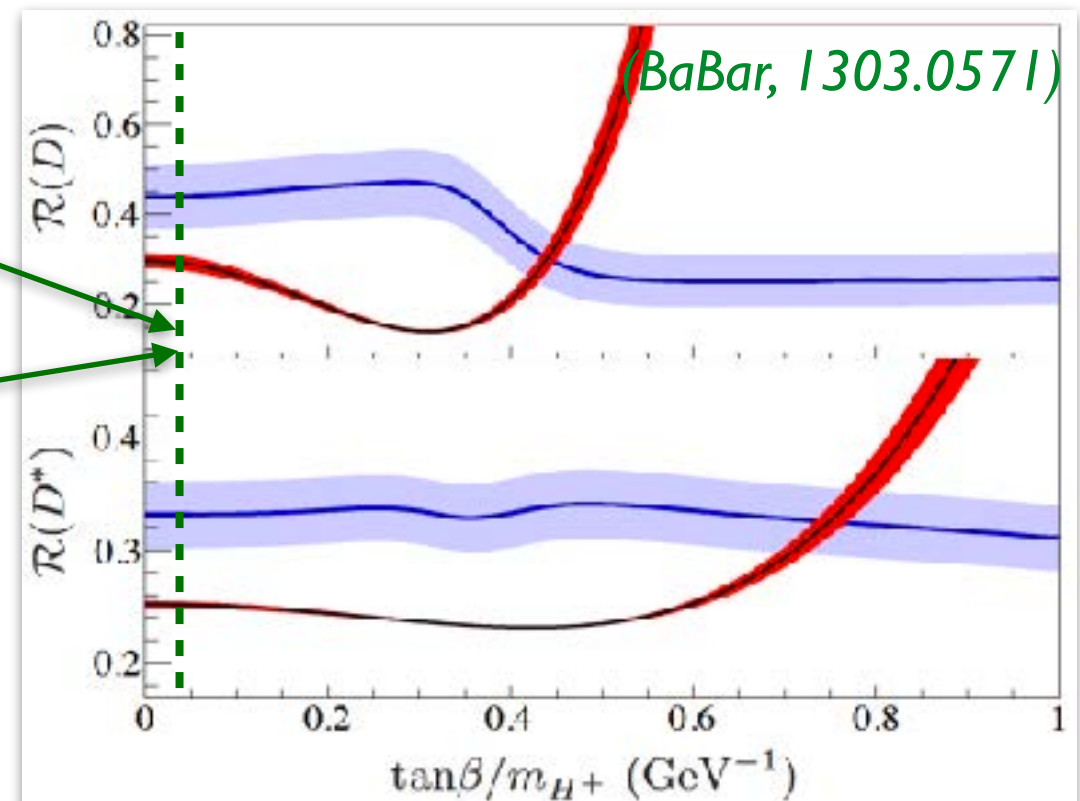
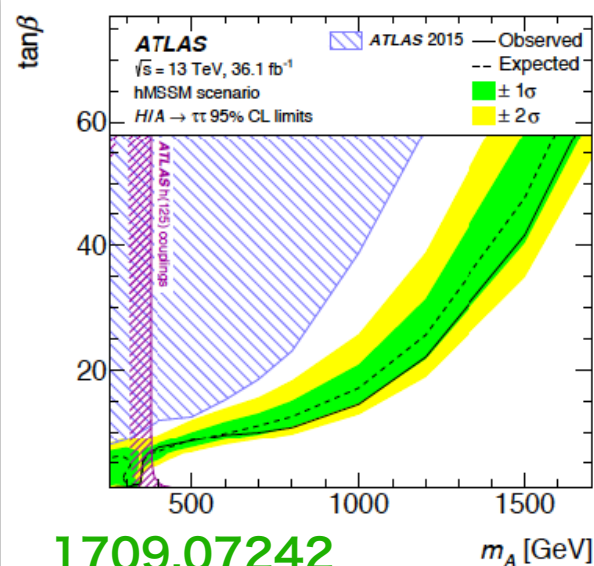
In type-II 2HDM



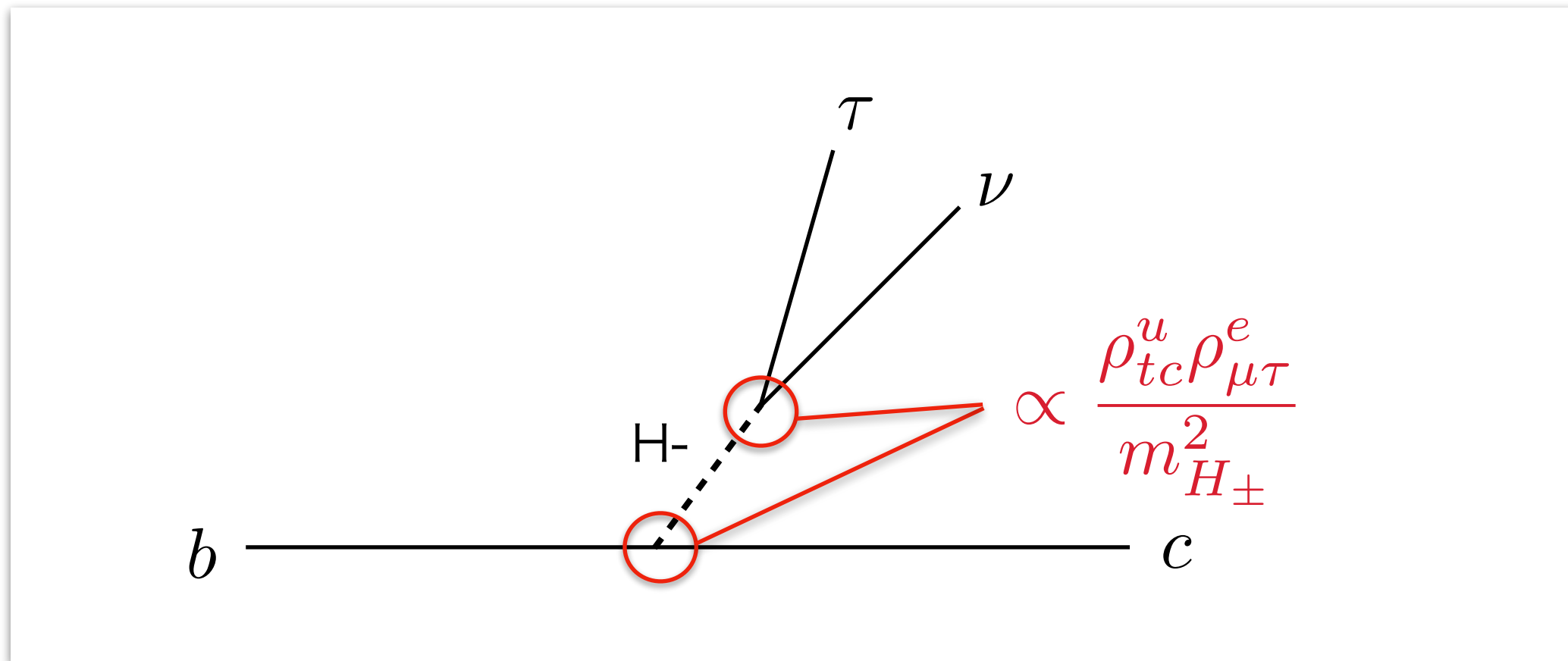
bound is from  $b \rightarrow s \gamma$ :

$$m_{H^\pm} > 580 \text{ GeV}$$

Misiak, Steinhauser, 1702.04571



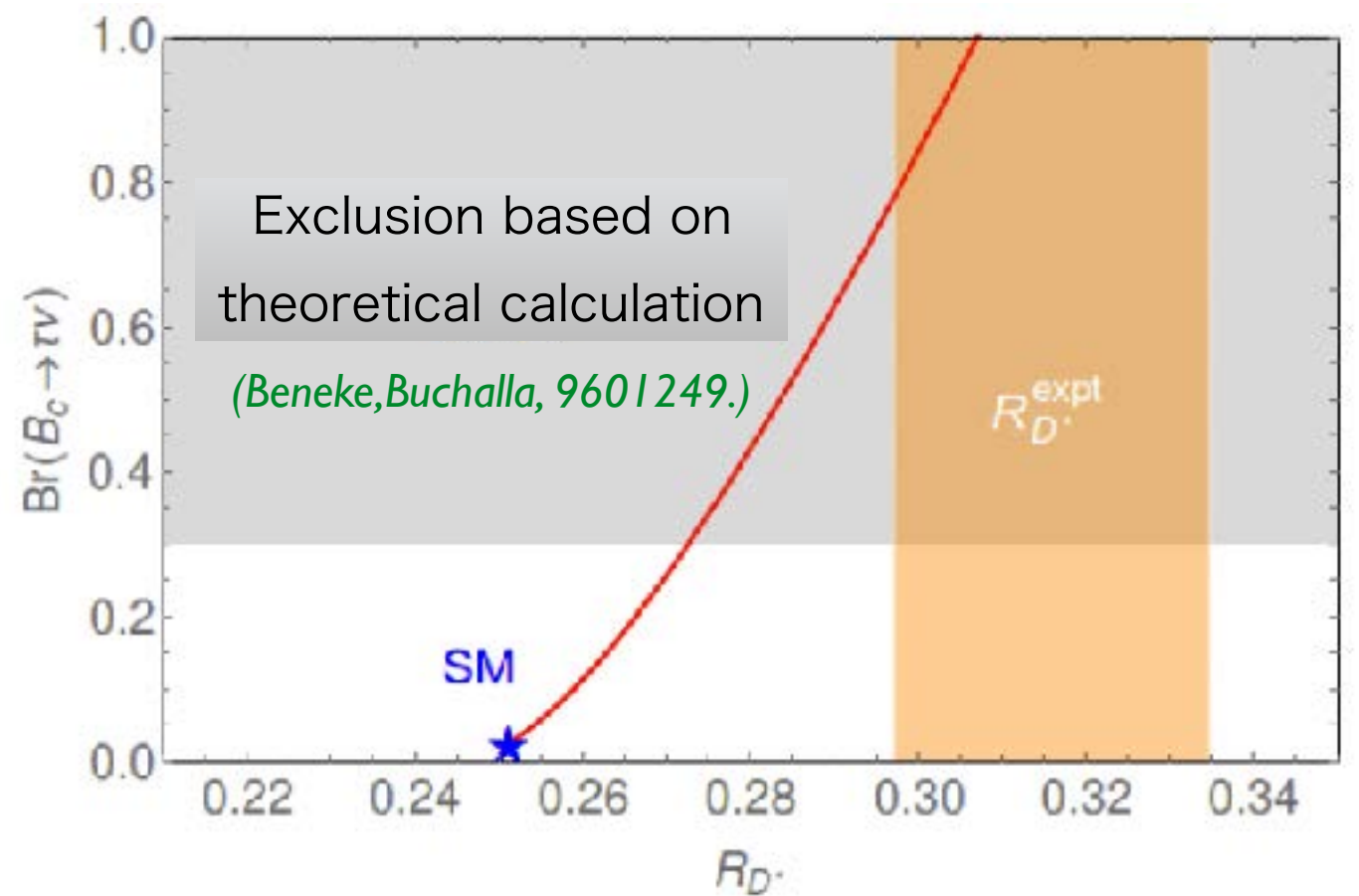
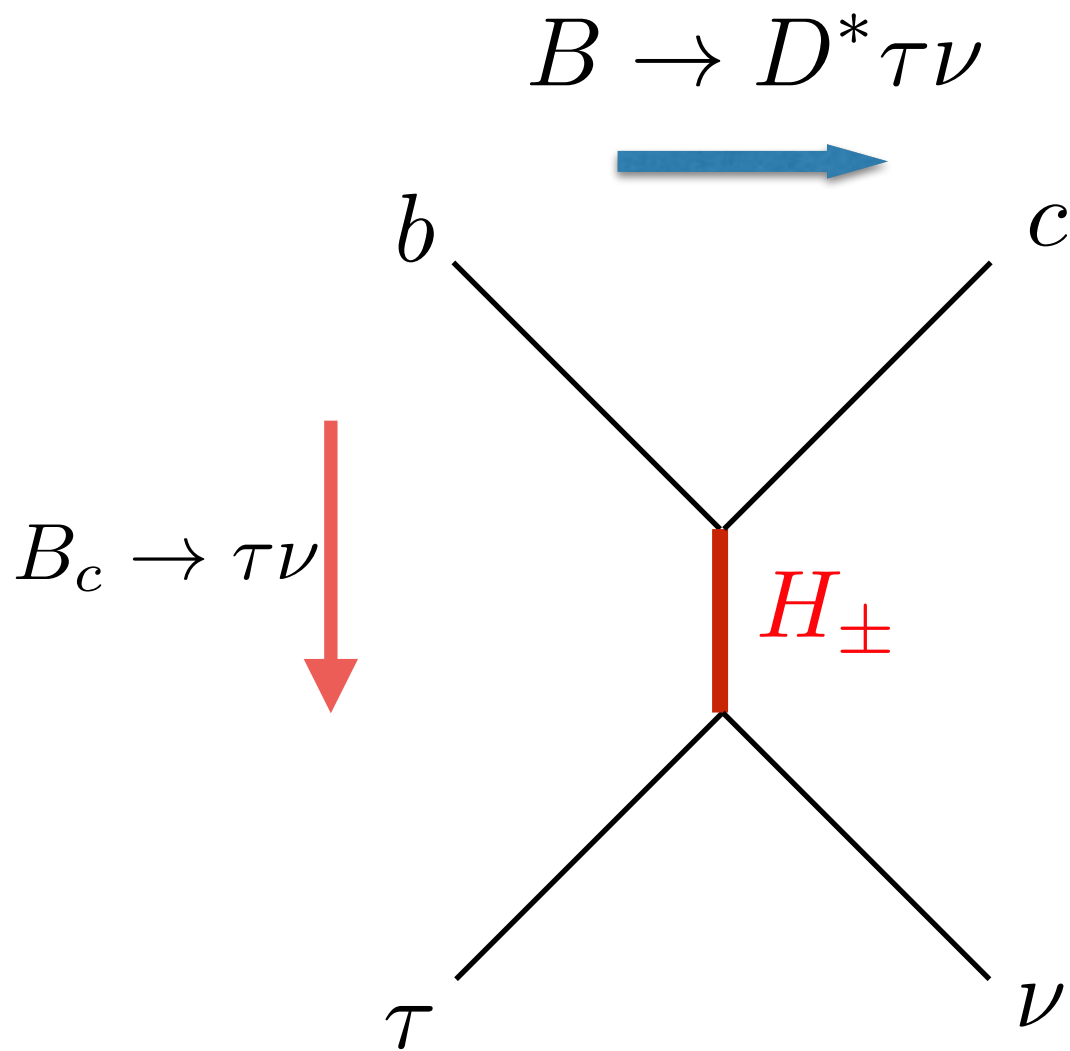
We have to think about the 2HDM with tree-level FCNCs,



although we have to tune many para. to avoid the flavor constraints.

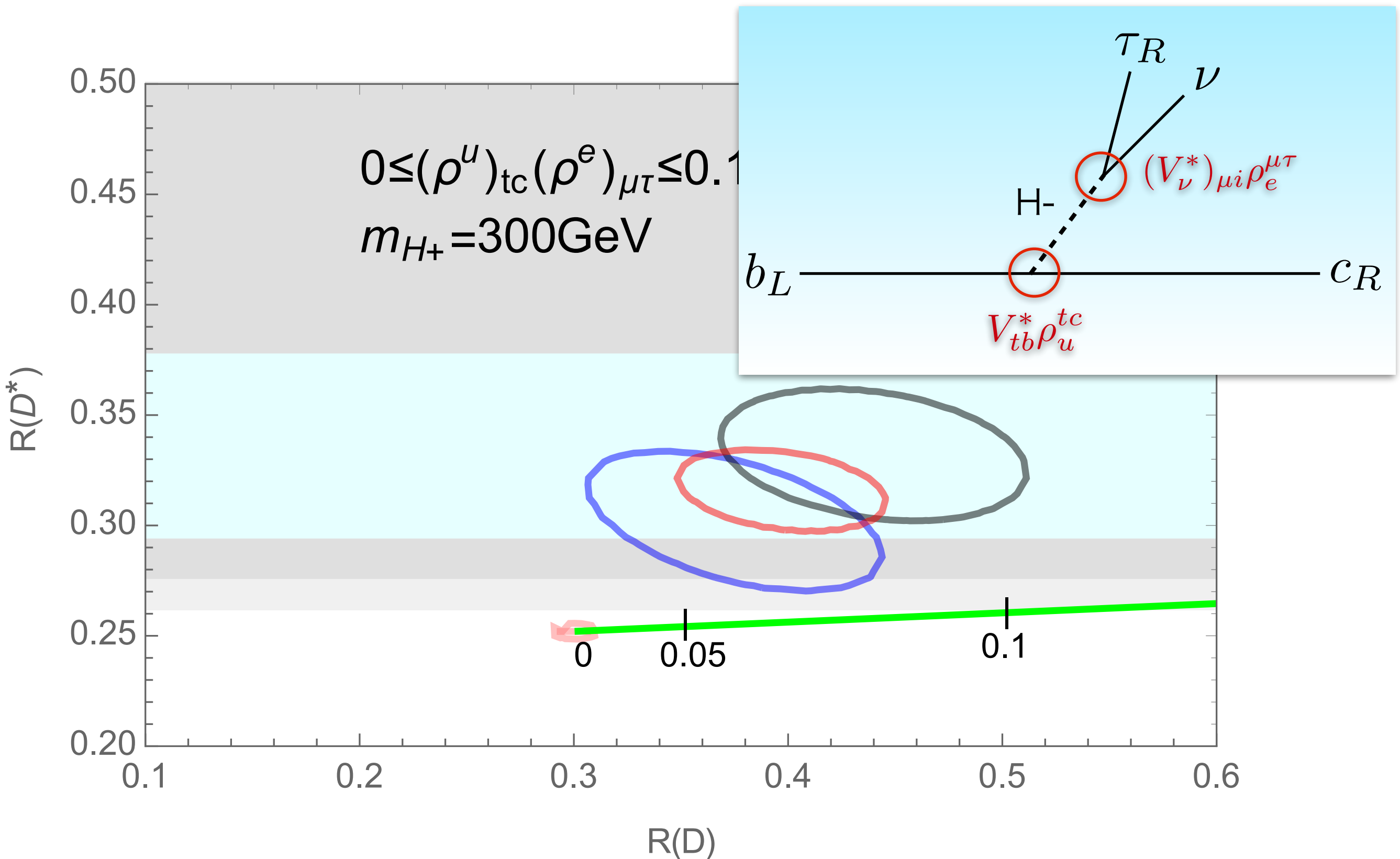
# Bc decay limits the R(D\*) in 2HDM.

(Alonso,Grinstein,et al., 1611.06676;Akeroyd,Chen,1708.04072)

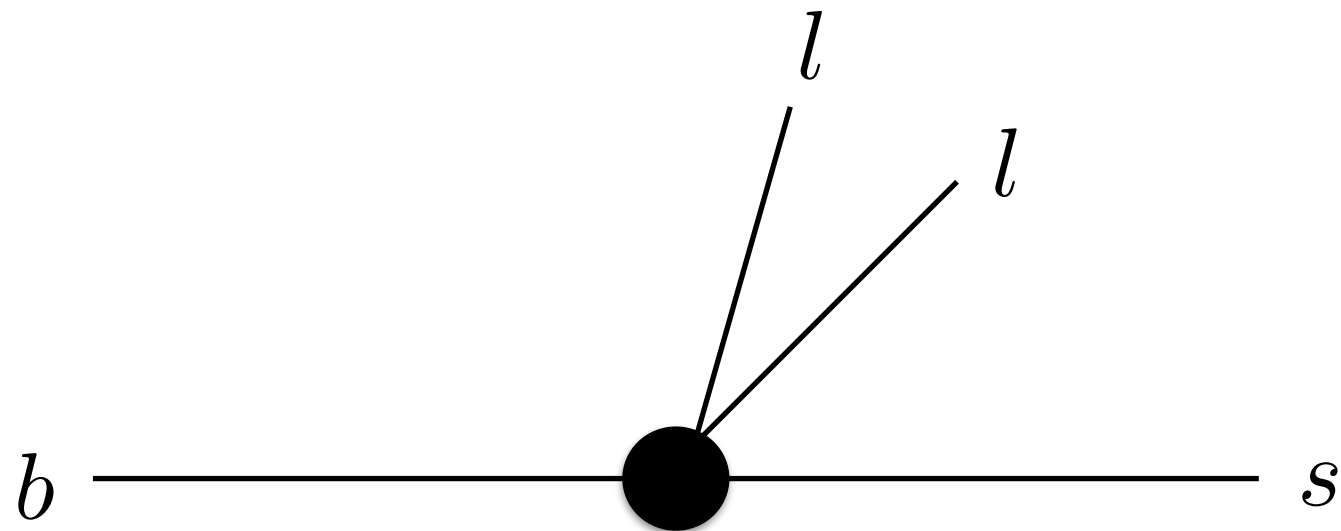


$$Br(B_c \rightarrow \tau \nu) = \left| 1 + \delta_{LQ} + \frac{m_{B_c}^2}{m_{\tau}(m_b + m_c)} \delta_{H_{\pm}} \right|^2 Br(B_c \rightarrow \tau \nu)_{SM}$$

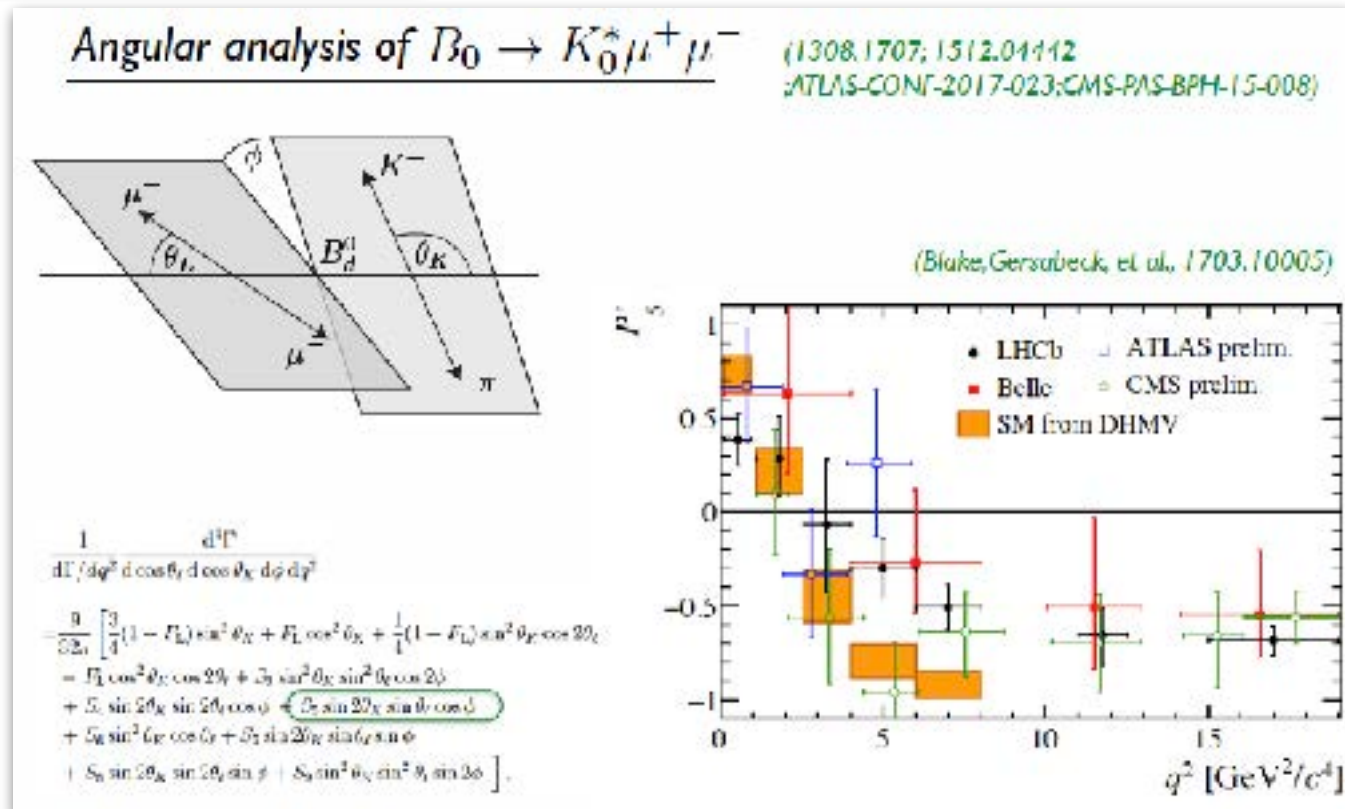
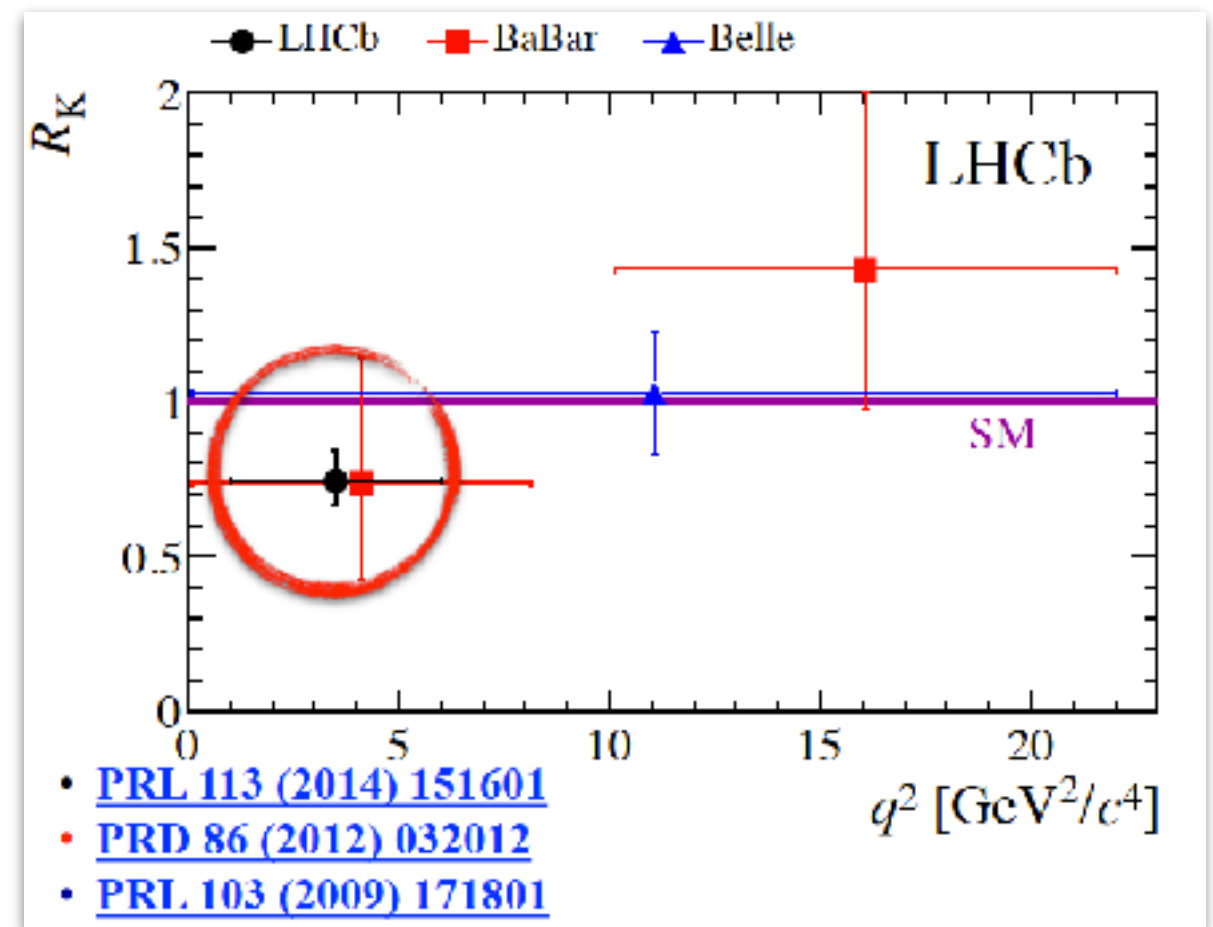
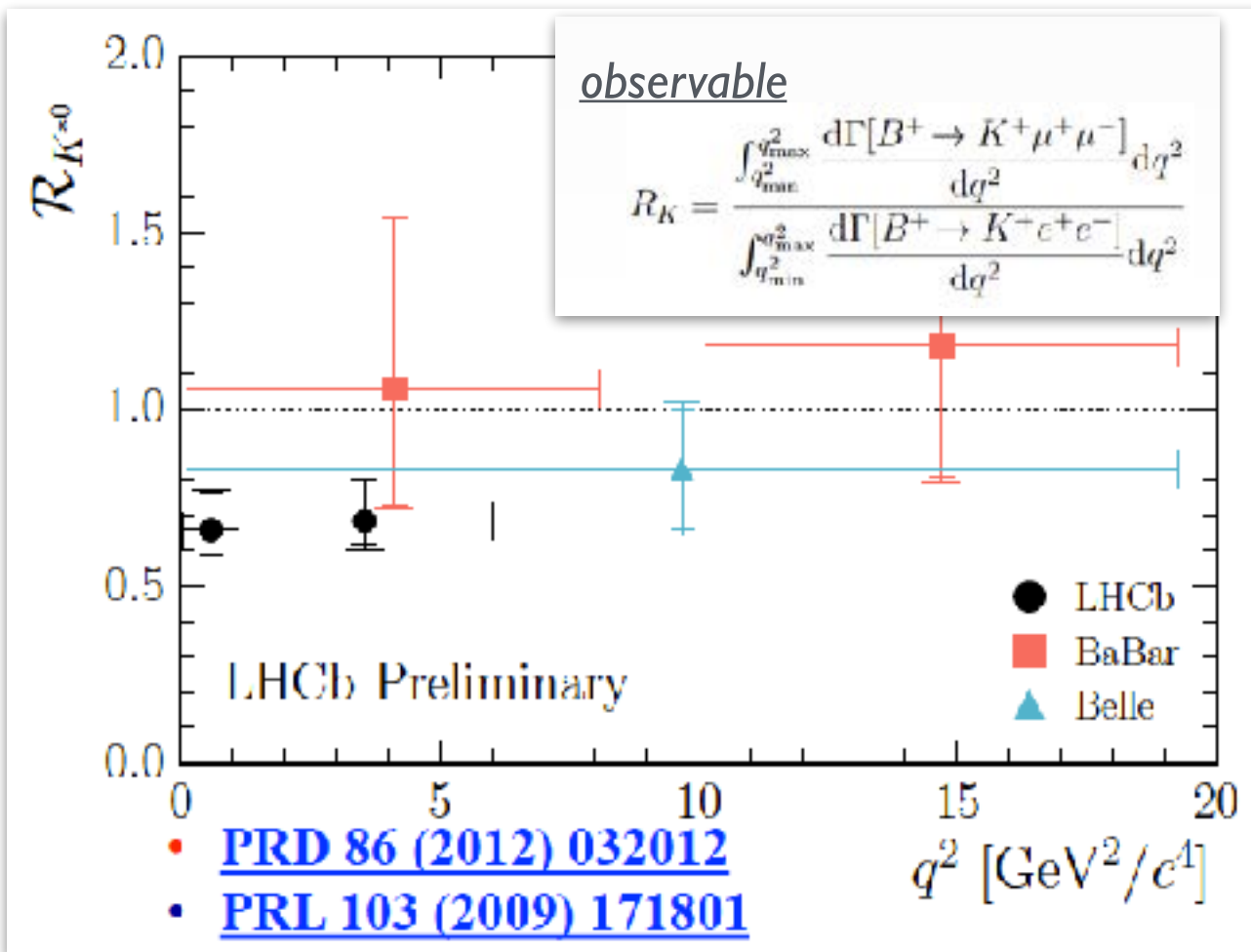
# R(D) and R(D\*) in a 2HDM



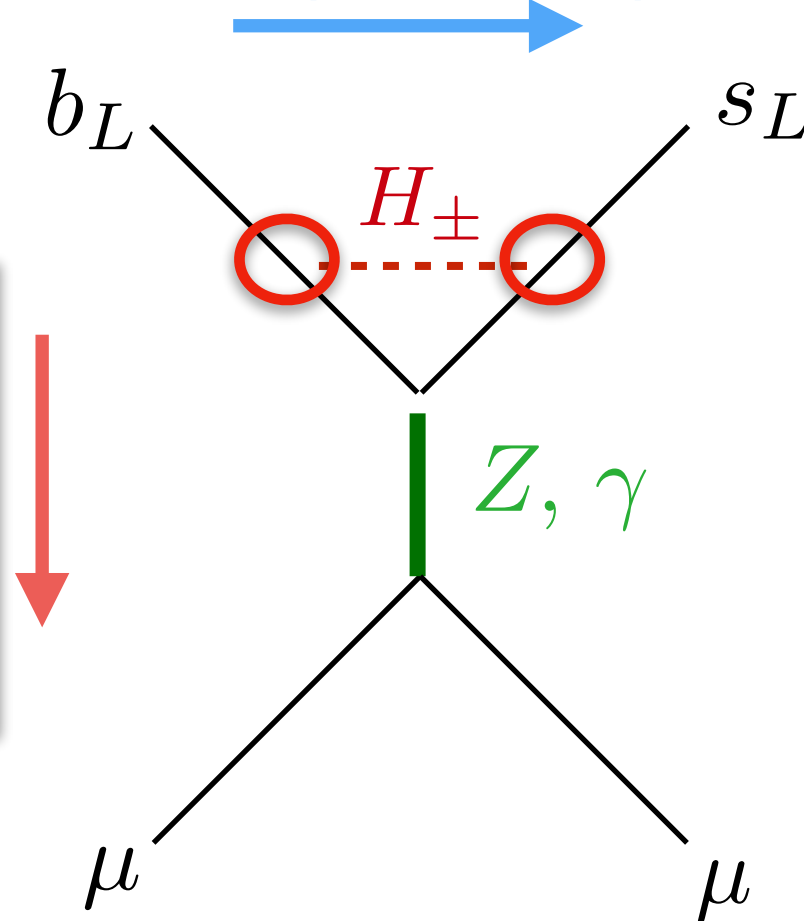
There are several excesses in  $B \rightarrow K(*)\Pi$



# There are several excesses in $B \rightarrow K(*)\ell\ell$



Semileptonic B decays



$$B_s \rightarrow \mu\mu$$

$$BR(B_s \rightarrow \mu^+\mu^-)^{\text{exp}} = (2.8^{+0.7}_{-0.6}) \times 10^{-9}$$

$$BR(B_s \rightarrow \mu^+\mu^-)^{\text{SM}} = (3.65 \pm 0.23) \times 10^{-9}$$

Generated ope. are  $C_9$  and  $C_{10}$  :

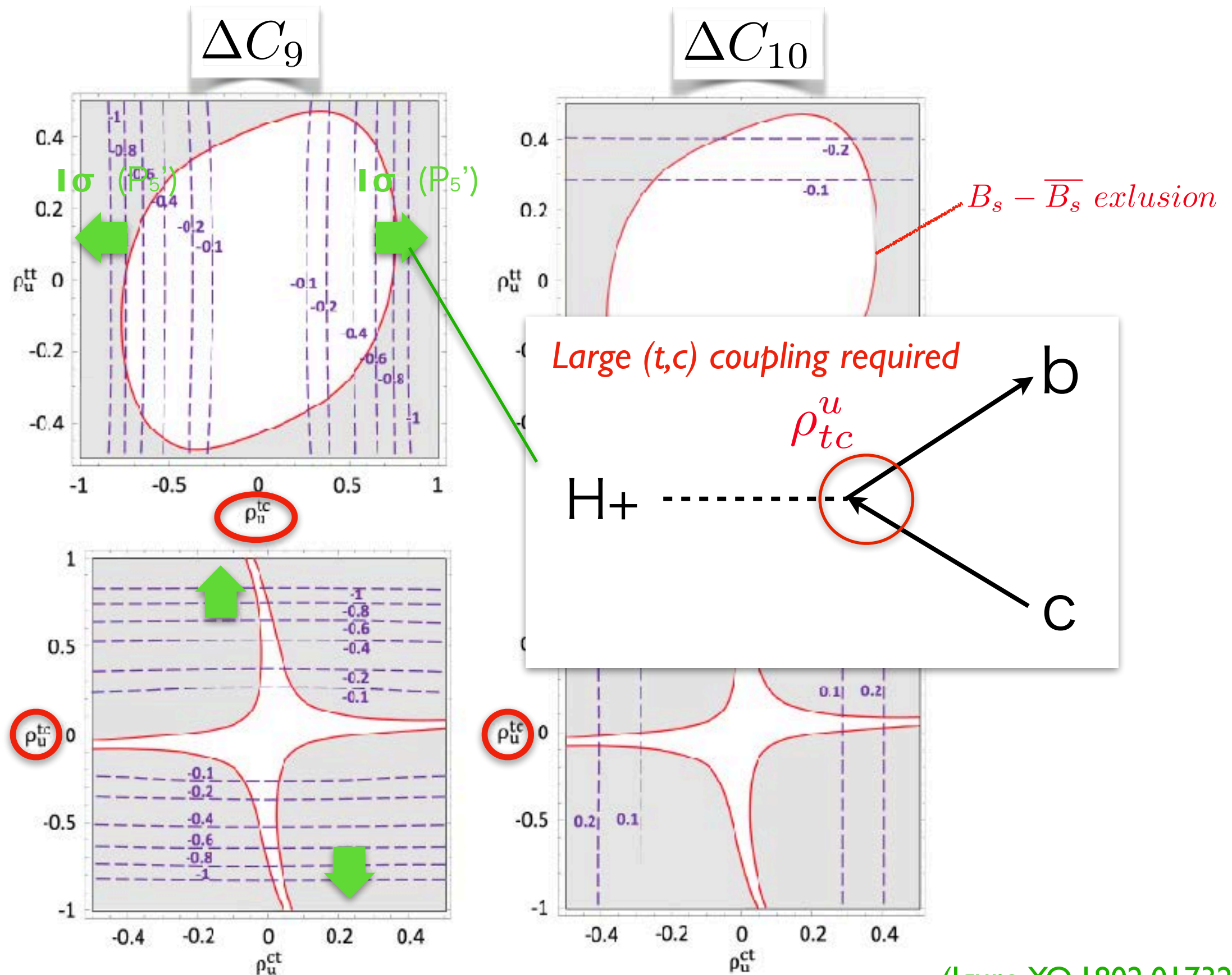
$$\mathcal{H}_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \frac{e^2}{16\pi^2} \{C_9(\bar{s}_L \gamma_\mu b_L)(\bar{\mu} \gamma^\mu \mu) + C_{10}(\bar{s}_L \gamma_\mu b_L)(\bar{\mu} \gamma^\mu \gamma_5 \mu) + h.c.\}$$

$C_{10}$  contributes to  $B_s \rightarrow \mu\mu$ :

$$\frac{BR(B_s \rightarrow \mu\mu)}{BR(B_s \rightarrow \mu\mu)_{SM}} = |1 - 0.24 C_{10}^\mu|^2$$



# Results in 2HDM with only $\rho u$ @ $m_{H_{\pm}} = 200$ GeV





What can we expect from  
the underlying theories?

In MSSM, for instance,

EW scale relates to SUSY breaking para.:

CP-odd scalar mass

$$m_A^2 = \frac{1}{\cos 2\beta} (m_{H_d}^2 - m_{H_u}^2) + M_Z^2$$

SUSY breaking para.

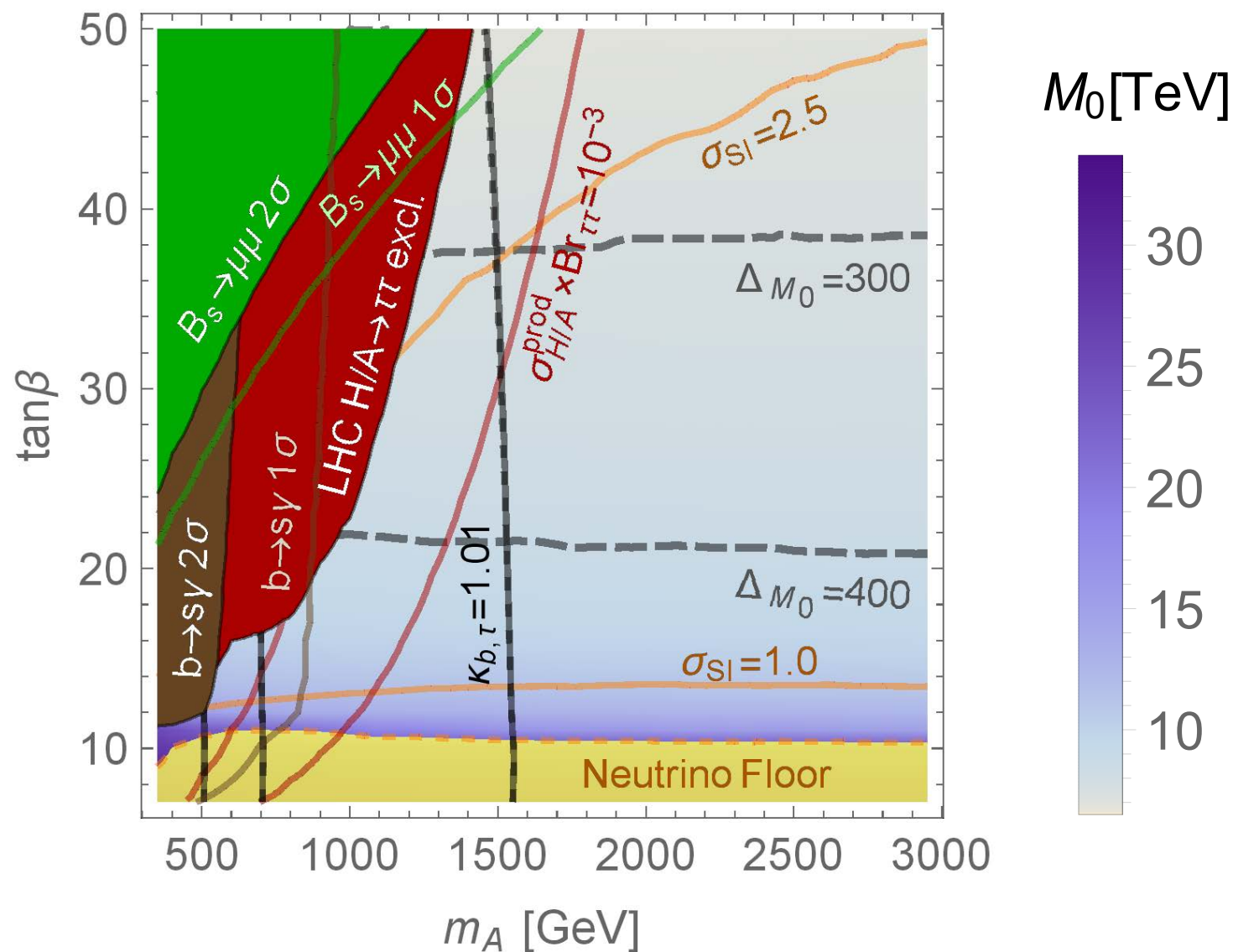
(Higgsino)

$$\mu^2 = \frac{m_{H_u}^2 \tan^2 \beta - m_{H_d}^2}{1 - \tan^2 \beta} - \frac{1}{2} M_Z^2$$

*SUSY scale. have to be much higher than we expected,  
so that the extra Higgs mass is expected to be also high.*

One scenario to predict Type-II 2HDM is  
*mirage mediation (moduli-anomaly mixture)*

Type-II 2HDM with Higgsino DM



(Kawamura,YO, arXiv:1710.03412)

*BSMs that unify the SM Yukawa also lead 2HDM at the low energy*

*(Iguro, Muramatsu, YO, Shigekami, arXiv:1804.07478)*

## Left-Right symmetric model

$$Y_{ij} \overline{\hat{Q}_L^i} \Phi \hat{Q}_R^j \quad \text{where } \Phi = (\tilde{H}_u, H_d) \quad \hat{Q}_R^j = (\hat{u}_R^j, \hat{d}_R^j)^T \text{ defined.}$$



*For the realistic Yukawa,*

$$Y_{ij}^1 \overline{\hat{Q}_L^i} \Phi \hat{Q}_R^j + Y_{ij}^2 \overline{\hat{Q}_L^i} \tilde{\Phi} \hat{Q}_R^j \quad \text{in non-SUSY;}$$

$$Y_{ij}^1 \overline{\hat{Q}_L^i} \Phi_1 \hat{Q}_R^j + Y_{ij}^2 \overline{\hat{Q}_L^i} \Phi_2 \hat{Q}_R^j \quad \text{in SUSY.}$$

After the LR symmetry breaking,

2HDM (4HDM) with large FCNCs appears.

in non-SUSY,

(Iguro, Muramatsu, YO, Shigekami, arXiv:1804.07478)

$$\rho_{ij}^u \overline{\hat{Q}_L^i} \tilde{H} \hat{u}_R^j + \rho_{ij}^d \overline{\hat{Q}_L^i} H \hat{d}_R^j$$

in SUSY,

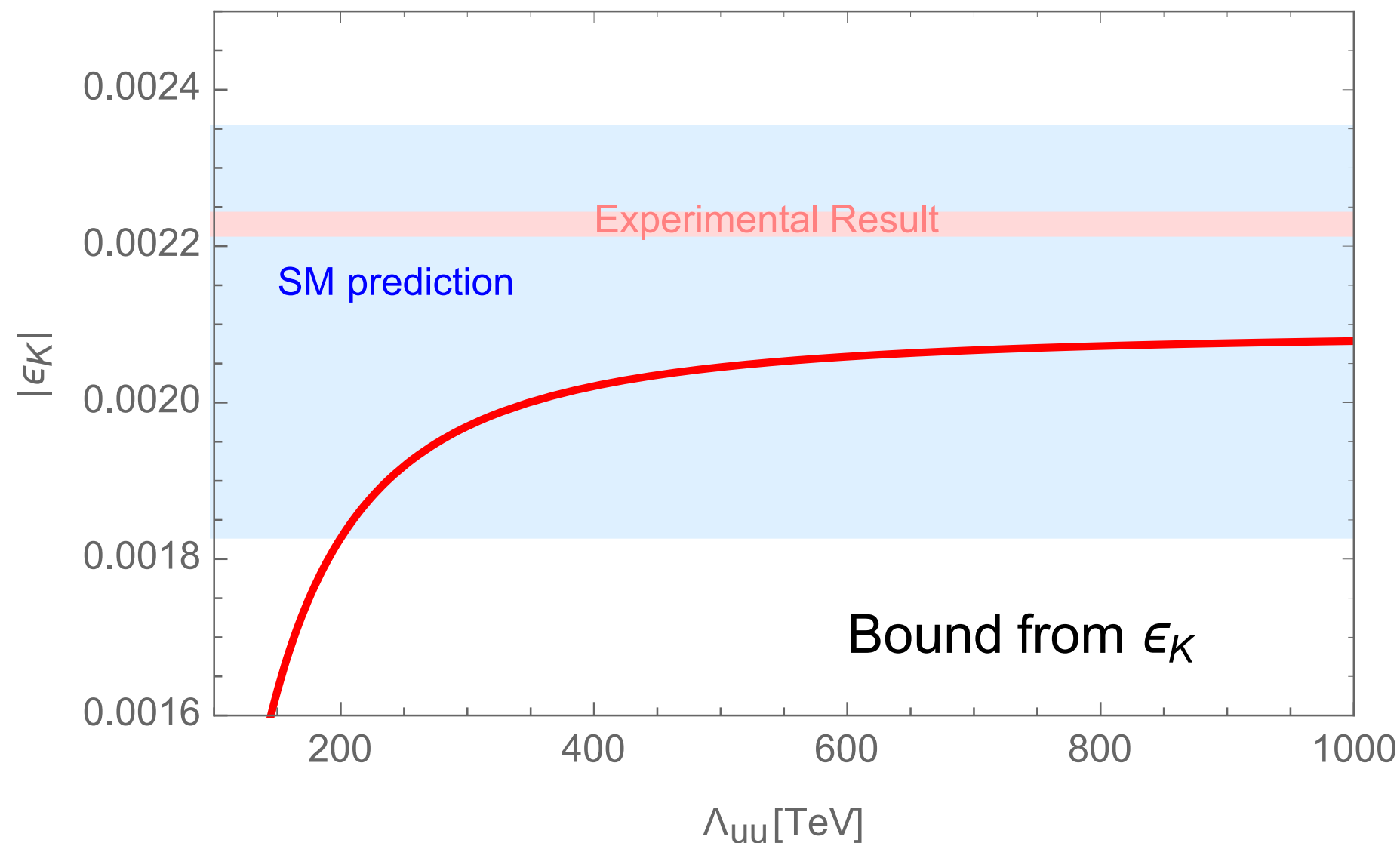
$$\sum_{A=1}^3 \left( \rho_{Aij}^u \overline{\hat{Q}_L^i} \tilde{H}_A \hat{u}_R^j + \rho_{Aij}^d \overline{\hat{Q}_L^i} H_A \hat{d}_R^j \right)$$

Relation between the FCNC and the measured values

$$\begin{pmatrix} \rho_{Aij}^u \\ \rho_{Aij}^d \end{pmatrix} = \begin{pmatrix} U_{11}^A & U_{12}^A \\ U_{21}^A & U_{22}^A \end{pmatrix} \begin{pmatrix} V_{ik}^\dagger \frac{\sqrt{2}m_k^u}{v} V_{kj} \\ \frac{\sqrt{2}m_i^d}{v} \delta_{ij} \end{pmatrix}$$

Very “predictable,”  
so that the flavor constraint is very severe.

(Iguro, Muramatsu, YO, Shigekami, arXiv:1804.07478)



Extra Higgs scales are naively  $O(100)$  TeV.

# Summary and Discussion

## Why do we need the scalar?

Additional symmetry often requires extra Higgs to realize the realistic Yukawa.

SUSY, GUT,  
flavor symmetry,  $U(1)_{PQ}$ , etc.

## Where is the scalar?

In “MSSM-like” 2HDM,

$b \rightarrow s \gamma$  constrains strongly:  $m_{H^\pm} > 580 \text{ GeV}$ .

125 GeV Higgs couplings need to be SM-like.

In the bottom-up approach, the scalars may be  $O(100) \text{ TeV}$ .

$H \rightarrow hh$  search, for instance, constrains a lot.

→ Integrated research would be important.

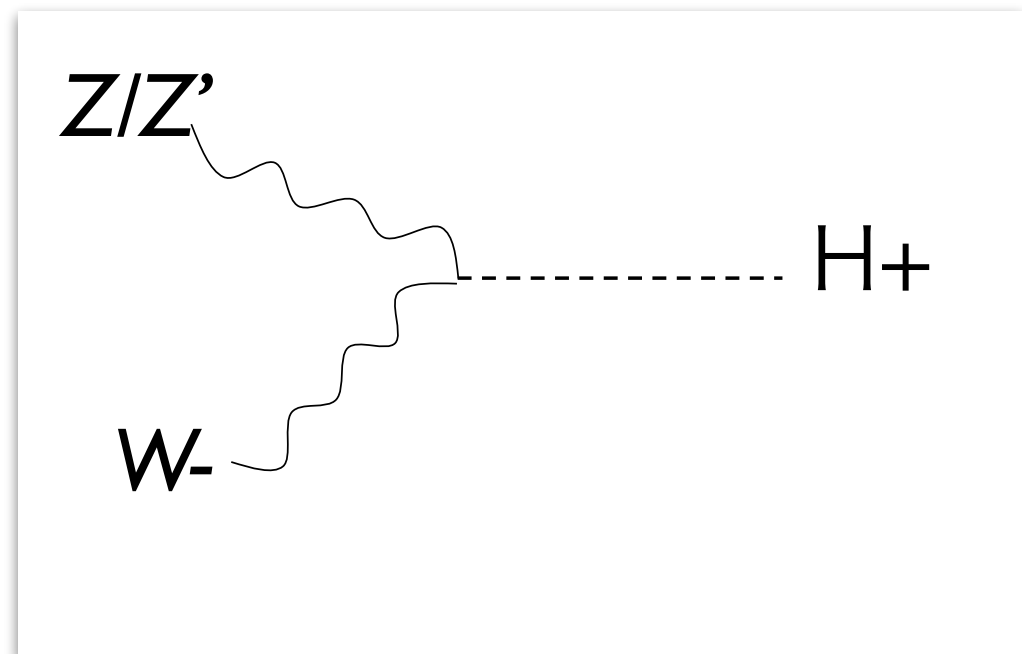
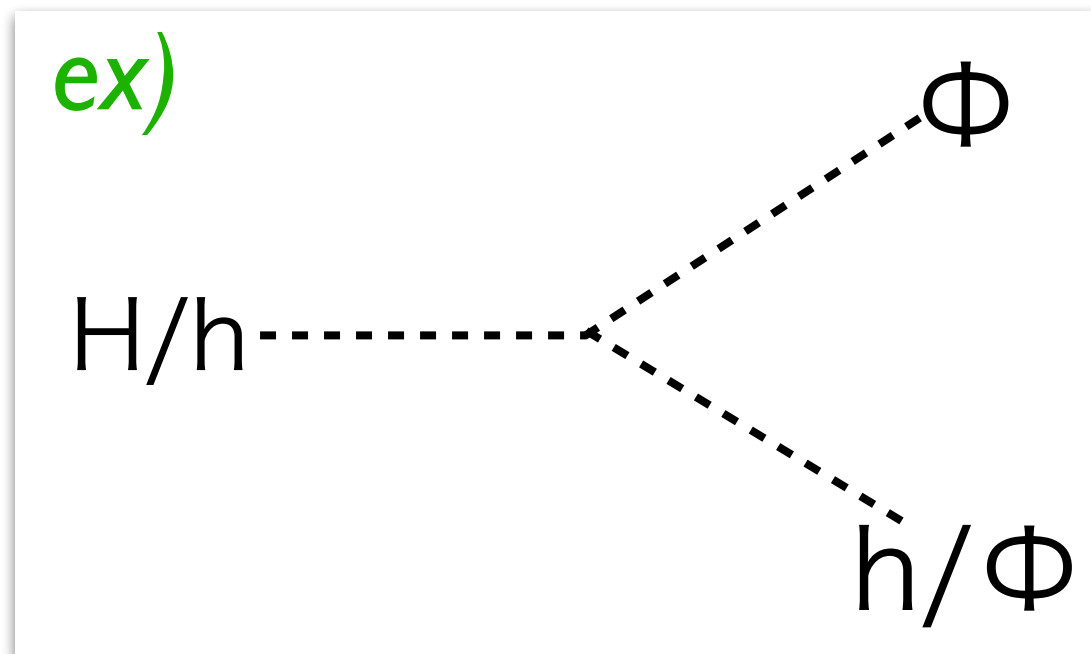
In other-type 2HDMs,

scalars can be light depending on the Yukawa couplings.



# What is interesting and what can we do?

- We can discuss the scenario where the predictions are deviated from the SM, in the other-type 2HDMs.
- Is there something new behind the excesses? How can we test?  
Large  $(t, c)$ ,  $(\tau, \mu)$  couplings in 2HDM.
- 2HDM may be too minimal  
→ 2HDM+(scalar, Higgsino, or  $Z'$ , etc.) is more realistic?



Backup

# In the Type-I 2HDM

